

Caustics* in UHE Cosmic Rays

G. Farrar, NYU

***Rapidly varying
magnification and demagnification
as a function of direction and rigidity**

Deflections of UHECRs in the Galactic magnetic field

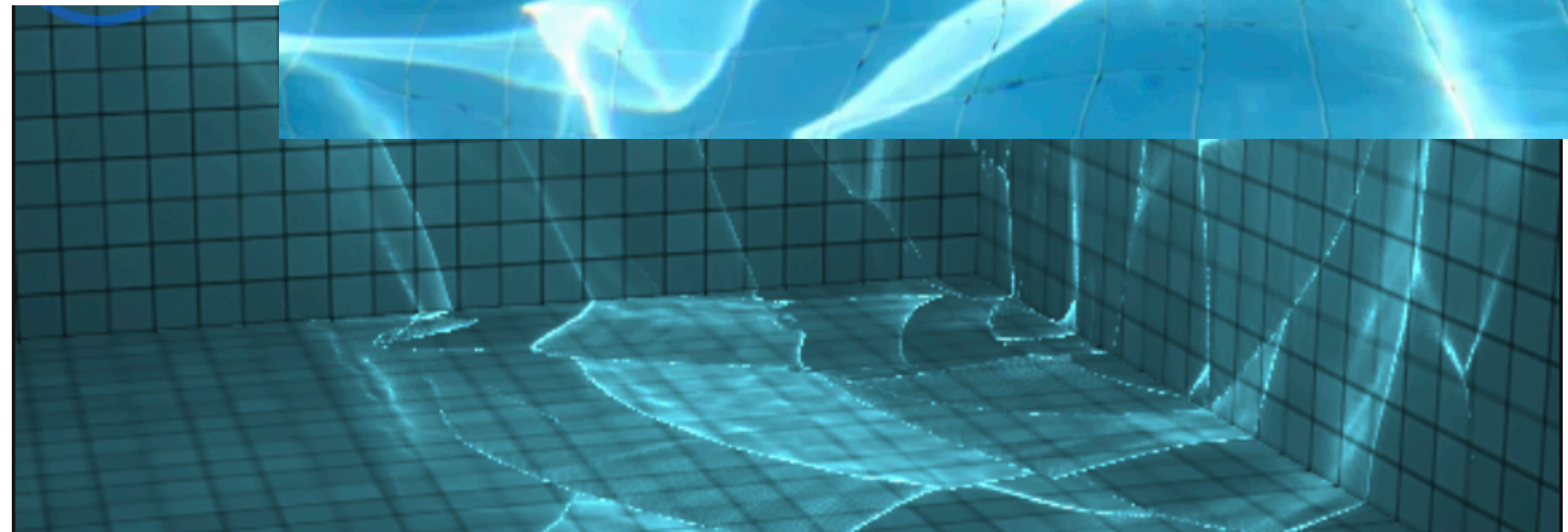
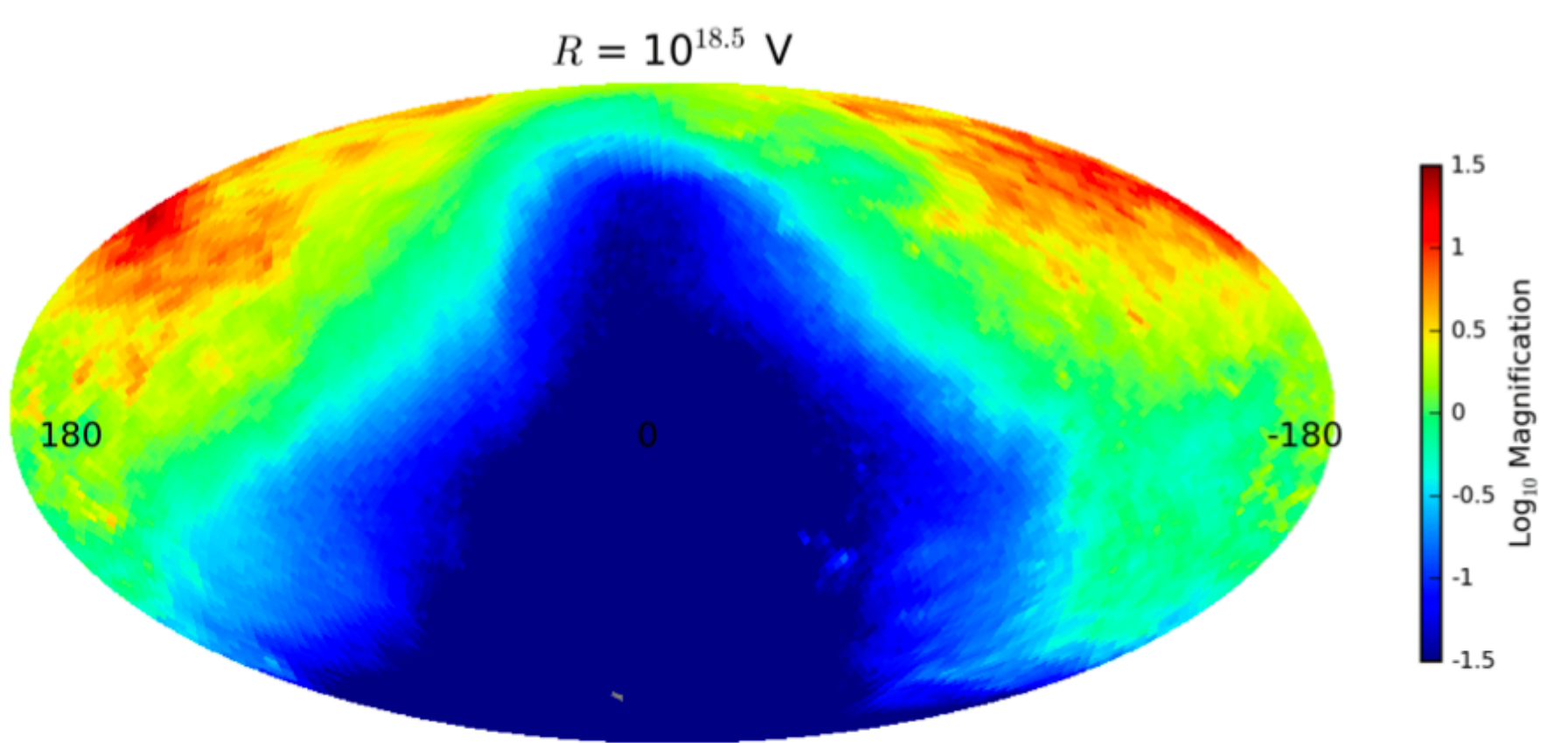
