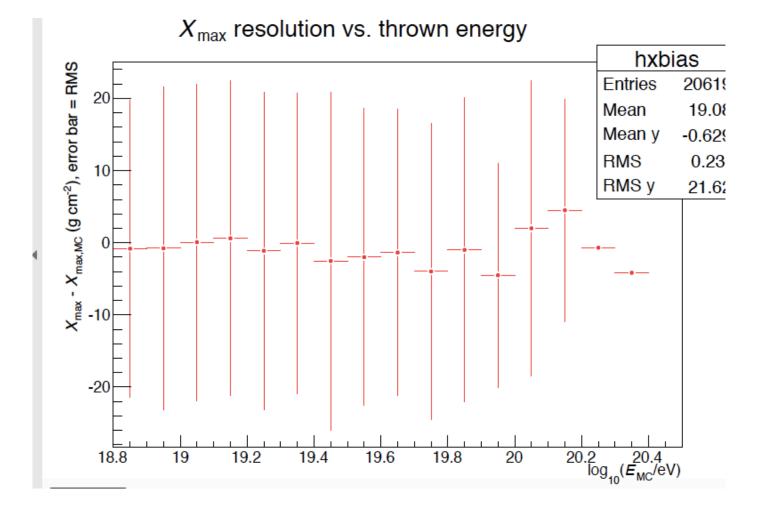
Snake Array

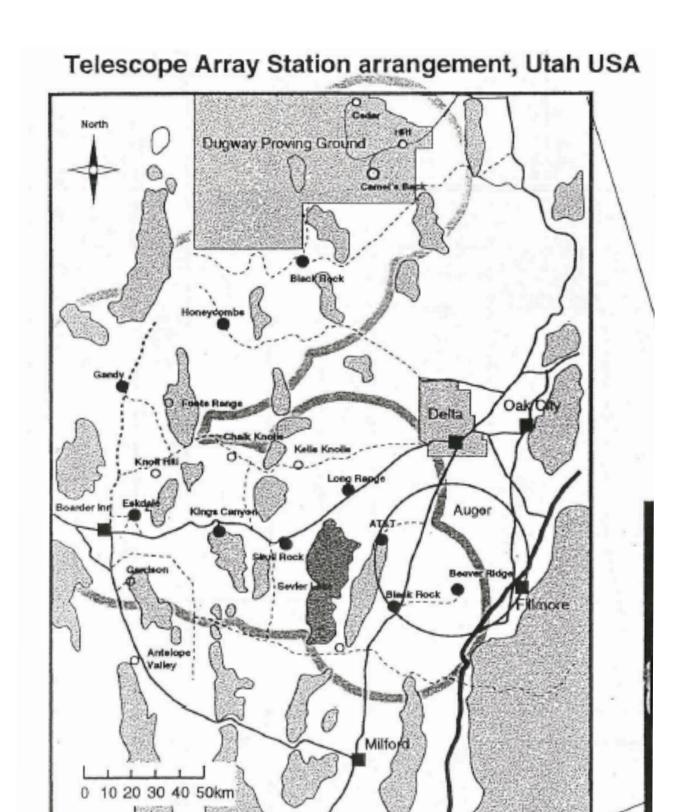
Pierre Sokolsky University of Utah GCOS 2022

