

TeV muon bundles in air showers detected with IceTop & IceCube

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ICECUBE
NEUTRINO OBSERVATORY



**GHENT
UNIVERSITY**

Introduction

➤ Indirect CR measurements

- Primary nucleus: E_0, A, θ

Air shower



Analysis



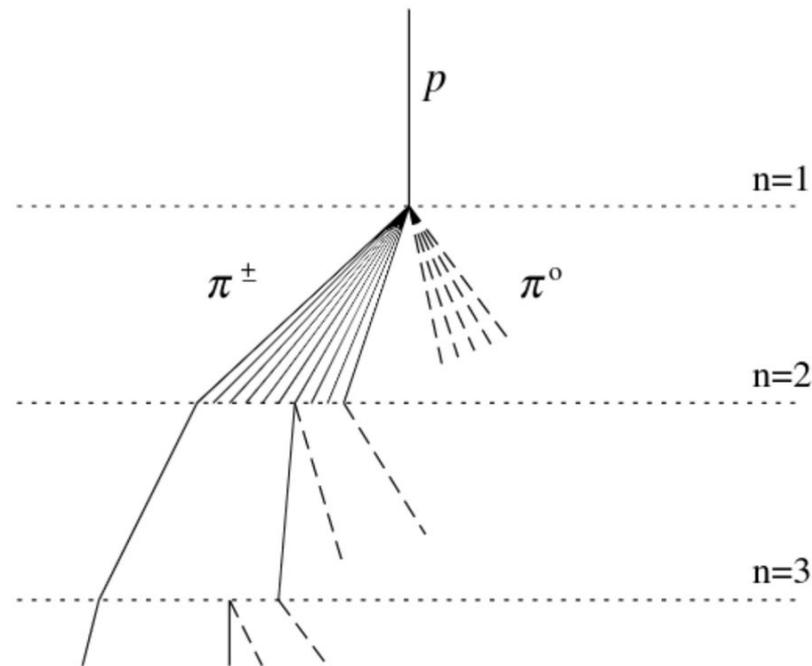
- Observables: $N_\mu, N_e, X_{\max}, \dots$

➤ Muons in air showers

- Mass sensitive
- Tracers of the hadronic cascade
- Heiter-Matthews model:

$$\cdot N_\mu = A \left(\frac{E_0}{AC} \right)^\beta$$

$$\cdot \ln N_\mu = (1 - \beta) \ln A + \beta \ln (E_0/C) \quad \beta \approx 0.9$$



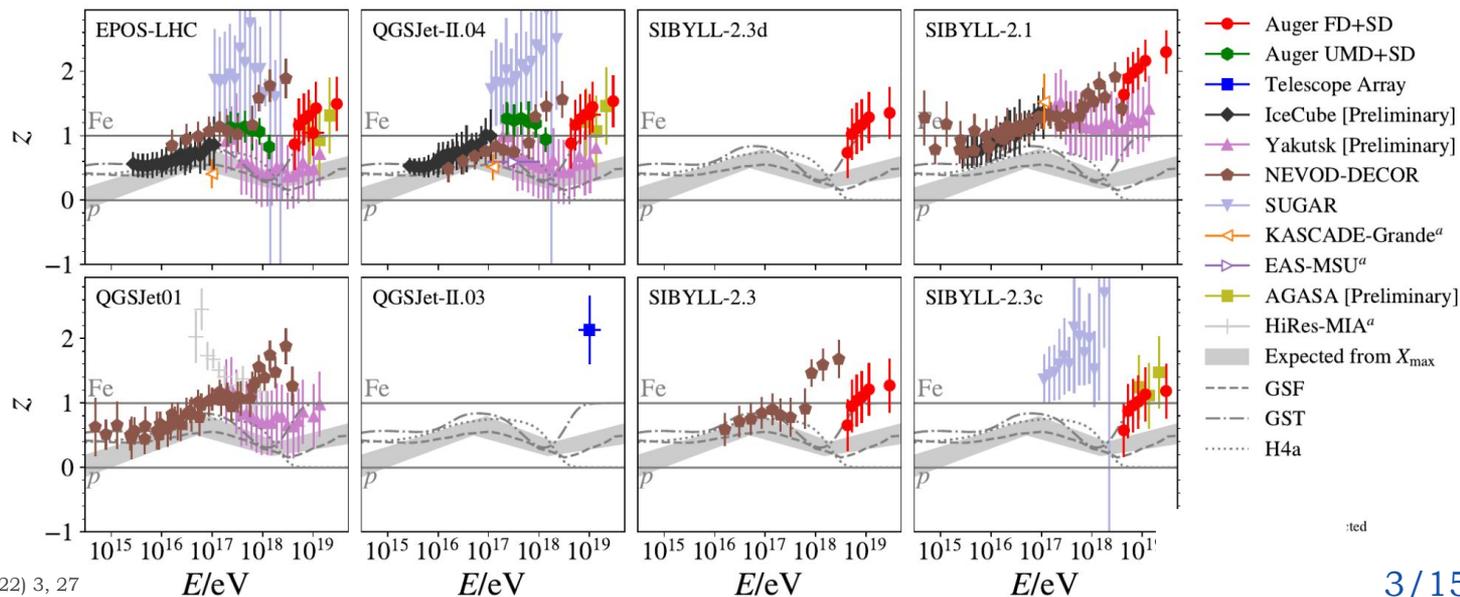
The Muon Puzzle

➤ Air shower simulations

- Necessary for interpretation of measurements
- Hadronic interaction models
- Uncertainties due to extrapolations outside of accelerator phase space

→ Discrepancies between data and MC established for state-of-the-art models

$$z = \frac{\ln \langle N_\mu \rangle_{\text{data}} - \ln \langle N_\mu \rangle_{\text{MC}}^p}{\ln \langle N_\mu \rangle_{\text{MC}}^{\text{Fe}} - \ln \langle N_\mu \rangle_{\text{MC}}^p}$$



IceCube Neutrino Observatory

➤ IceCube

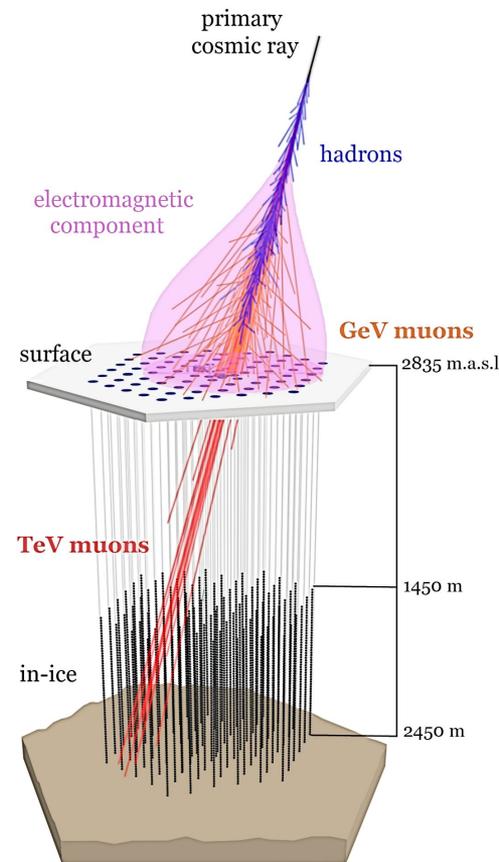
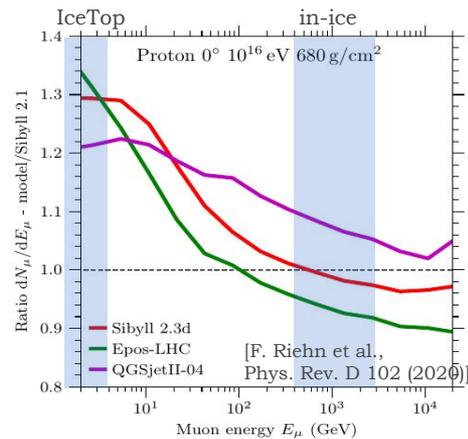
- $\sim 1 \text{ km}^3$ instrumented volume
- 86 strings with ~ 5000 Digital Optical Modules (DOMs)