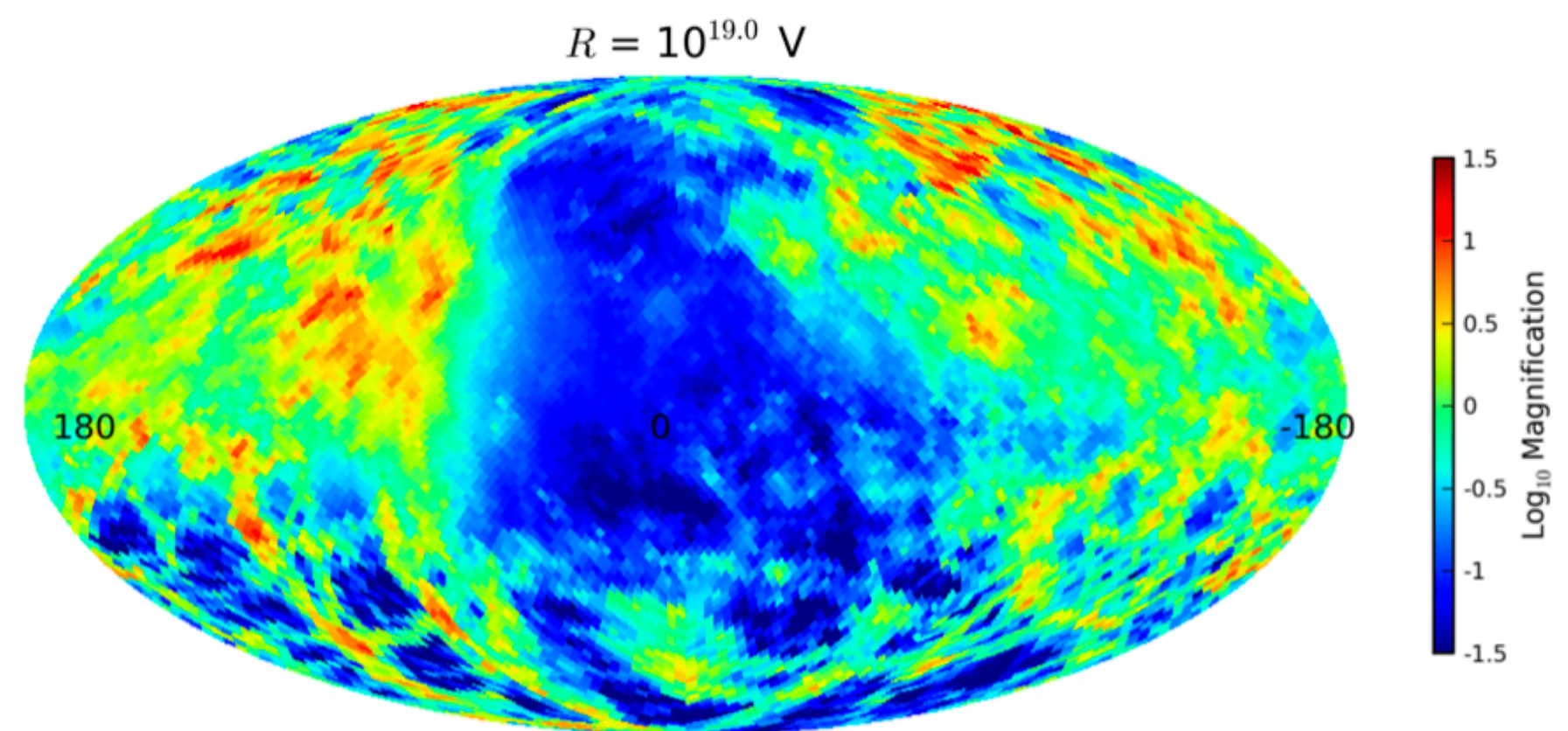
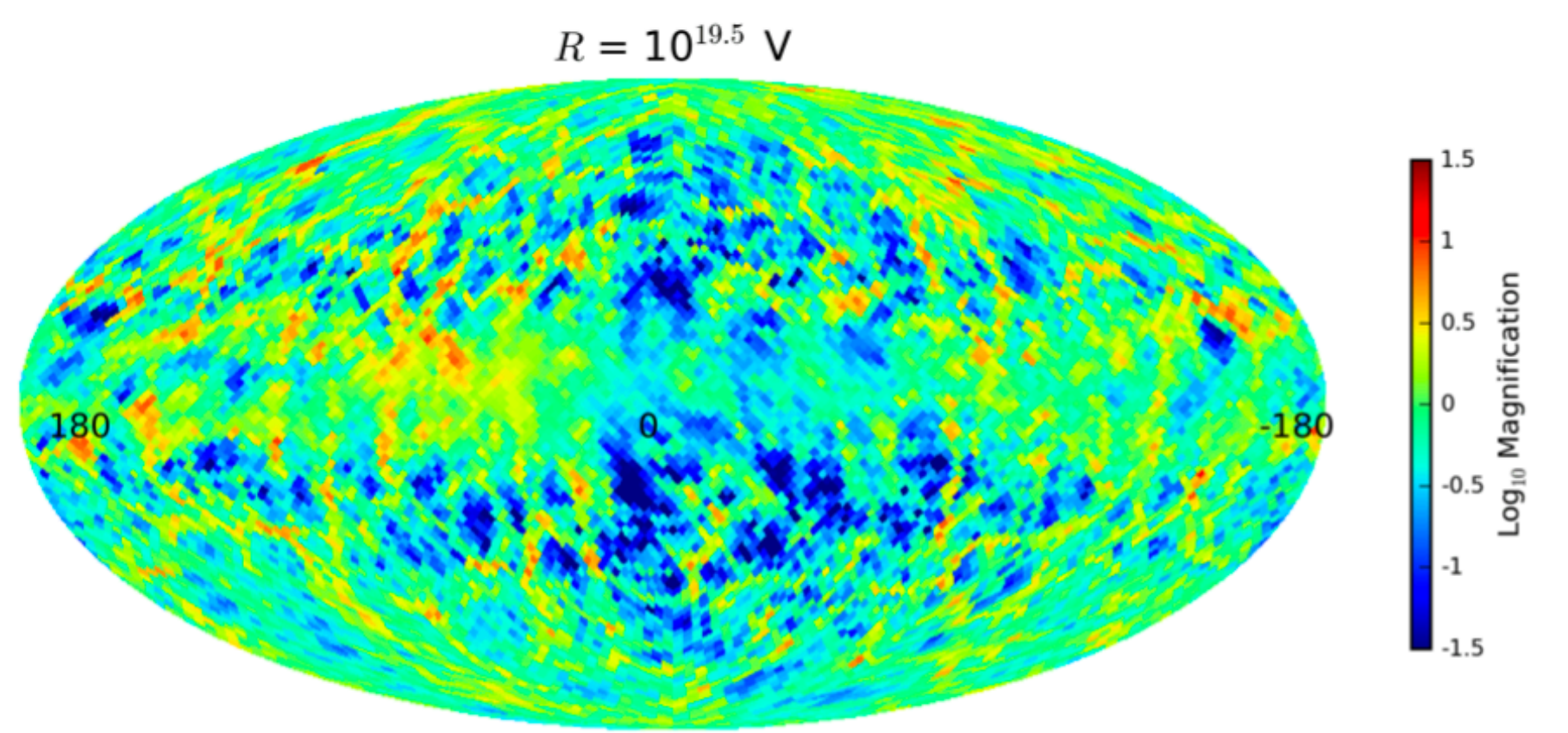
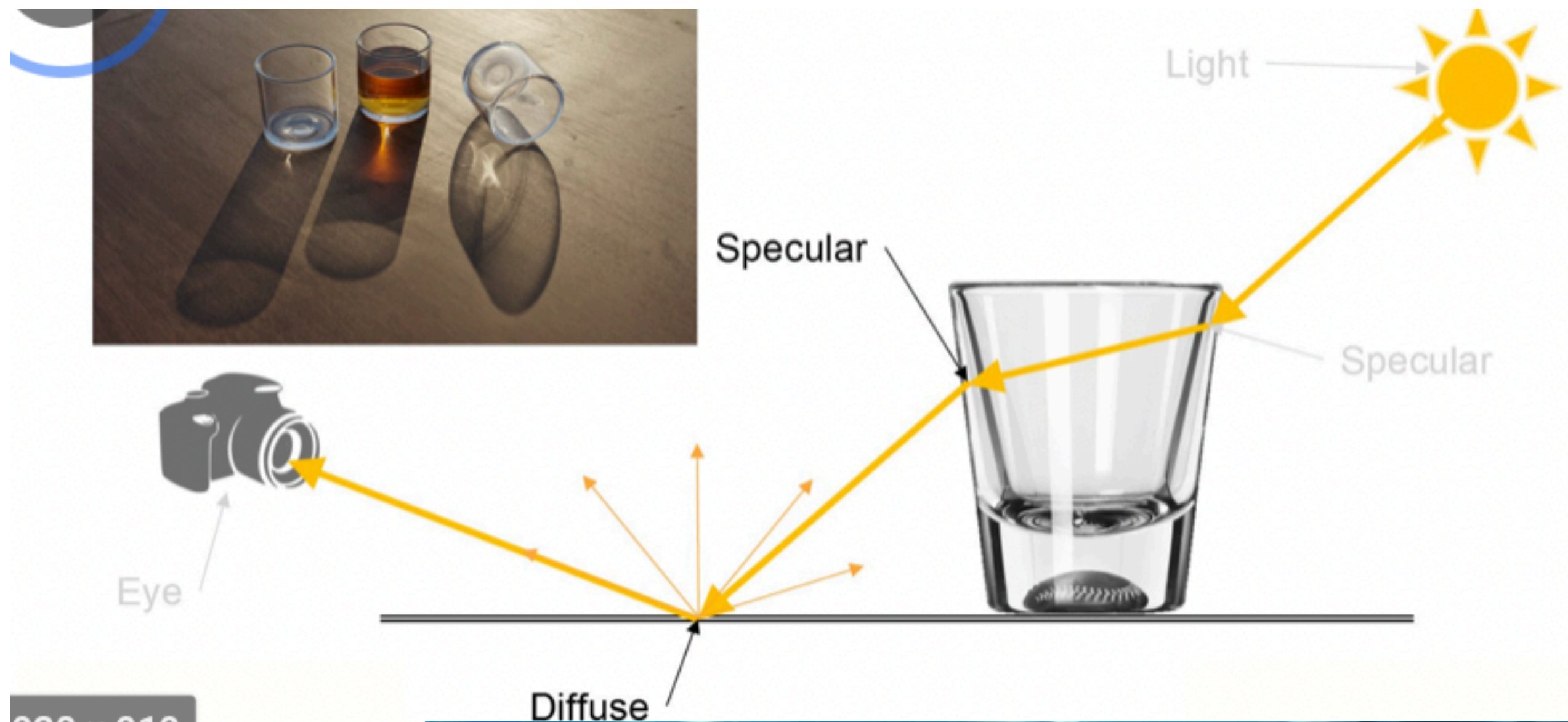
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Published in: *JCAP* 05 (2019) 004 • e-Print: [1711.02730](#) [astro-ph.HE]