Stereo Xmax resolution



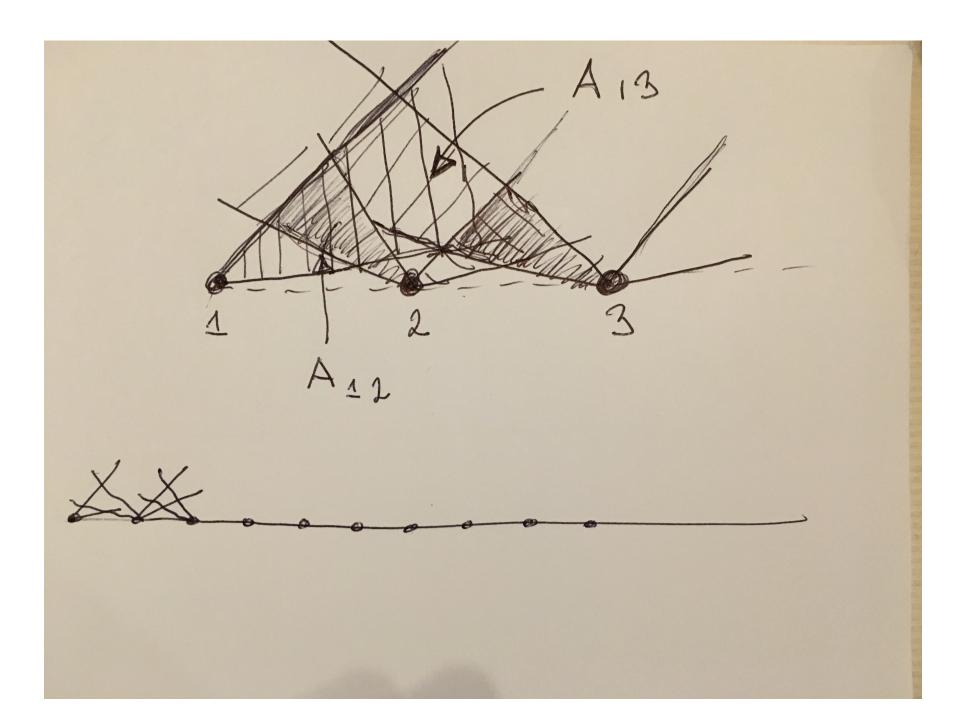
Before TA there was the Snake Array

- Discussions between Gene Loh, M. Teshima and myself (help from L. Wiencke,S. Yoshida)
- Maximize FD Stereo, no SD.
- Optimization -> minimize redundacy only 2 detector stereo -> linear array of detectors.
- Site survey was done and > 10 sites found
- ~ 30-40 km apart.

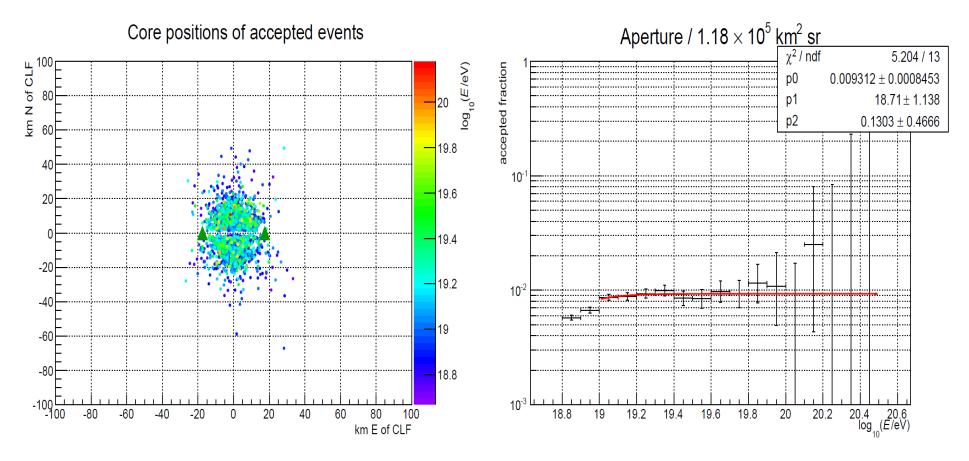


Refine "Snake" based on TA stereo

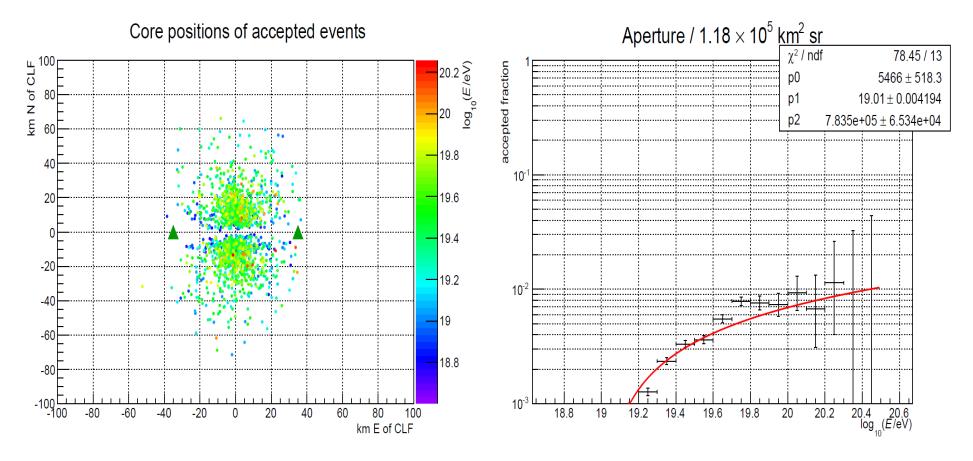
- Use current stereo reconstruction and data cuts. Use TA mirror/pmt layout
- Guaranteed Xmax and Energy resolution.
- Aperture verified cross calibrated with hybrid results.
- Linear array minimizes "non-productive" aperture.