➤ IceTop

- $\sim 1 \text{ km}^2$ air shower array
- Atmospheric depth $\sim 690 \text{ g/cm}^2$
- 81×2 Ice Cherenkov Tanks with 2 DOMs
- Primary energies $\sim \text{PeV} - \text{EeV}$

➤ Combined: Unique EAS Detector

- Electromagnetic component
- GeV muon content
- **TeV muon content**



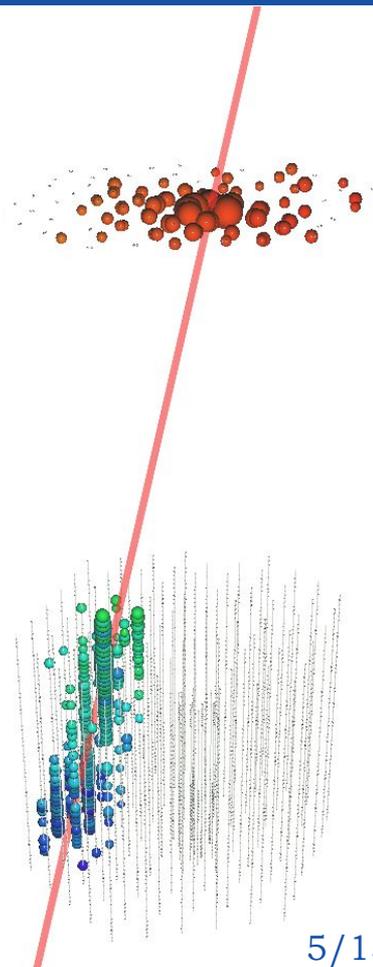
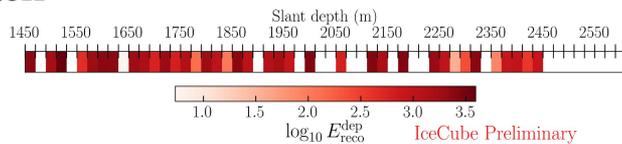
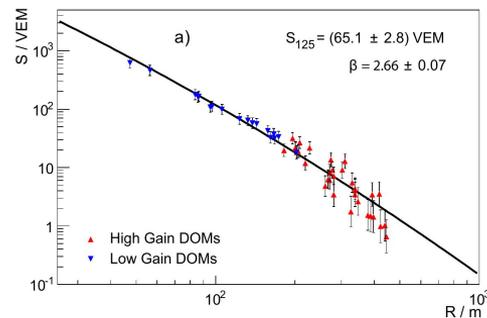
EAS Reconstruction

➤ IceTop

- Fit to IceTop signals
 - Lateral distribution function (charge)
 - Shower front (time)
- Direction & core position
- Shower size S_{125} : proxy for primary energy

➤ In-Ice

- Energy loss reconstruction
 - Along reconstructed IceTop track
 - In segments of 20 m
- Vector of deposited energy along track



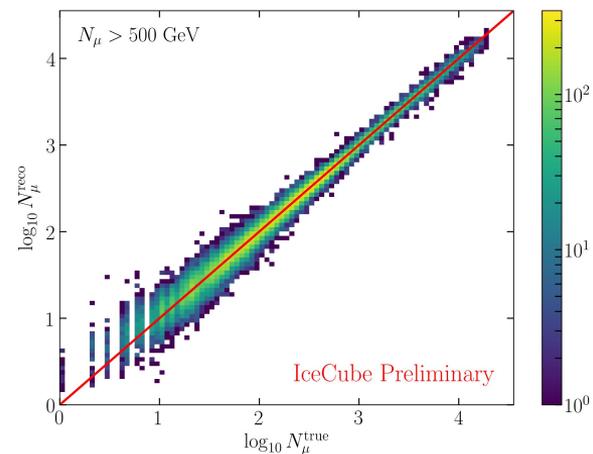
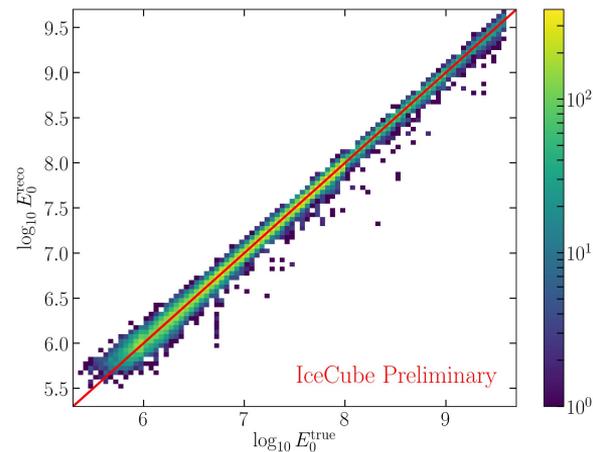
Neural Network

➤ Neural network reconstruction

- Inputs
 - Shower size S_{125}
 - Zenith θ
 - Energy loss vector
- Outputs
 - **Primary energy E_0**
 - **Number of muons > 500 GeV**
in shower at surface N_μ
- RNN + Dense layers

➤ MC Dataset

- Sibyll 2.1
- p, He, O, Fe
- Coincident events, contained in IceTop
- $\cos \theta > 0.95$ ($\theta \lesssim 18^\circ$)



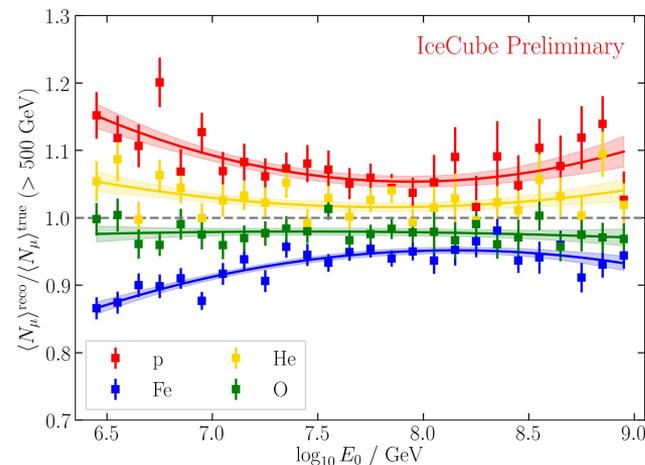
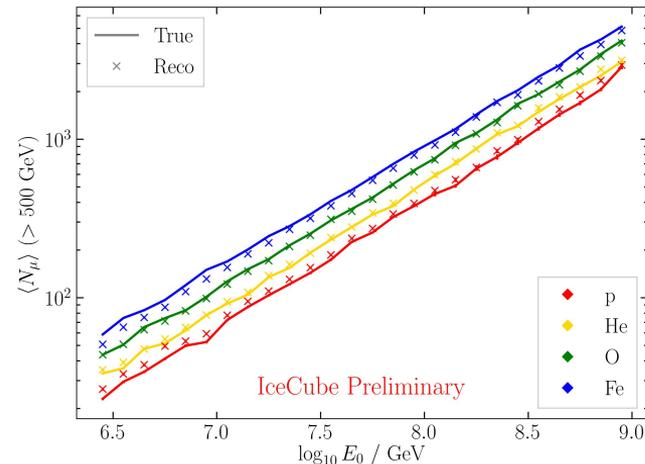
Correction factor