[The Galactic Magnetic Field and UHECR Optics](#) #

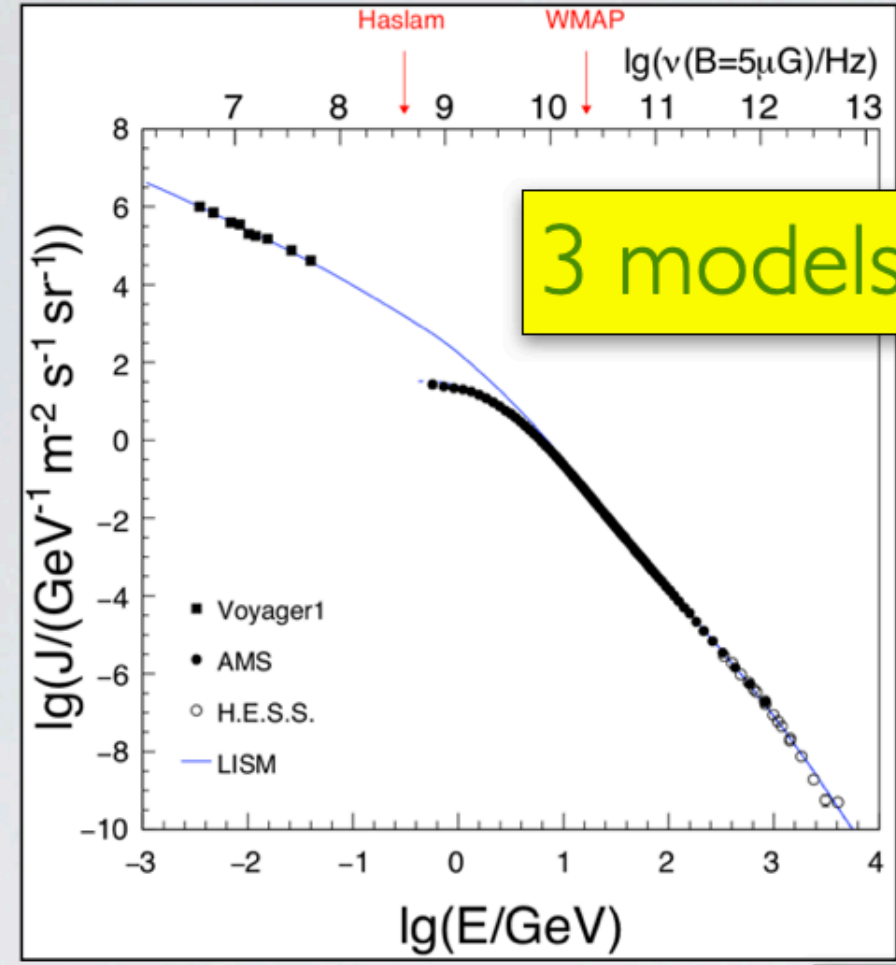
[Glennys R. Farrar \(New York U.\)](#), [Nafiun Awal \(New York U.\)](#), [Deepak Khurana \(New York U.\)](#), [Michael Sutherland \(Ohio State U., CCAPP\)](#) (Aug 19, 2015)

Published in: *PoS ICRC2015* (2016) 560 • Contribution to: [ICRC 2015](#), 560 • e-Print: [1508.04530](#) [astro-ph.HE]