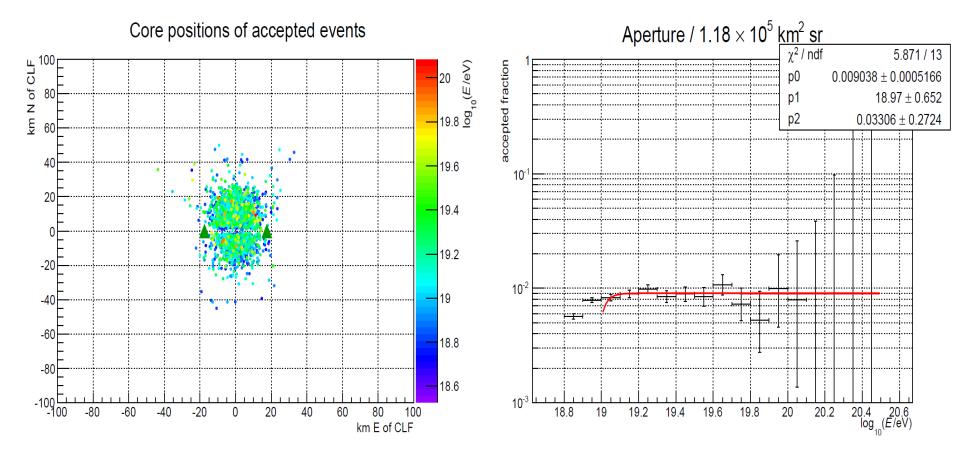
35km spacing, head-on



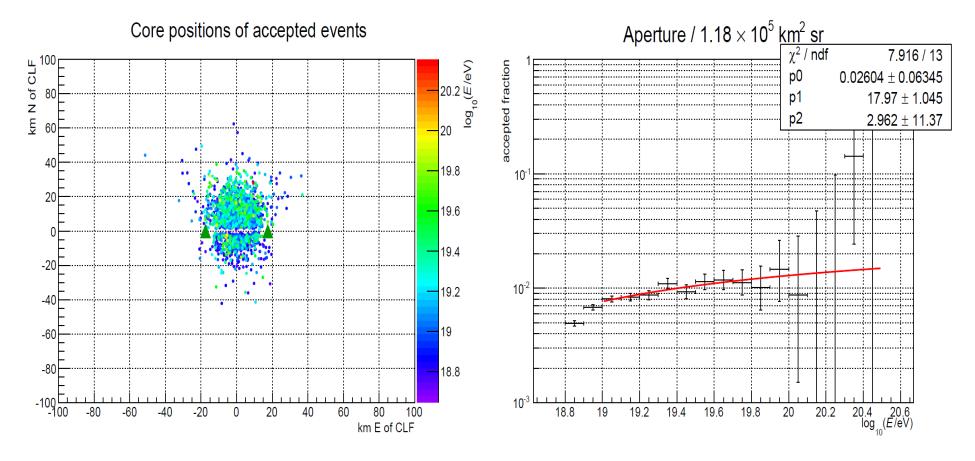
70km spacing, head-on



35km, 20° combined rotation



35km, 40° combined rotation



Estimated reconstructed and quality cut aperture

- Pair of stations 35 km spacing, 40 degree angle -1.5x10³ km2str @ 5x10¹⁹ eV
- Additional aperture for 70 km, 40 degree angle pair ~ 0.7x10³ km2str
- Total A12 + (A13) aperture 2.2x10³ km2str
- Total for 10 sites ~ 2.2x10⁴ km2str
- Assuming 15% on time effective aperture is 3.3x10³ km2str.
- Extrapolating from TA stereo data, expect ~1000 wellreconstructed events > 3x10¹⁹ eV in 10 yr run

Total Snake Cost

- Could build Snake for ~ 45-50M\$
- Would provide "well-reconstructed, good resolution Xmax data" = Auger SD aperture.
- Based on TA stereo data, expect ~ 1000 well reconstructed events above 3 x 10^19 eV in ~ 10 years.