➤ Determination of $\langle N_{\mu} > 500 \text{ GeV} \rangle$

- Bins of $\log_{10} E_0$
- Low-energy limit: IceTop threshold
- Comparison between
 - MC true values
 - neural-network reconstructions

➤ Correction factor

- Composition dependent over/underestimation
- Ratios fitted with quadratic function
- Used to correct bias



Iterative Correction

➤ Reconstruction bias

- Bias / correction composition dependent
 - $\langle N_\mu \rangle$ has composition information
- Iterative procedure

➤ Iterative correction procedure

- Linear combination of p & Fe corrections

$$C_{\text{eff}} = f_p C_p + f_{\text{Fe}} C_{\text{Fe}}$$

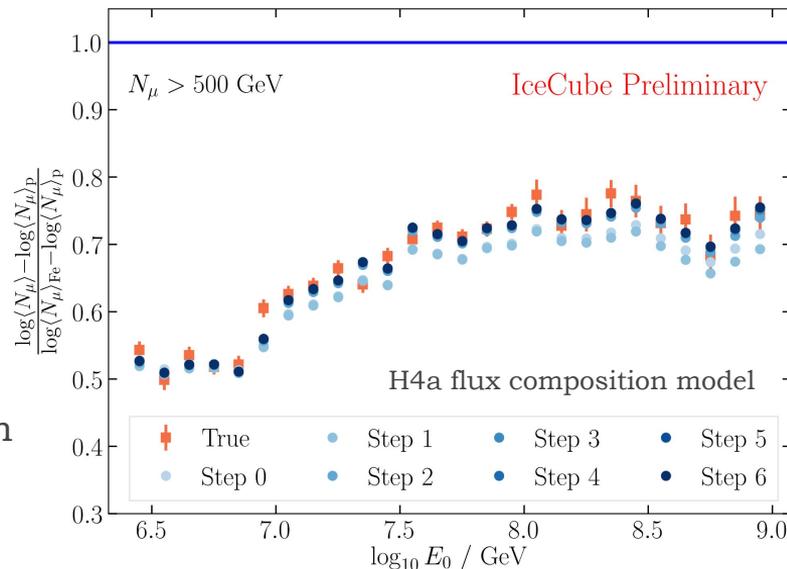
- Fractions f_p and f_{Fe} describe average composition

$$f_p \ln A_p + f_{\text{Fe}} \ln A_{\text{Fe}} = \langle \ln A \rangle$$

- Composition estimate:

$$z = \frac{\ln \langle N_\mu \rangle - \ln \langle N_\mu \rangle_p}{\ln \langle N_\mu \rangle_{\text{Fe}} - \ln \langle N_\mu \rangle_p} \approx \frac{\langle \ln A \rangle}{\ln 56}$$

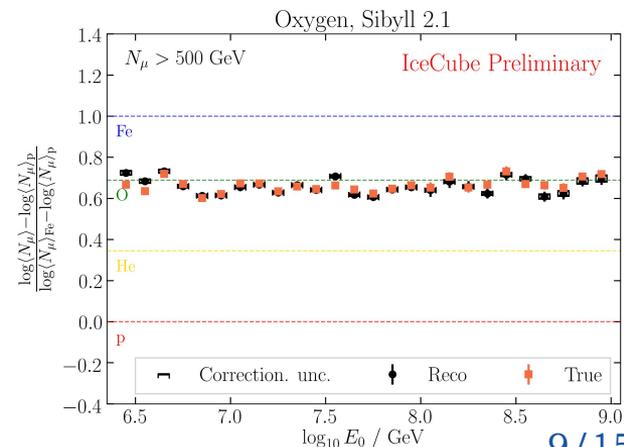
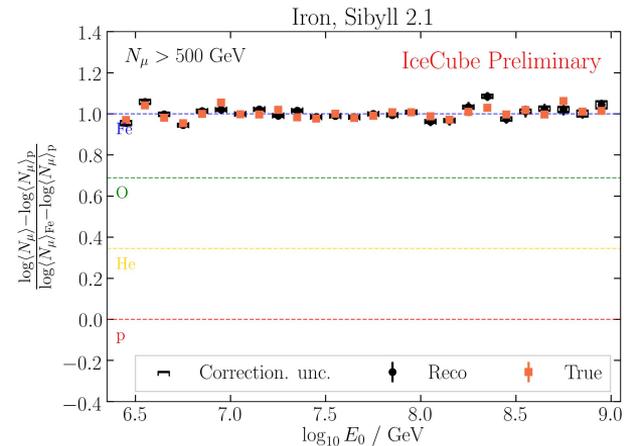
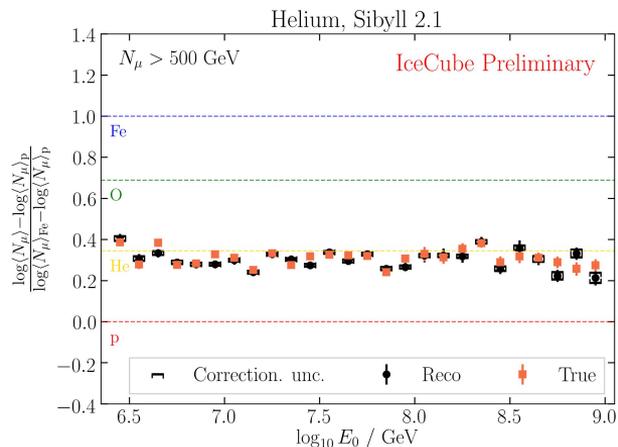
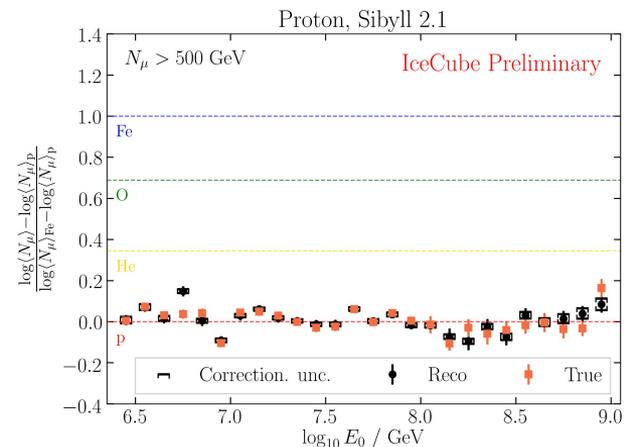
- Update $\langle N_\mu \rangle \rightarrow$ update $C_{\text{eff}} \rightarrow$ etc. until convergence



MC Tests

➤ Application of Neural Network & Correction to MC

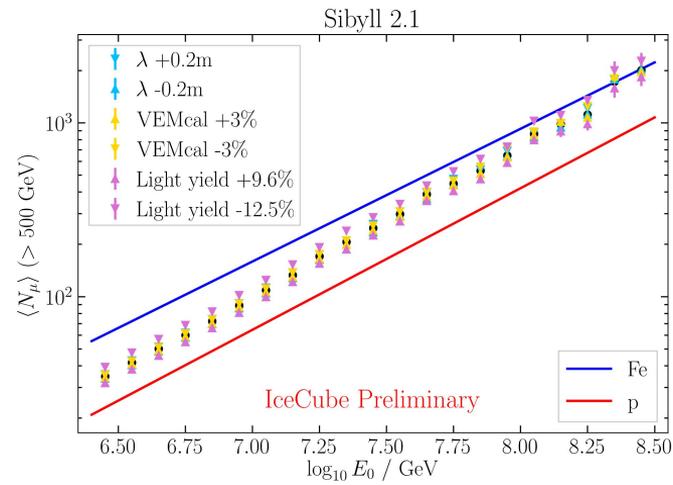
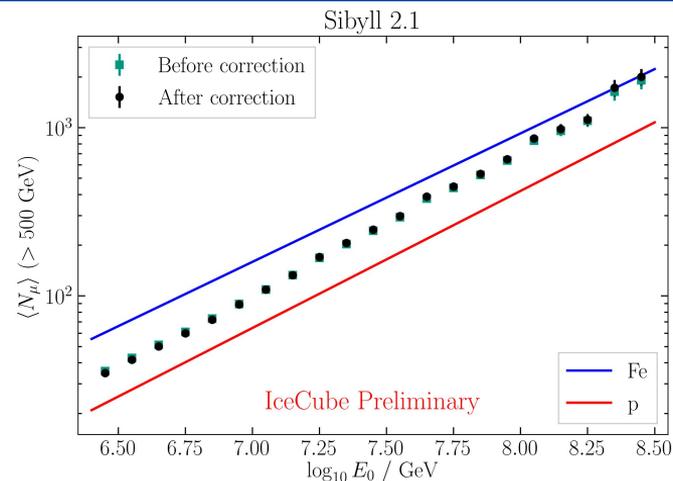
- Pure p, He, O, Fe
- Random combinations (see backup)
- Good agreement between true and reconstructed!