PROGRESS MODELING THE GMF

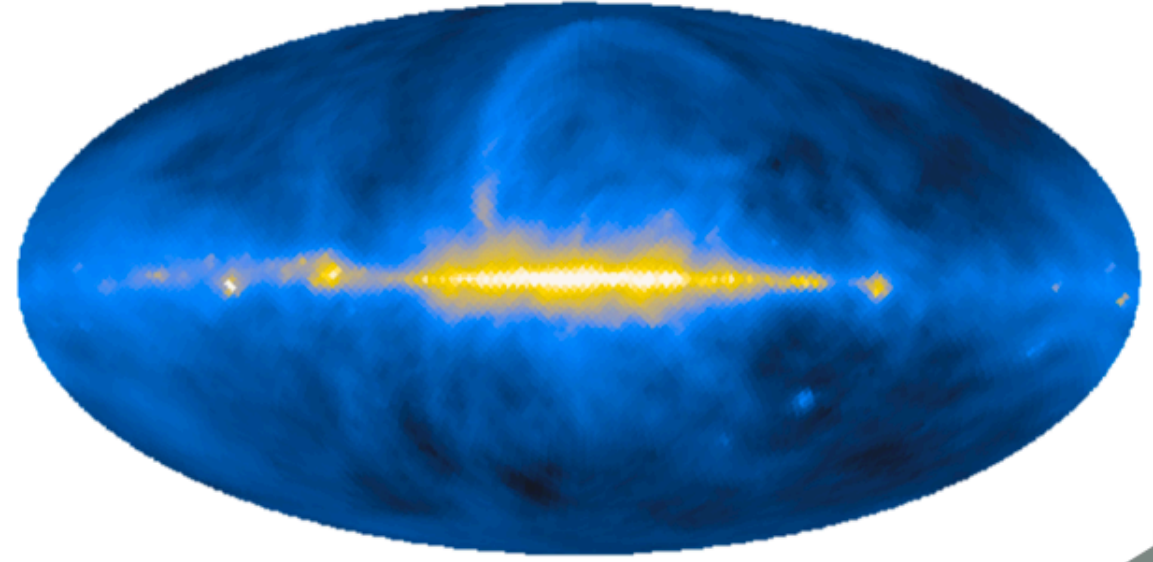
M. Unger
G. Farrar



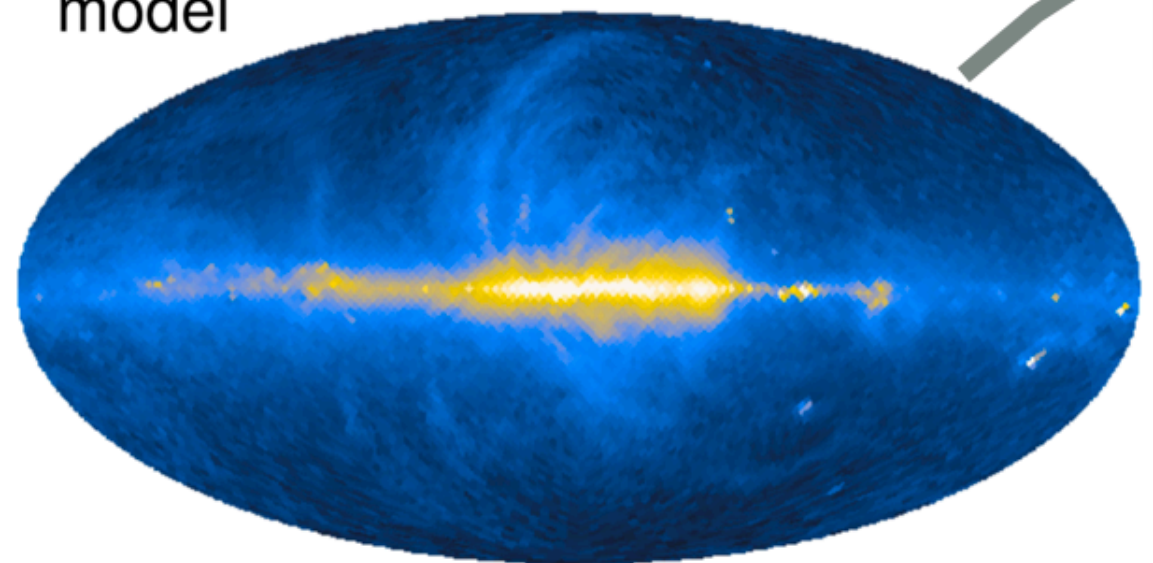
3 models of n_{cre}

Improved determination of synchrotron emission

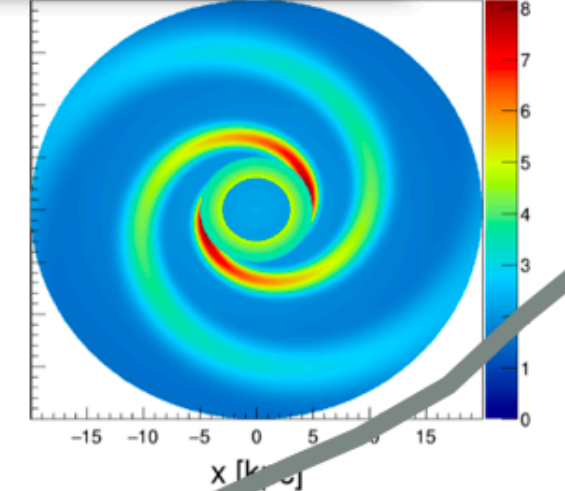
Haslam intensity



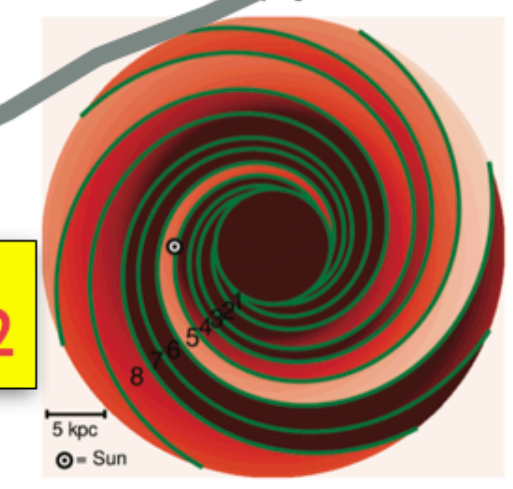
model



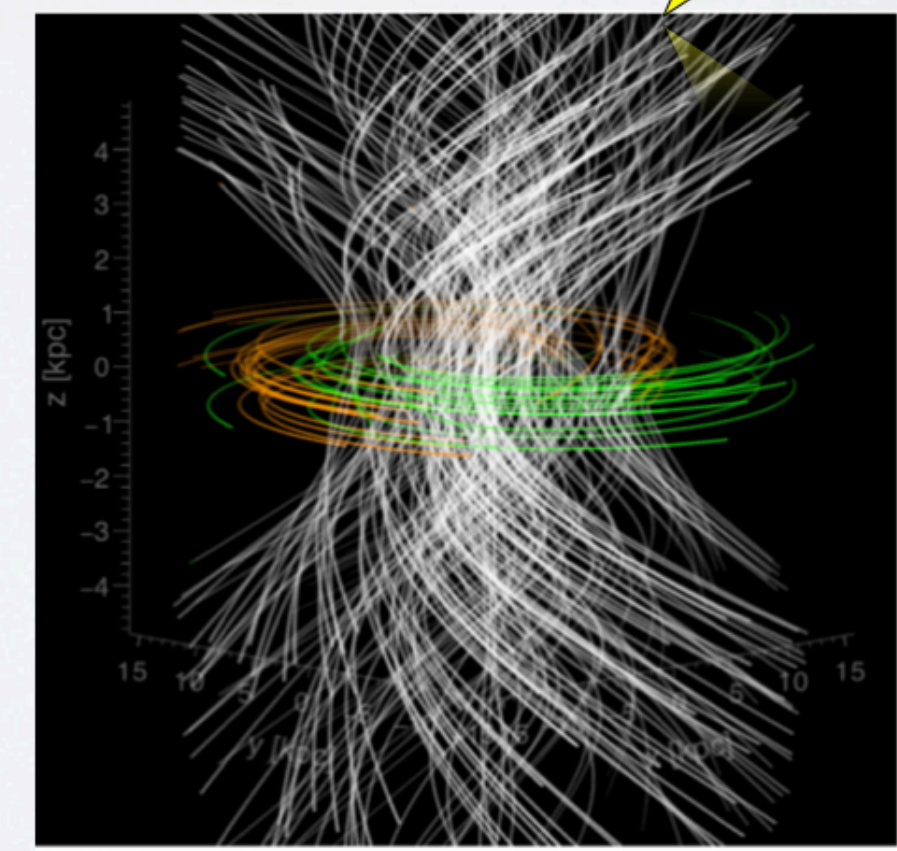
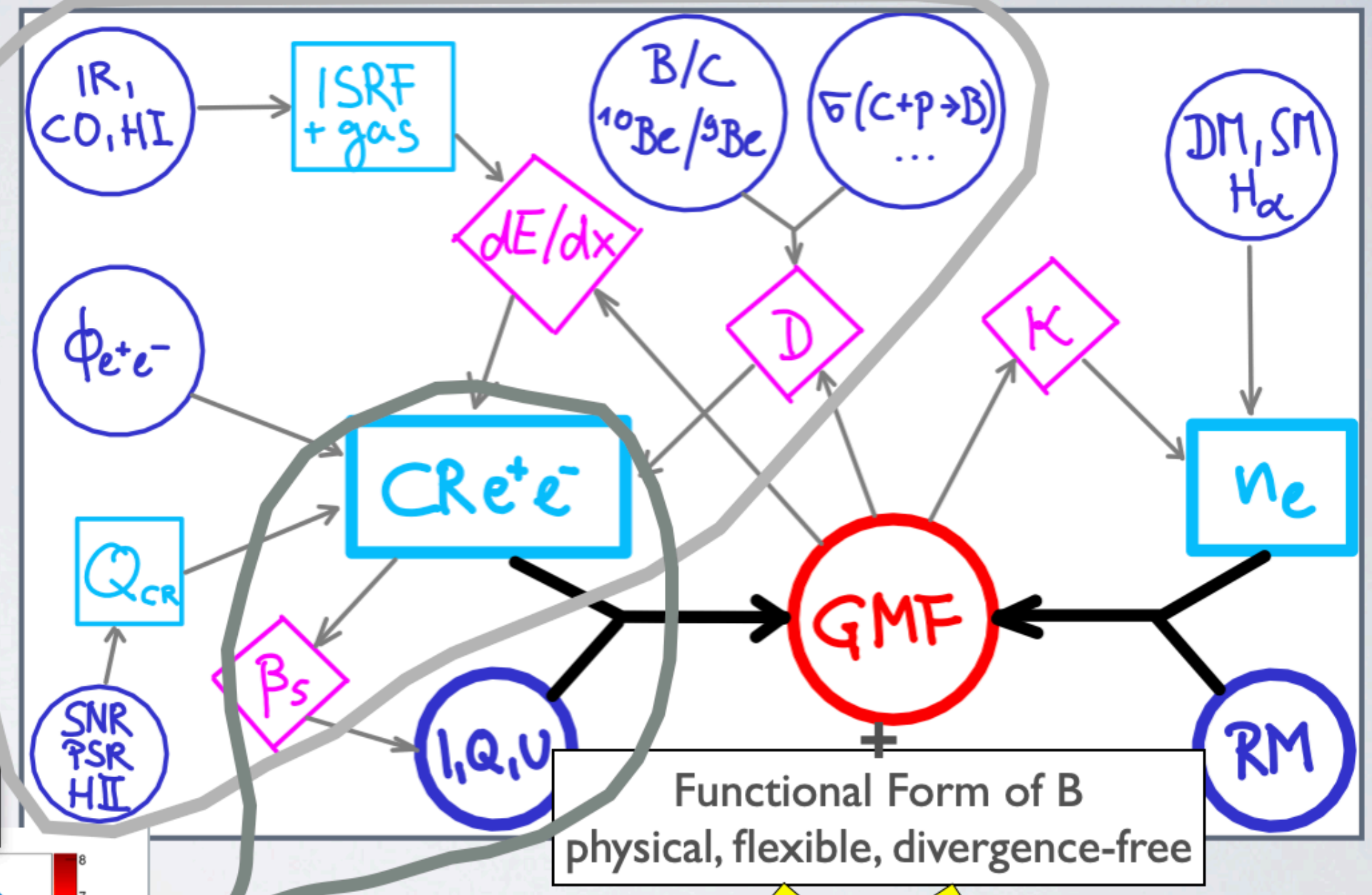
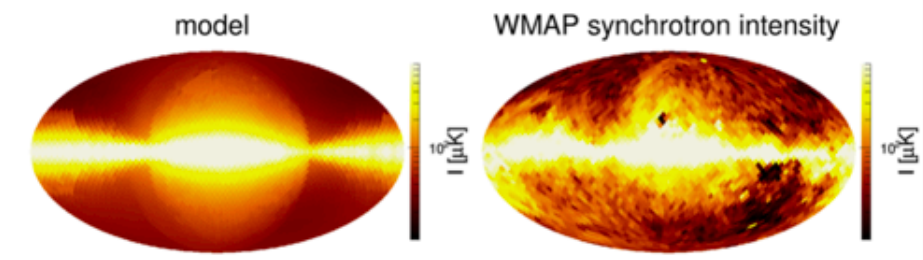
$B_{rand,UF19}$ (prelim)



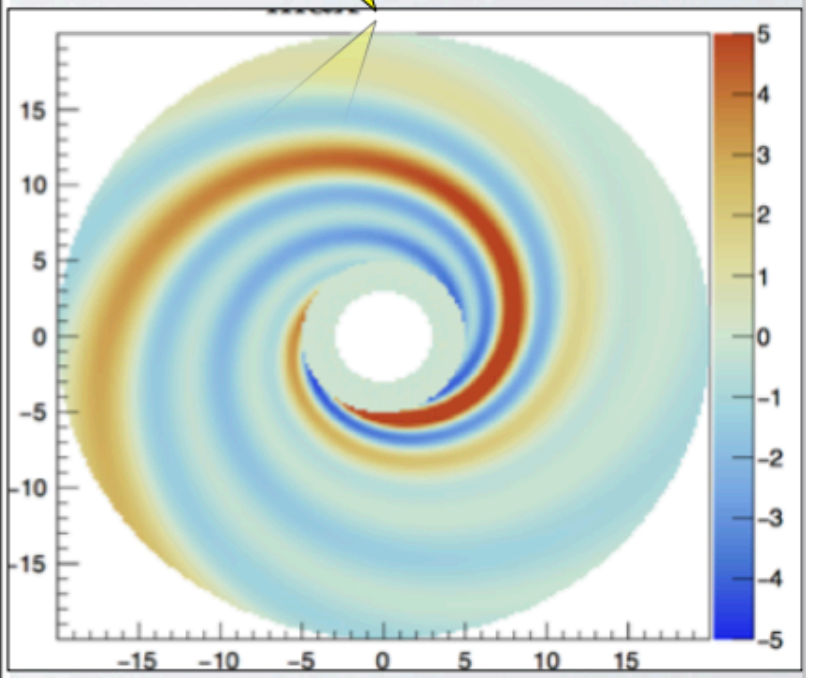
$B_{rand,JF12}$



Analogous comparison JF12



Toroidal field from differential rotation



Fourier spiral

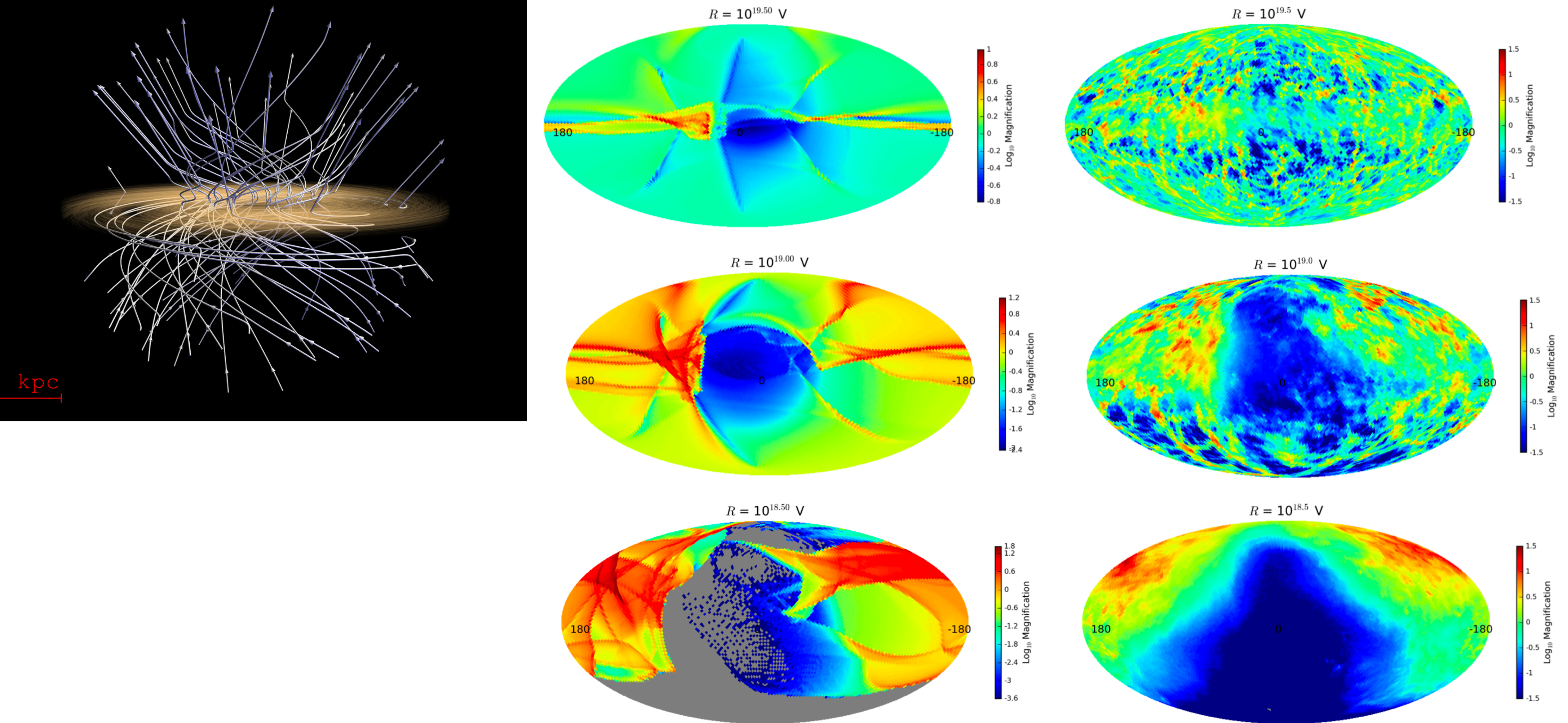
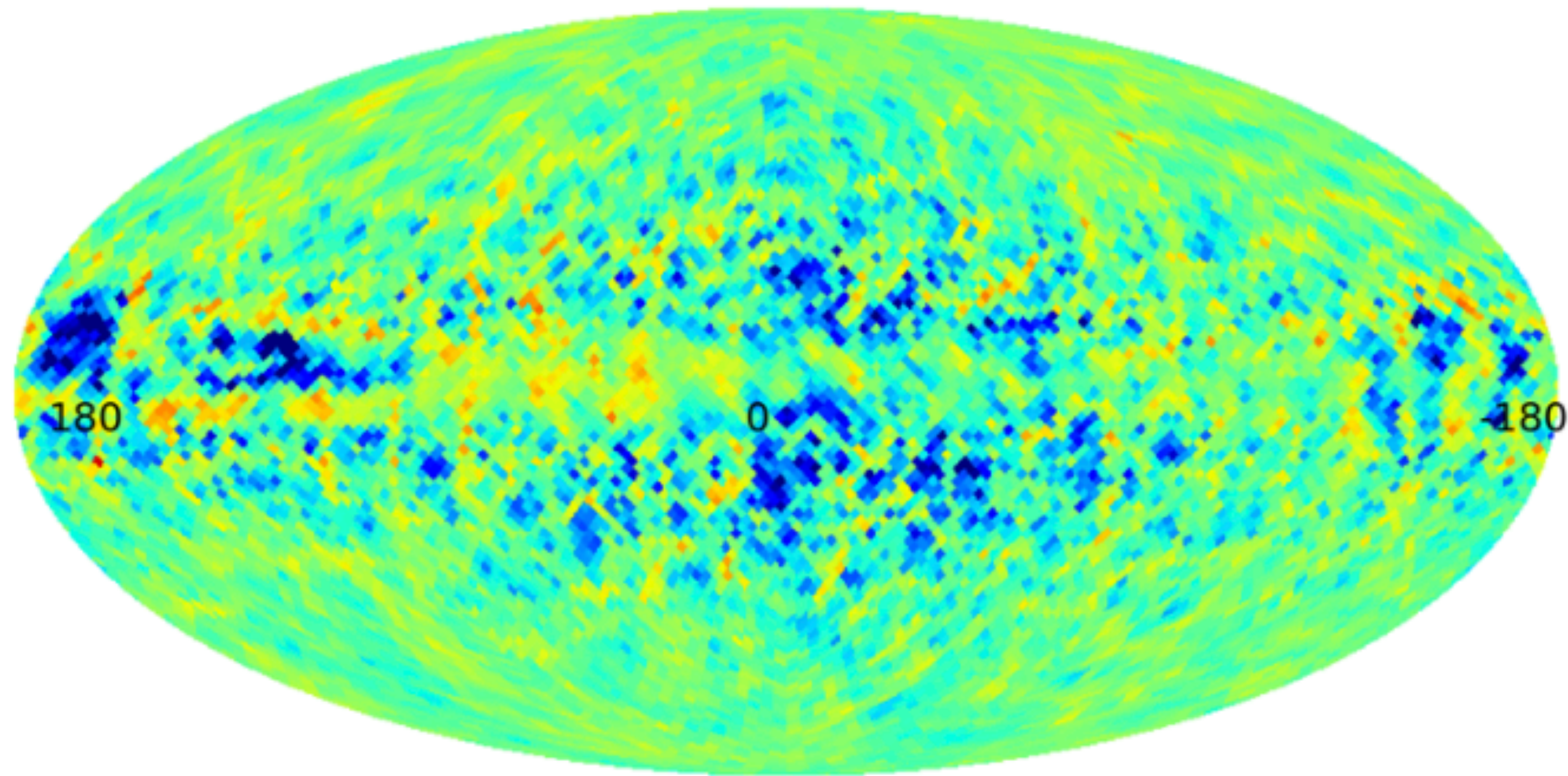
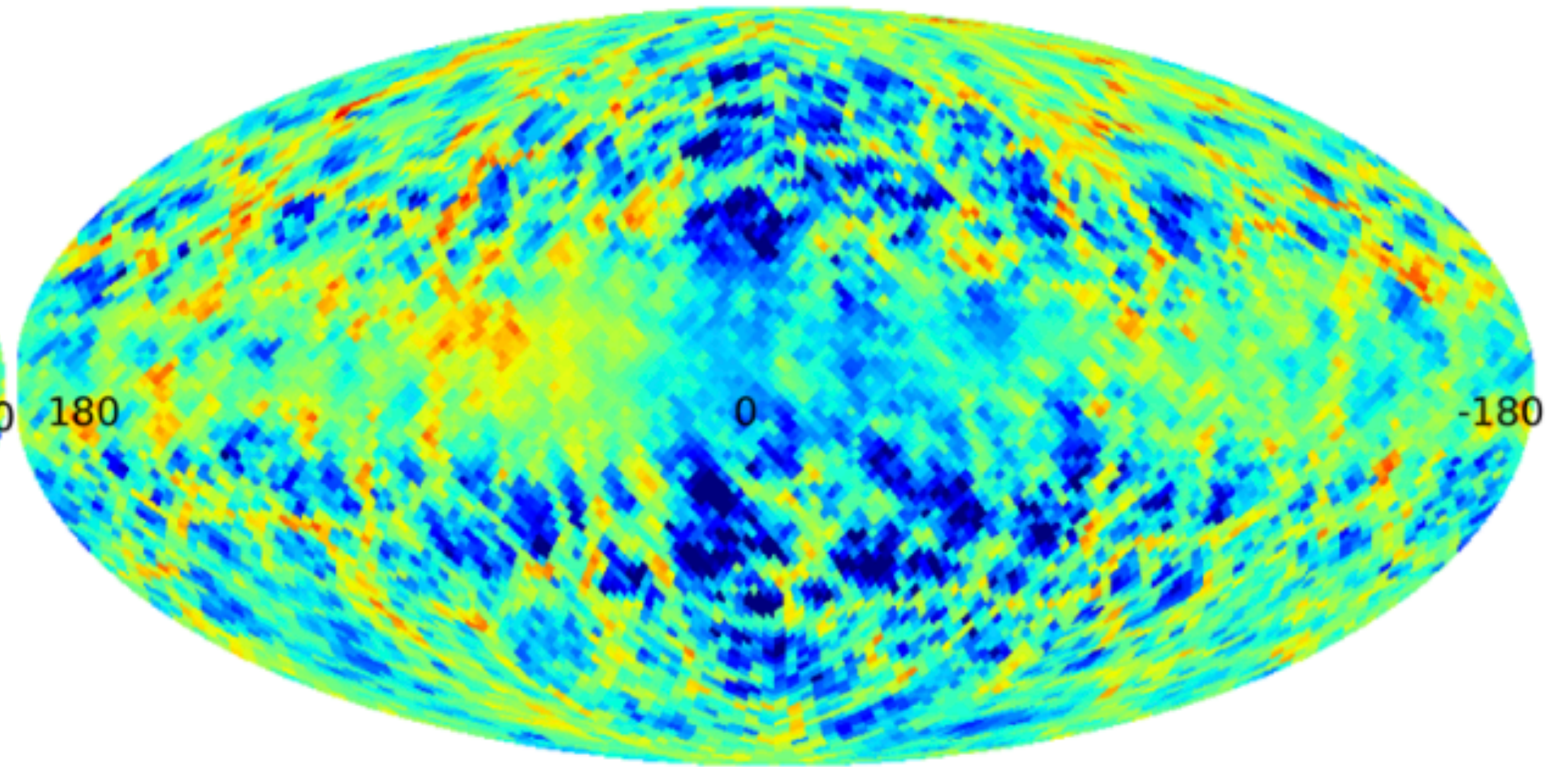