Results

- Application to experimental data
 - 10% of 1 year (05/2012 - 05/2013)
 - Compared to expectations from Sibyll 2.1

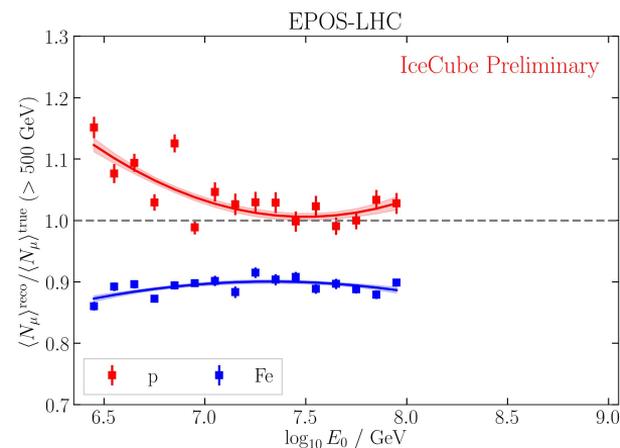
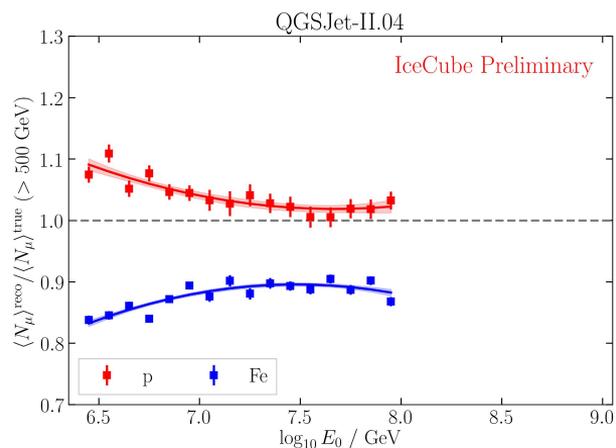
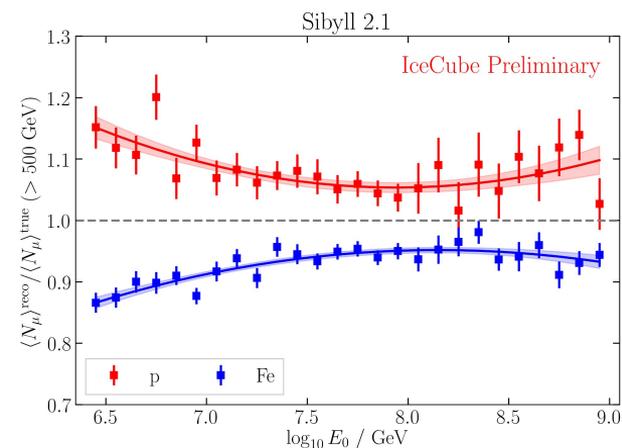
- Systematic uncertainties
 - Correction uncertainty
 - Detector uncertainties
 - Snow accumulation on IceTop
 - IceTop VEM definition / Energy scale
 - IceCube light yield (ice model, DOM eff.)



Other Hadronic Models

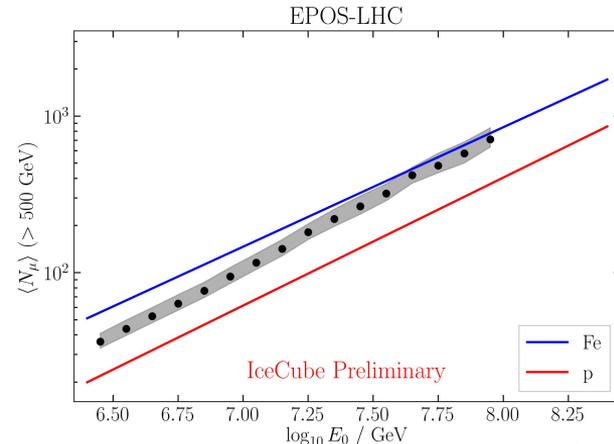
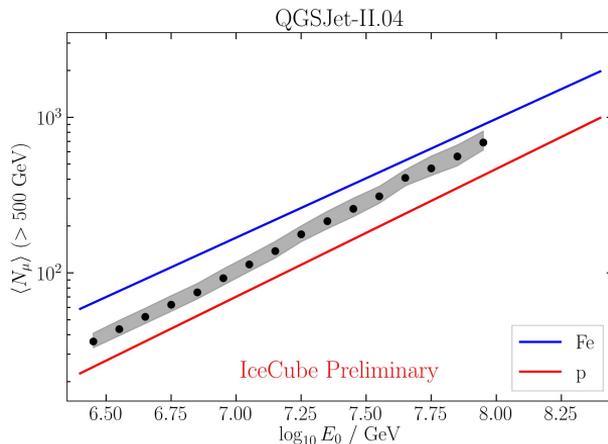
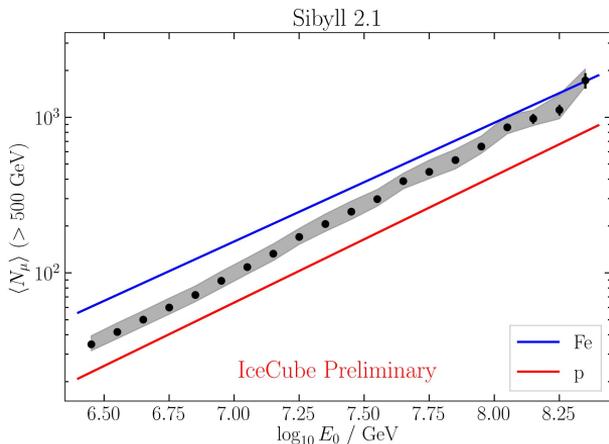
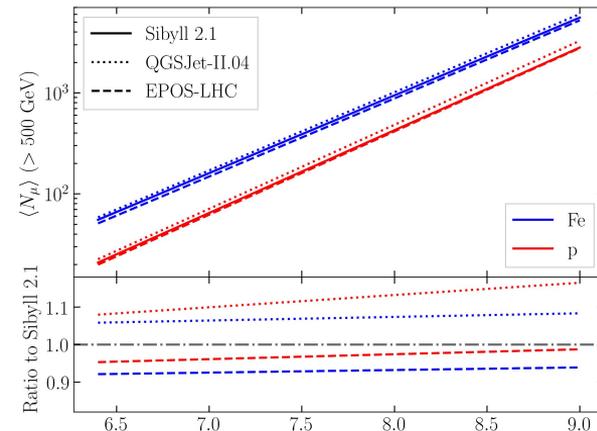
➤ Correction factors