Figure 4. Magnification maps for $\log(R/V)=19.5$, 19.0, and 18.5 from top to bottom. The magnification is defined to be the flux from a standard source in the given direction, summed over its images, relative to the flux in the absence of the GMF. The left column shows the simulation results with only the coherent JF12 component and the right column shows the coherent plus the KRF6 turbulent realization with $L_{\text{coh}} = 100$ pc. The colorbar shows the \log_{10} factor of the magnification of each pixel. The linear features in the coherent-field-only maps on the left result from the abrupt

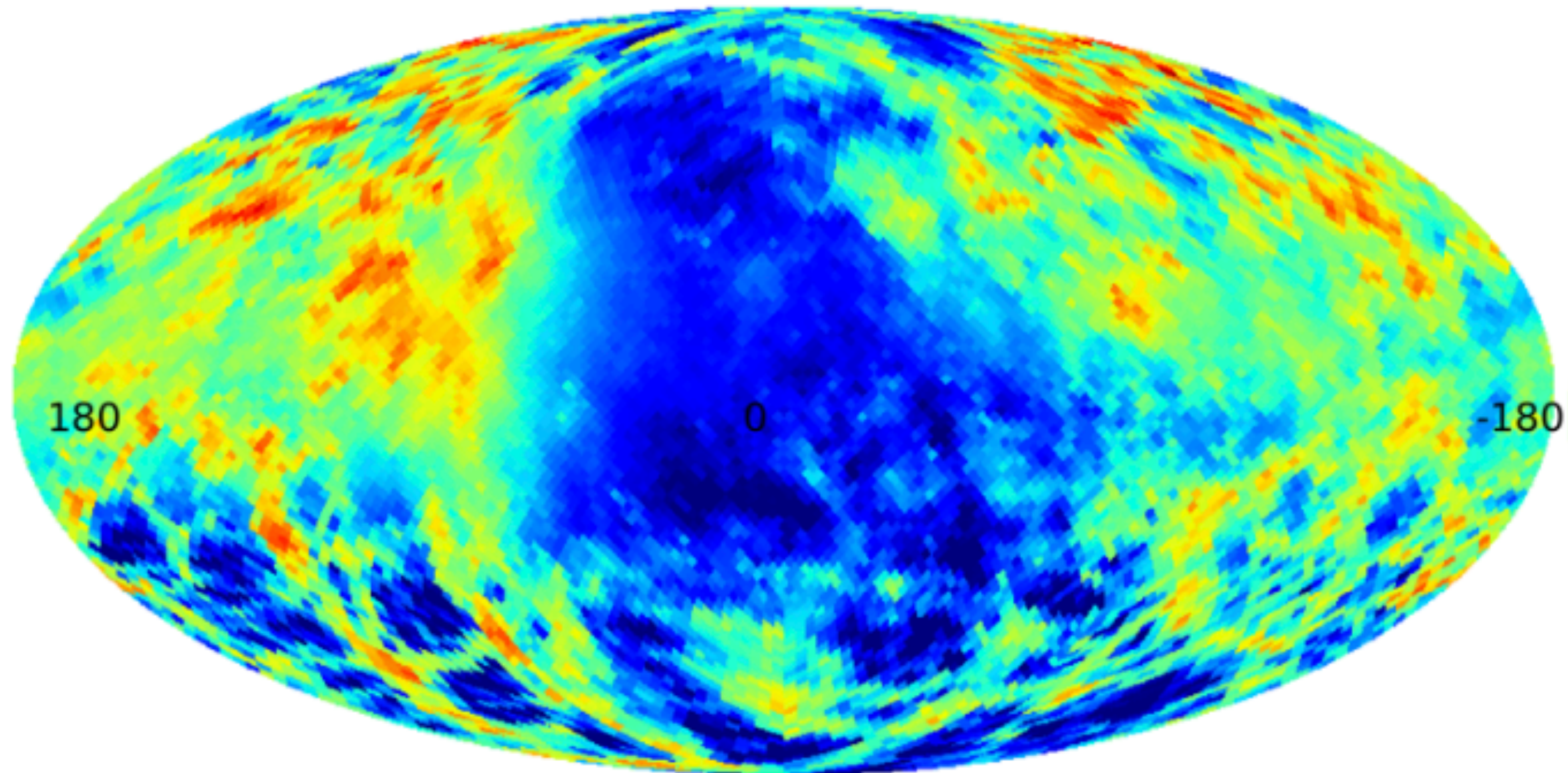
$R = 10^{19.8} \text{ V}$



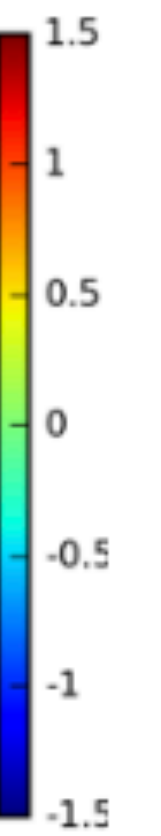
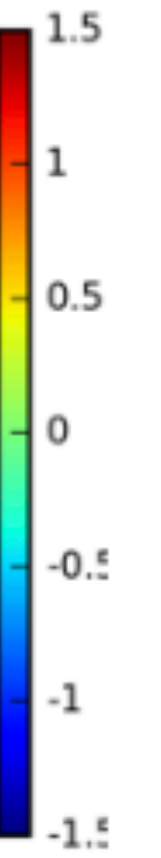
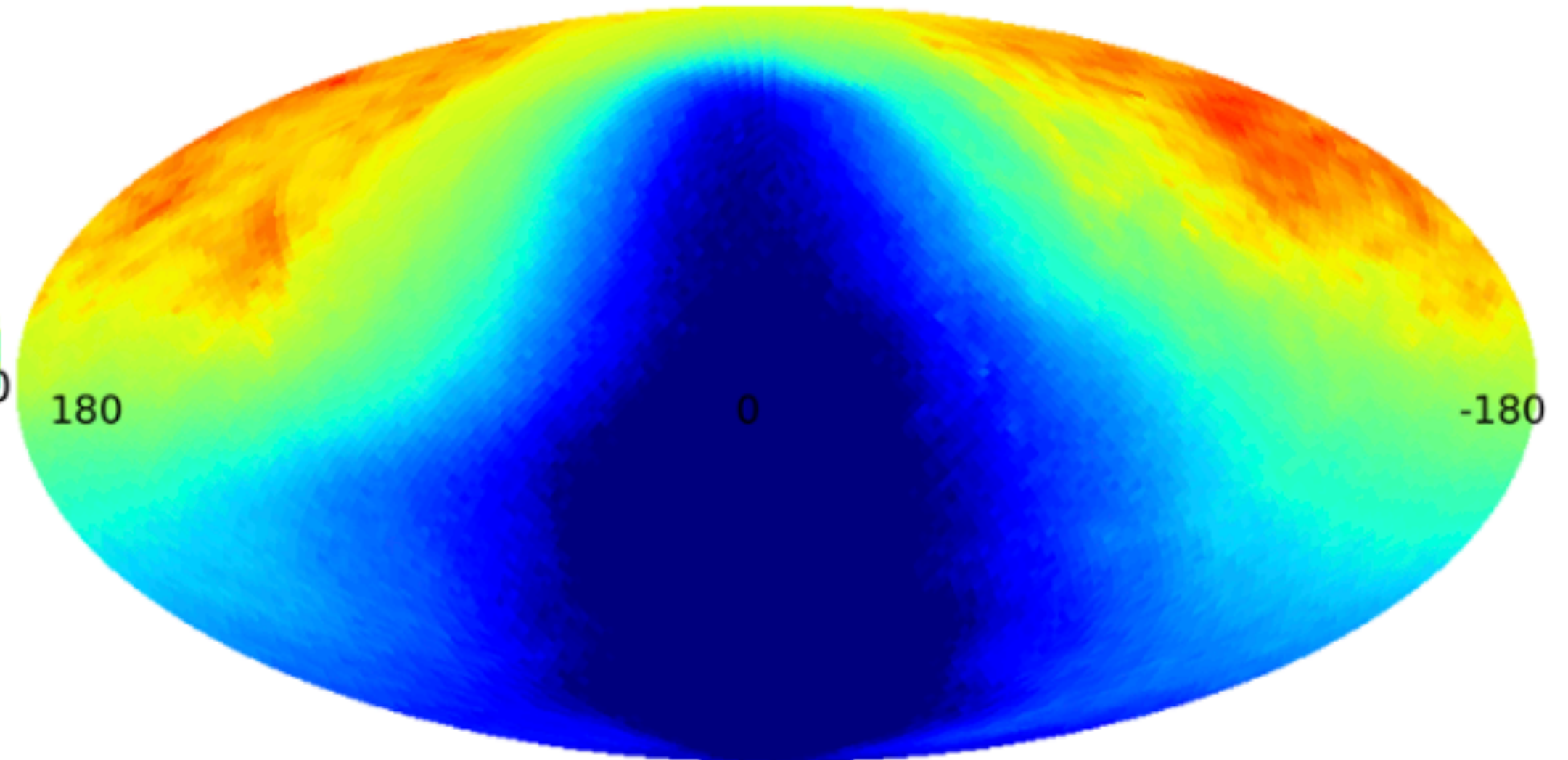
$R = 10^{19.4} \text{ V}$