- From MC → model dependent results
- Include other hadronic interaction models
 - QGSJet-II.04
 - EPOS-LHC
 - Limited to 100 PeV



Results

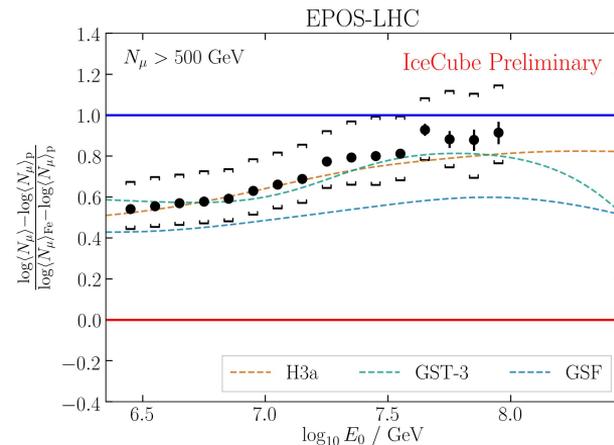
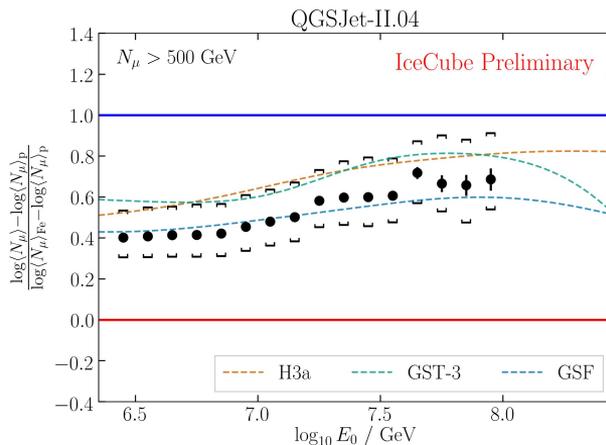
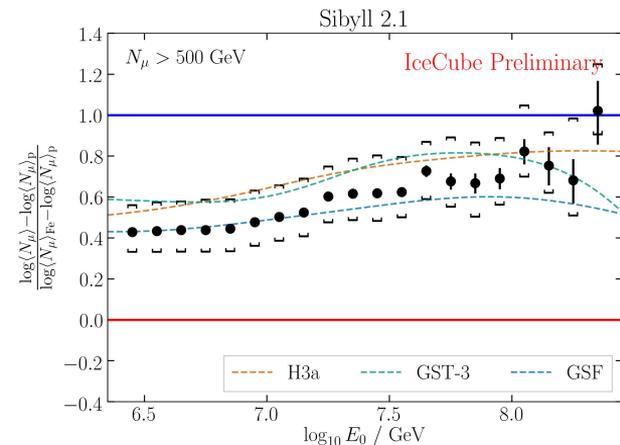
- Average muon multiplicity > 500 GeV
 - Hadronic model dependent
 - Compared to corresponding MC predictions
 - Shaded area: total systematic uncertainty



Results

➤ Results in “z-values”

- $$z = \frac{\ln \langle N_\mu \rangle - \ln \langle N_\mu \rangle_p}{\ln \langle N_\mu \rangle_{\text{Fe}} - \ln \langle N_\mu \rangle_p}$$
- Comparison to composition models H4a, GST-3, GSF
- Brackets: total systematic uncertainty



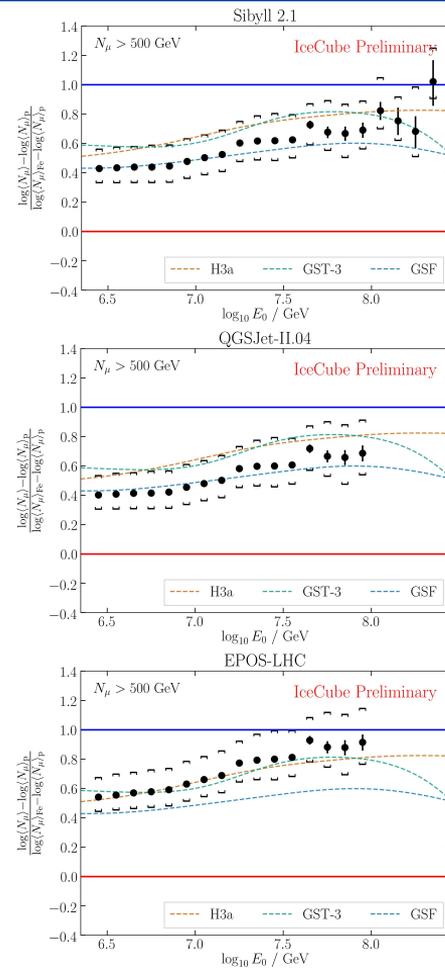
Summary & Conclusions

➤ Measurement of TeV muon content in EAS

- IceTop-IceCube coincident events
- # muons > 500 GeV in showers at surface
- Energies between 2.5 PeV –
 - 250 PeV (Sibyll 2.1)
 - 100 PeV (QGSJet-II.04, EPOS-LHC)

➤ Conclusions

- No excess/deficit
- Sibyll 2.1 and QGSJet-II.04: good agreement with composition models
- EPOS-LHC yields slightly heavier mass composition



Outlook

➤ TeV muon analysis

- Update with more data coming soon
- Several possible improvements (zenith range, in-ice systematics, seasonal variations...)

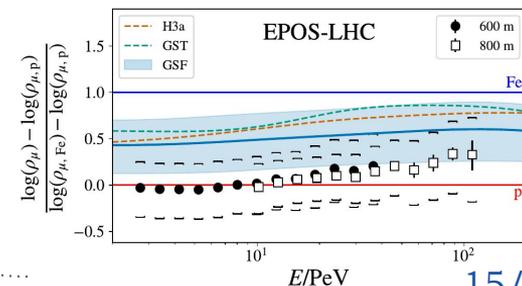
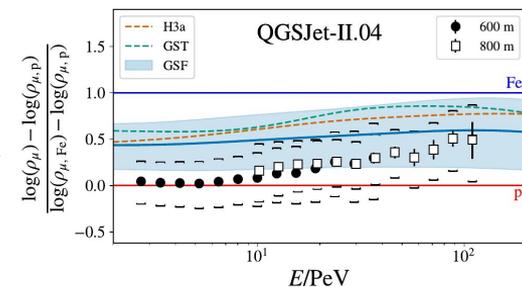
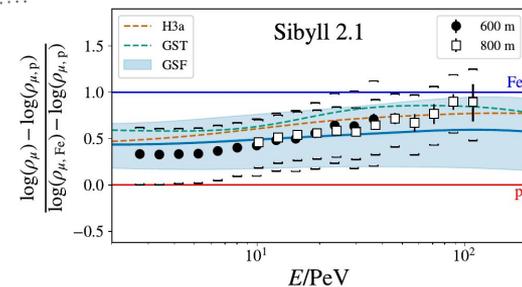
➤ Coincident measurements of GeV and TeV muons

- Unique tests of hadronic interaction models
- Density of GeV muons in IceTop [\[arXiv:2201.12635\]](#)
 - Agreement with TeV muons for Sibyll 2.1
 - Tension for QGSJet-II.04 and EPOS-LHC

→ Implies models do not correctly describe interactions

➤ IceCube Gen2 & Surface Enhancement

- Solid angle, EM/muon separation, energy scale, X_{\max} ...
[\[PoS\(ICRC2021\)407\]](#)



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THE ICECUBE COLLABORATION

FUNDING AGENCIES

Fonds de la Recherche Scientifique (FRS-FNRS)
Fonds Wetenschappelijk Onderzoek-Vlaanderen
(FWO-Vlaanderen)

Federal Ministry of Education and Research (BMBF)
German Research Foundation (DFG)
Deutsches Elektronen-Synchrotron (DESY)

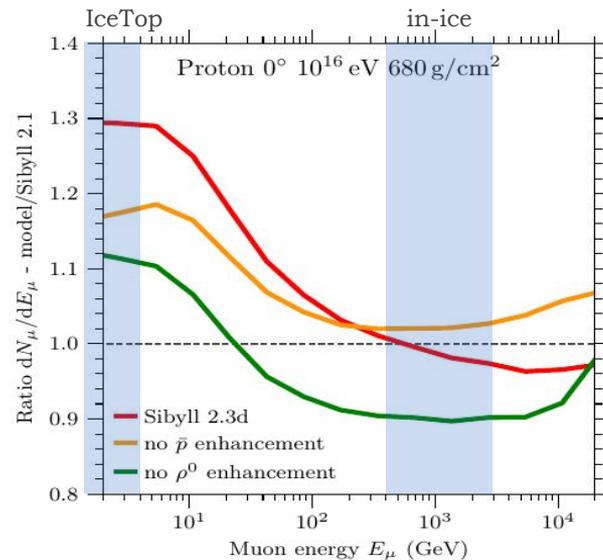
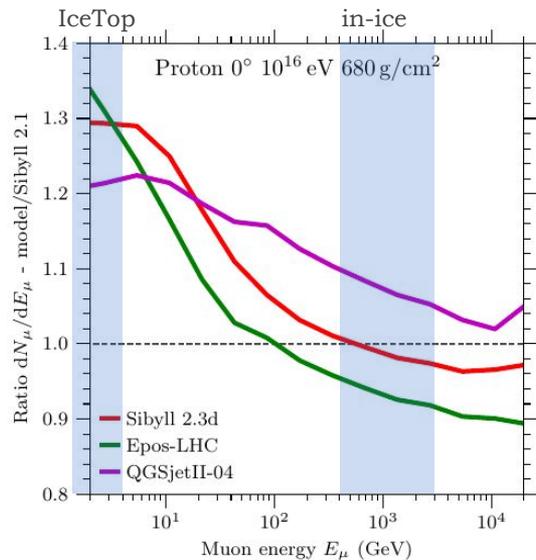
Japan Society for the Promotion of Science (JSPS)
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Swedish Polar Research Secretariat

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University of Wisconsin Alumni Research Foundation (WARF)
US National Science Foundation (NSF)

Backup

Hadronic interaction models

- Measurements of GeV and TeV muons can uniquely constrain hadronic interaction models

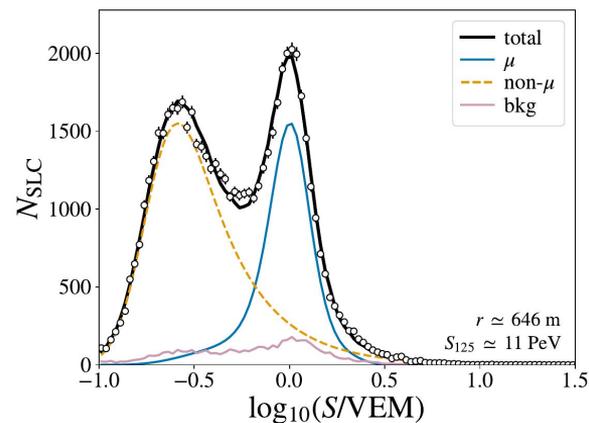
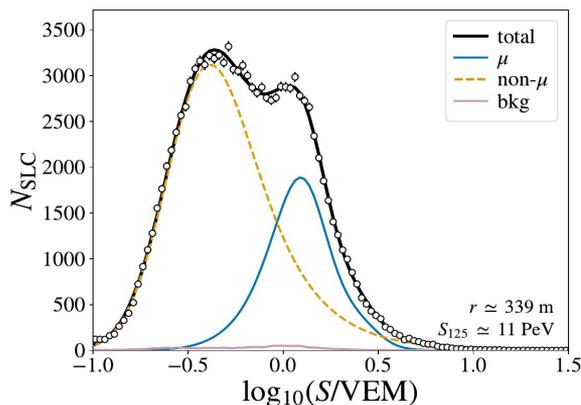
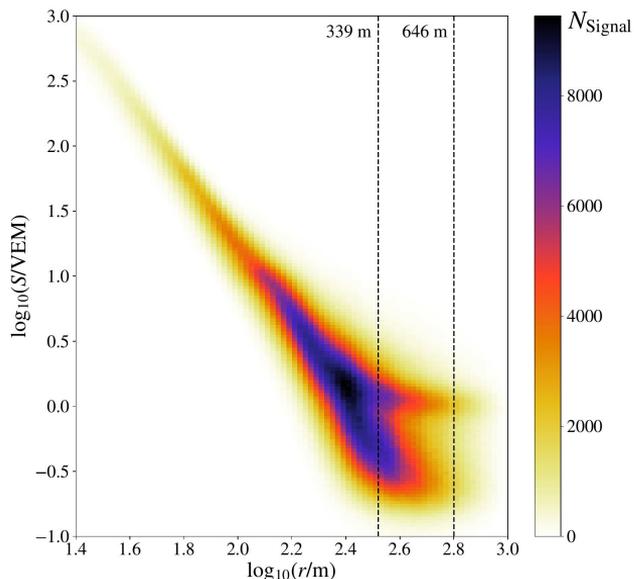


[F. Riehn et al., Phys. Rev. D 102 (2020)]

Density of GeV muons in IceTop

➤ Analysis method

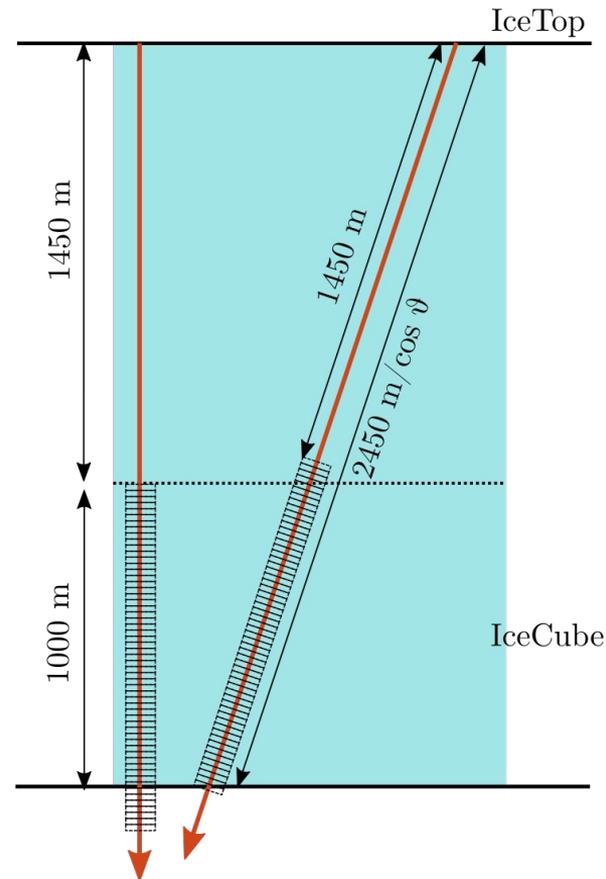
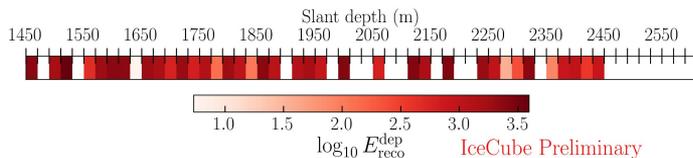
- At large lateral distance: typical 1 VEM muon signal > EM signal
- Fit signals with different components
- Fit muon LDF, obtain density @ 600 m, 800 m
- Apply MC corrections