$R = 10^{19.0} \text{ V}$



$R = 10^{18.3} \text{ V}$

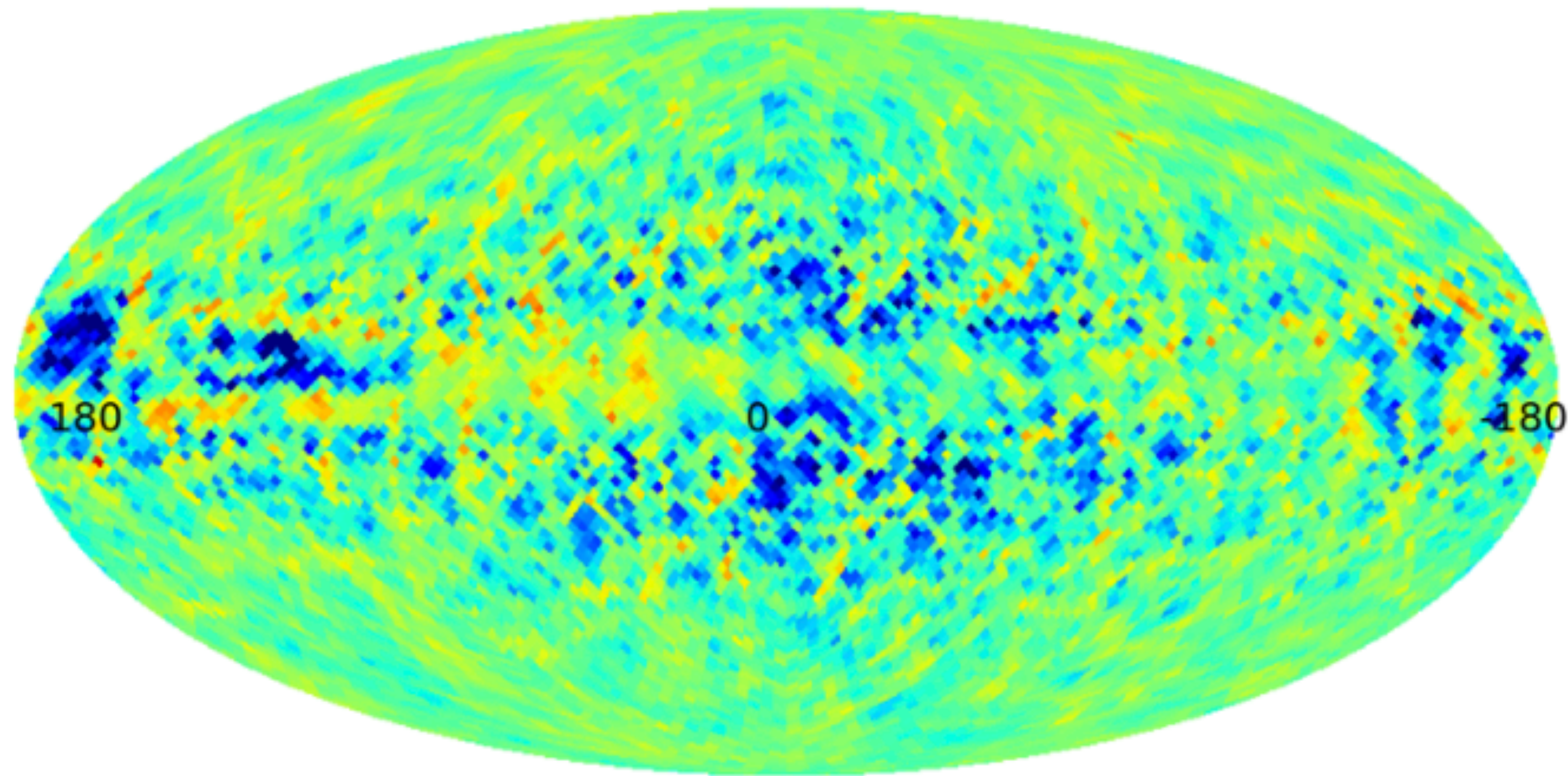


Key Features of UHECR Caustics

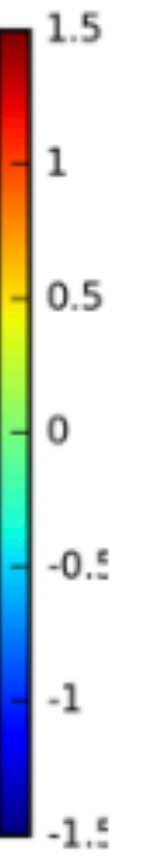
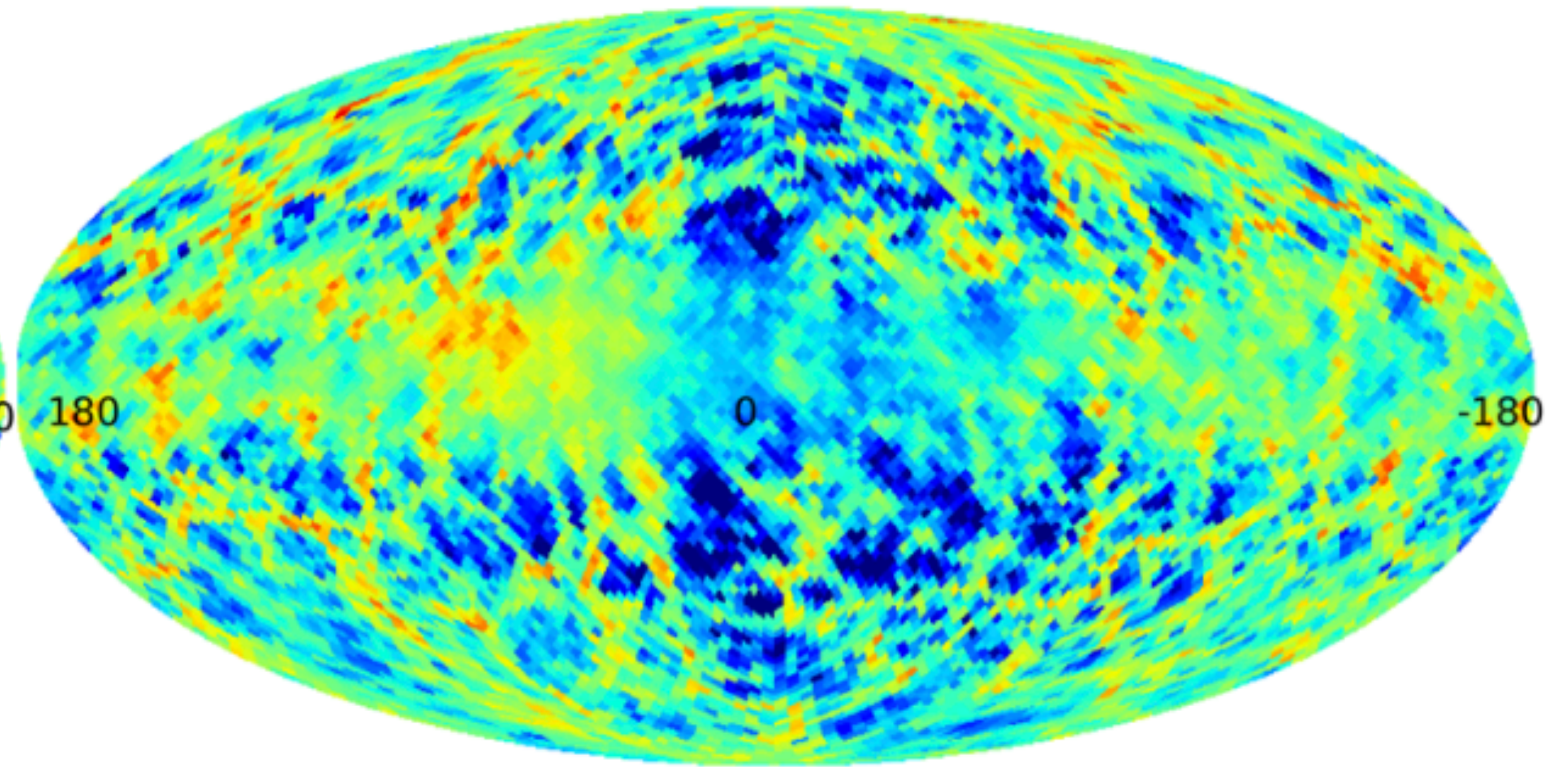
- Magnification rapidly changing in RIGIDITY
 - structure @ 0.1 in $\log_{10} (E/Z)$
 - composition a challenge!
- Present at all rigidities
- Rigidity dependence reveals scale of inhomogeneities
- Opportunity for serendipitous discovery
 - c.f. very high z sources via gravitational lensing