Energy loss input

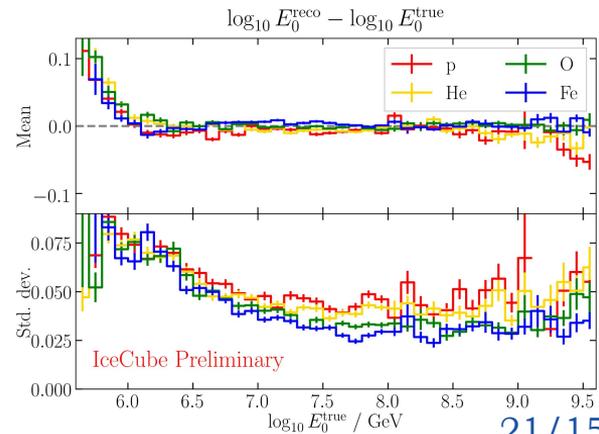
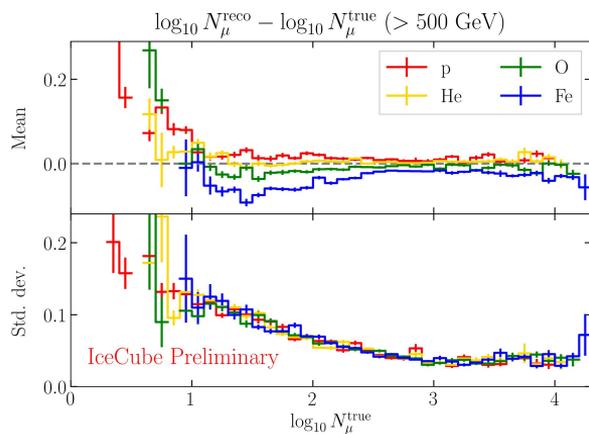
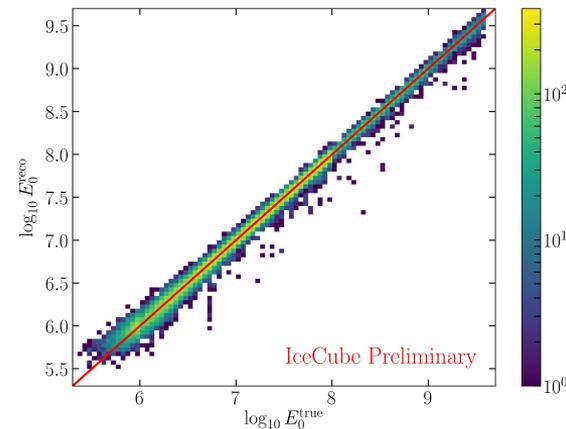
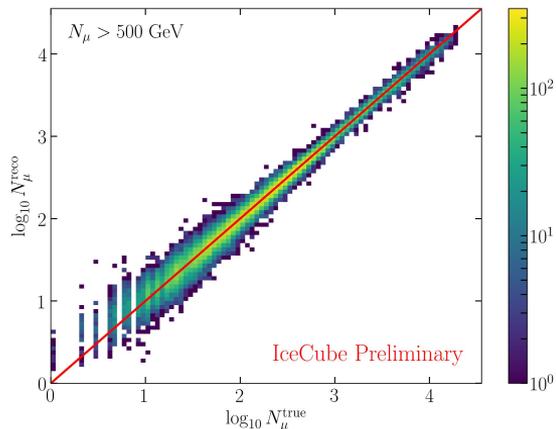
➤ Energy loss input

- Deposited energy reconstruction in segments along shower axis track
- Remove segments outside detector
- Pad to vector of fixed length 57 (based on zenith angle, limited to $\cos \theta > 0.95$)
- Vertical event example



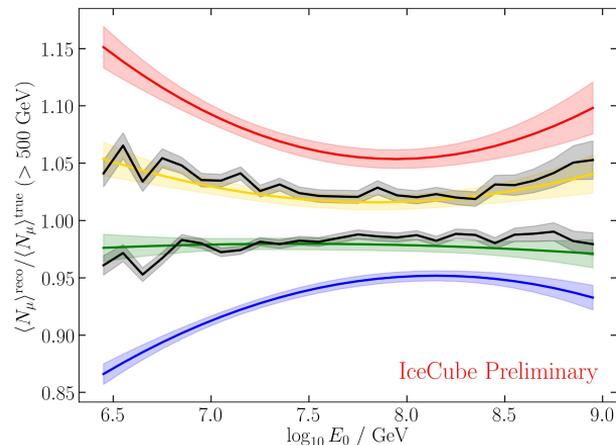
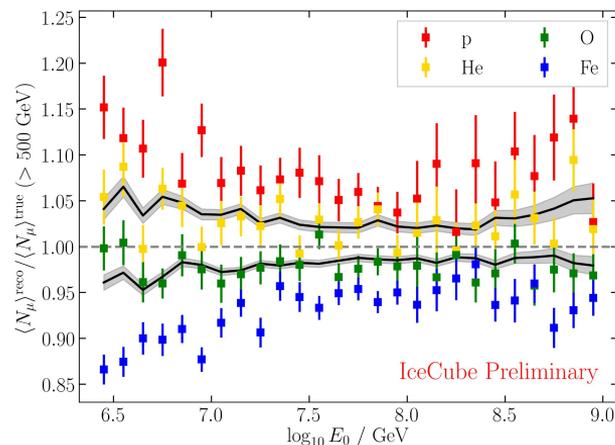
Neural Network Performance

- Performance on test set
 - Correlation plots (p, He, O, Fe combined)
 - Bias & resolution plots (by primary type)



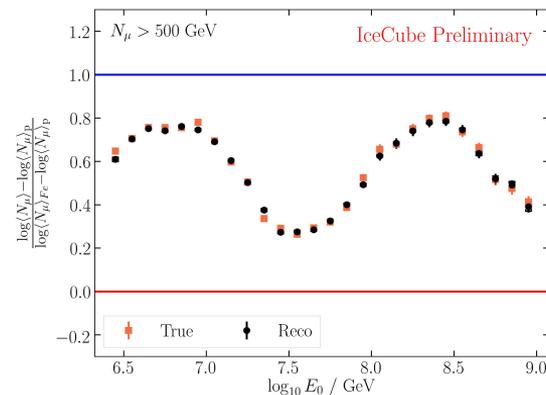
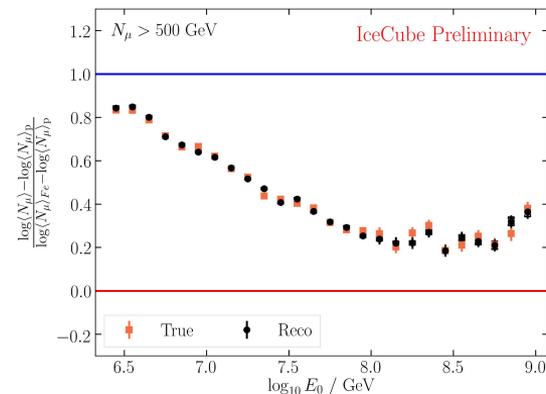
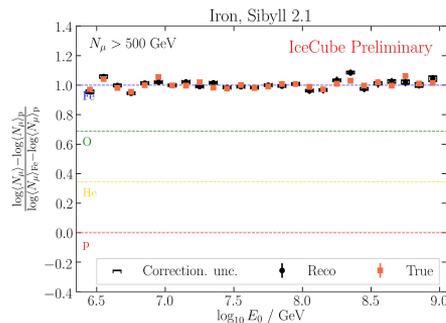
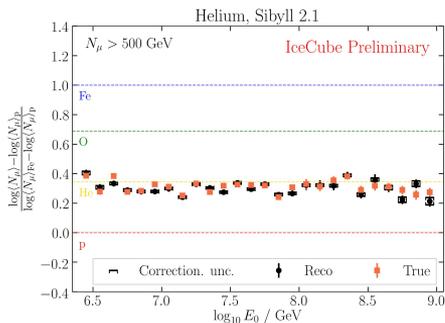
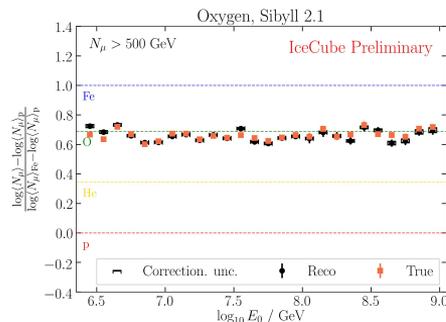
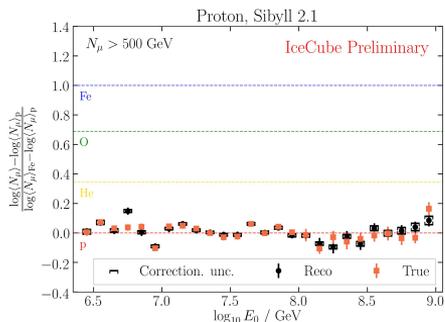
Iterative Correction