logR	mag	$\langle \ell \rangle$ ($^\circ$)	$\langle b \rangle$ ($^\circ$)	$\langle \Delta \rangle, \sigma_\Delta$ ($^\circ$)
HPX 2: $\ell = 135^\circ, b = 41.8103^\circ$				
20.0	1.31 (0.88) {1.05}	132 (135) {135}	42 (42) {42}	2.1, 0.8 (0.5, 0.7) {0.6, 0.7}
19.8	1.18 (0.76) {1.07}	131 (134) {134}	43 (42) {43}	3.4, 0.8 (0.7, 0.6) {0.9, 0.7}
19.6	0.77 (0.71) {1.12}	128 (134) {134}	43 (43) {43}	5.4, 0.6 (1.5, 0.8) {1.5, 0.7}
19.5	0.54 (1.06) {1.16}	126 (133) {134}	43 (44) {43}	6.5, 0.5 (2.1, 0.8) {2.0, 0.8}
19.4	0.39 (1.23) {1.21}	125 (133) {133}	43 (44) {44}	7.8, 0.5 (2.5, 0.8) {2.6, 0.8}
19.2	0.07 (1.71) {1.37}	119 (131) {132}	42 (45) {46}	11.7, 2.2 (4.3, 1.3) {4.5, 0.8}
19.0	3.4 (1.45) {1.76}	146 (128) {128}	25 (49) {49}	19.0, 12.4 (8.4, 1.9) {8.6, 0.9}
18.8	6.64 (3.99) {4.59}	127 (101) {113}	27 (61) {58}	16.4, 23.6 (27.8, 8.1) {21.5, 2.3}
18.6	5.43 (2.37) {1.74}	132 (111) {102}	11 (-34) {-46}	30.8, 29.0 (79.2, 22.0) {92.7, 7.4}
18.5	3.81 (2.72) {2.68}	134 (103) {127}	-6 (-27) {-27}	48.0, 50.5 (74.6, 36.8) {68.9, 7.4}
18.4	4.92 (3.31) {3.01}	157 (99) {106}	-9 (-25) {-28}	54.8, 44.1 (74.9, 40.3) {75.0, 27.0}
18.3	3.24 (5.07) {5.53}	180 (111) {77}	-20 (-25) {-24}	74.8, 69.6 (70.3, 47.5) {85.2, 53.7}
18.2	4.19 (4.36) {1.1}	-167 (91) {11}	-12 (-29) {7}	75.6, 65.3 (81.3, 56.3) {109.4, 46.7}
18.1	7.75 (5.0) {8.01}	-162 (-154) {41}	8 (-26) {-55}	65.1, 49.7 (94.3, 70.9) {125.1, 25.6}
18.0	5.42 (3.53) {2.06}	152 (34) {23}	5 (1) {21}	39.4, 85.0 (97.8, 53.2) {91.2, 34.2}
HPX 3: $\ell = -135^\circ, b = 41.8103^\circ$				
20.0	1.14 (1.06) {1.01}	-135 (-135) {-135}	41 (41) {42}	1.3, 0.8 (0.4, 0.7) {0.4, 0.7}
19.8	1.15 (1.17) {1.01}	-136 (-135) {-135}	40 (41) {41}	2.0, 0.8 (0.6, 0.8) {0.6, 0.7}
19.6	1.04 (1.22) {1.02}	-136 (-136) {-136}	39 (41) {41}	3.1, 0.7 (0.9, 0.8) {0.9, 0.7}
19.5	1.26 (1.16) {1.03}	-137 (-136) {-136}	38 (41) {41}	3.6, 0.8 (1.2, 0.8) {1.1, 0.7}
19.4	1.27 (1.09) {1.04}	-137 (-136) {-136}	38 (41) {41}	4.2, 0.9 (1.4, 0.9) {1.4, 0.7}
19.2	3.63 (0.76) {1.06}	-139 (-137) {-137}	38 (40) {40}	4.7, 4.9 (2.2, 1.0) {2.3, 0.7}
19.0	2.9 (0.76) {1.11}	-135 (-138) {-138}	22 (39) {39}	19.9, 11.5 (3.7, 2.1) {3.9, 0.7}
18.8	2.25 (1.65) {1.2}	-141 (-141) {-141}	7 (35) {37}	34.9, 19.3 (8.0, 3.5) {6.5, 0.8}
18.6	2.53 (3.48) {2.07}	-134 (-142) {-140}	6 (7) {11}	36.3, 25.5 (35.4, 28.5) {30.9, 22.8}
18.5	3.73 (4.55) {4.54}	-159 (-136) {-136}	23 (1) {-1}	27.1, 53.4 (41.0, 33.1) {43.2, 32.4}
18.4	3.93 (4.89) {5.33}	-36 (-130) {-136}	11 (4) {5}	89.7, 79.3 (37.7, 37.4) {37.3, 23.6}
18.3	6.24 (5.32) {3.62}	60 (-93) {-38}	32 (4) {-4}	104.4, 57.5 (53.0, 64.0) {98.0, 30.4}
18.2	5.77 (3.82) {1.81}	79 (-109) {-60}	34 (30) {53}	98.1, 44.0 (24.4, 47.3) {49.8, 8.0}
18.1	1.71 (3.24) {2.04}	93 (-118) {-97}	16 (33) {60}	107.4, 75.4 (16.2, 44.3) {29.5, 9.3}
18.0	2.47 (2.89) {2.25}	67 (-123) {-138}	4 (23) {51}	130.2, 83.6 (21.8, 54.8) {9.4, 12.2}

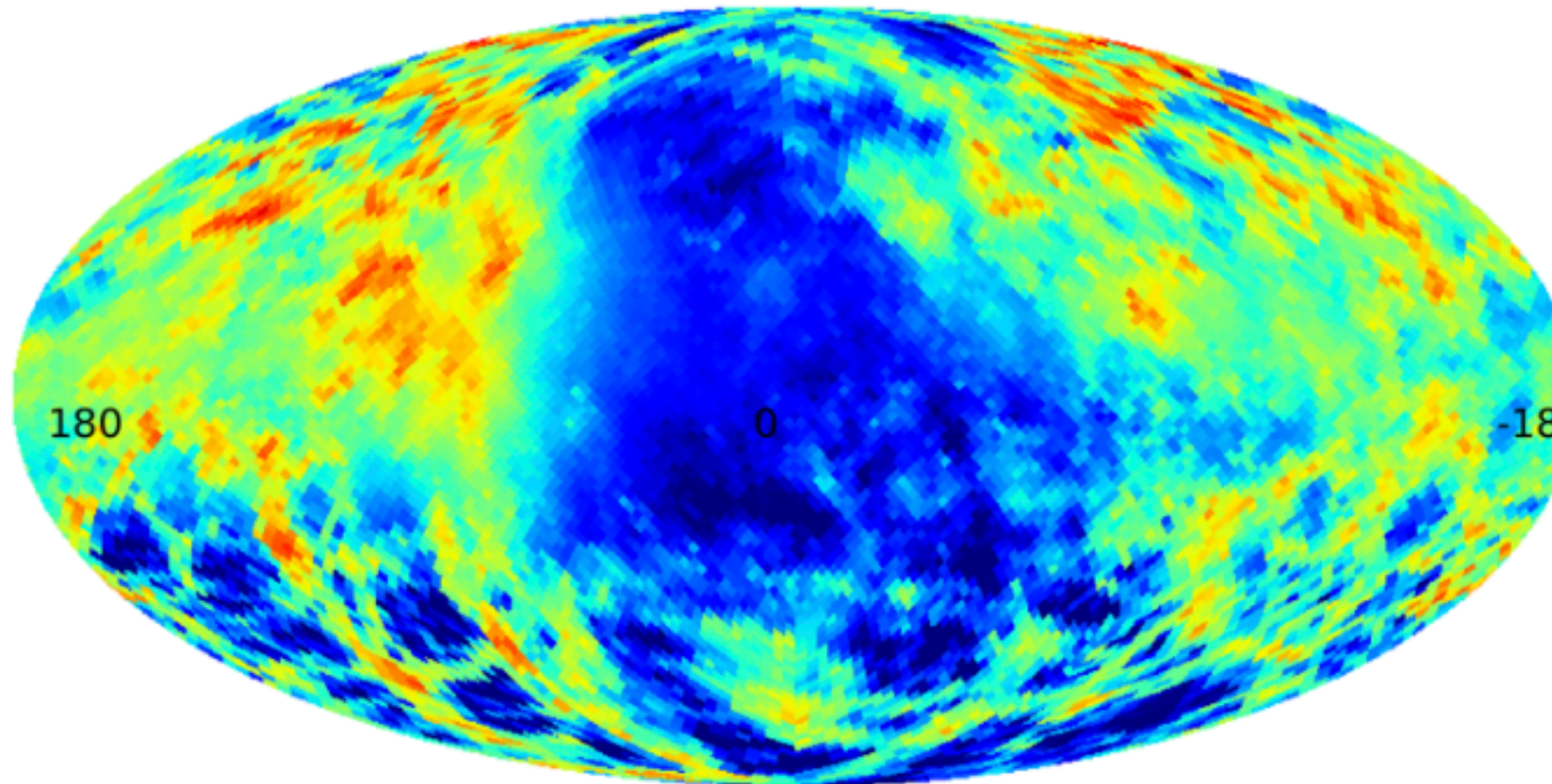
$R = 10^{19.8} \text{ V}$



$R = 10^{19.4} \text{ V}$



$R = 10^{19.0} \text{ V}$



$R = 10^{18.3} \text{ V}$