- Important check: can correction factor of intermediate elements be obtained by combining p & Fe correction factors?
 - Use pure He and O MC
 - Use true $\langle N_\mu \rangle$ in He and O
 - Based on this, calculate fractions f_p and f_{Fe}
 - Combine p & Fe correction factors with these fractions → Grey lines in plots
 - Agrees with true He and O correction factors!



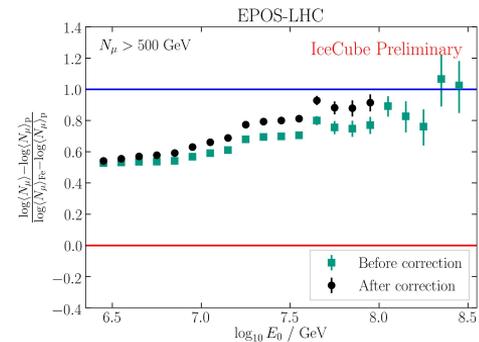
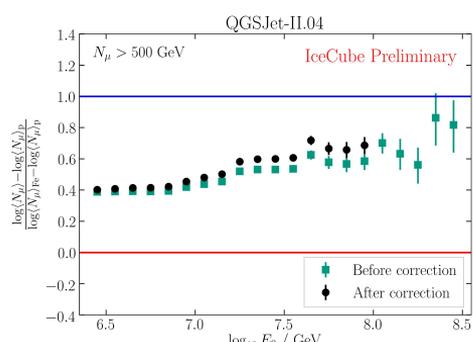
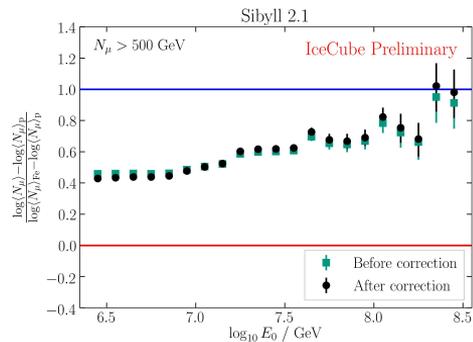
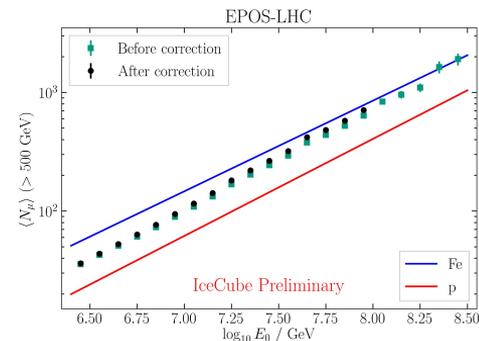
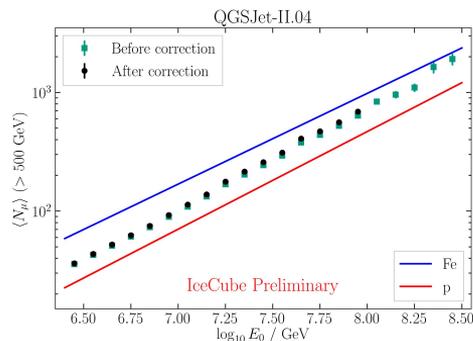
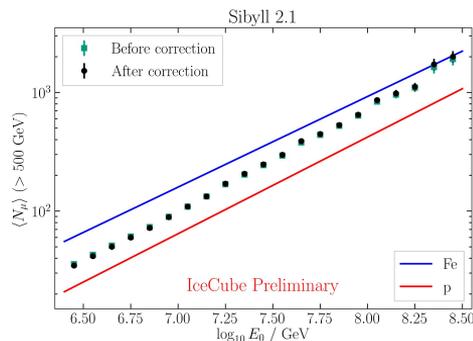
MC checks

- Application of reconstructions and correction to different composition cases
 - 1 component MC (left)
 - 4 component weighted to artificial composition (right)



Application to data

- Application of reconstructions and correction to experimental data
 - 10% of IC86.2012
 - Different model dependent results



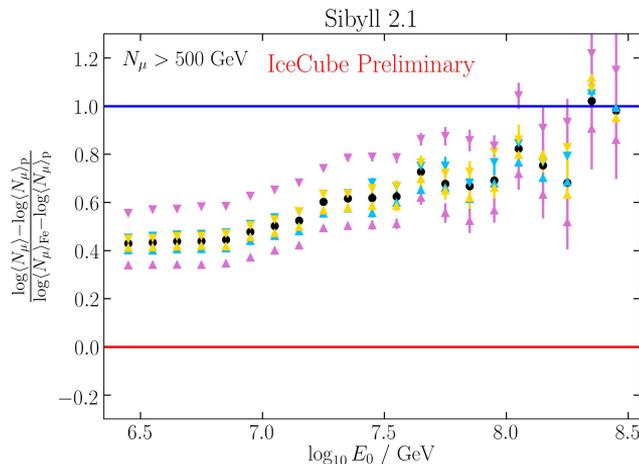
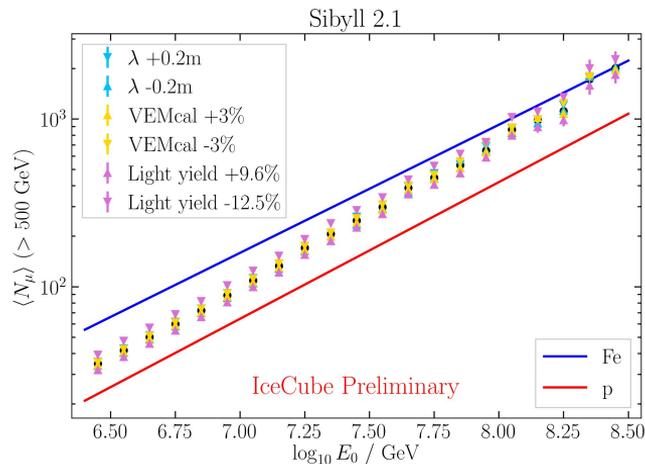
Systematic Uncertainties

➤ Correction uncertainty

- Propagated from p & Fe correction factor uncertainty

➤ Detector uncertainties

- Following 3-year composition & spectrum paper [M. G. Aartsen et al., Phys. Rev. D 100 (2019)]
 - Snow correction $\lambda \pm 0.2\text{m}$
 - VEMCal $\pm 3\%$
 - InIce combined light yield uncertainty $+9.6\%$, -12.5%



Results

➤ How do individual results compare?

- Average given with envelope describing model differences
- Less than $\pm 5\%$ variation around average

