



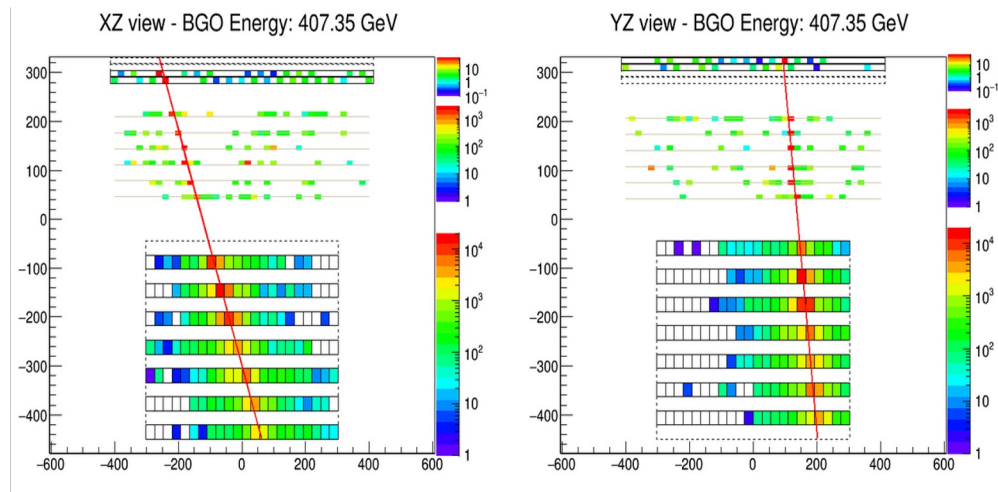
## Lesson 2

Direct measurement  
of galactic cosmic rays

The example of DAMPE

Paolo Bernardini

Università del Salento and INFN, Lecce, Italy



## To do:

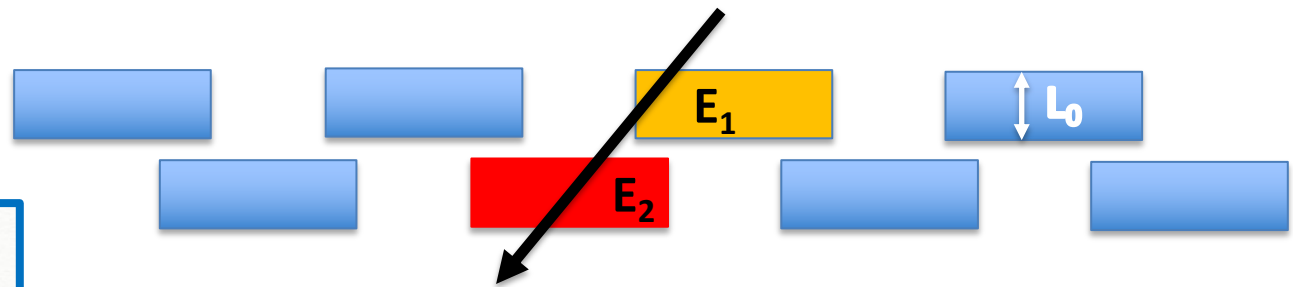
- Reconstruct the track in STK and BGO
- Distinguish the primary charge ( $Z$ )
- Select a "clean" sample of events
- Estimate event energy ( $E_{\text{BGO}} \rightarrow E_{\text{Primary}}$ )
- Measure the live time ( $\Delta t$ )
- Estimate effective acceptance by means of simulation
- Validate simulation by comparison with data

# PSD - Charge measurement

Bethe-Bloch formula for energy loss by ionization and atomic excitation

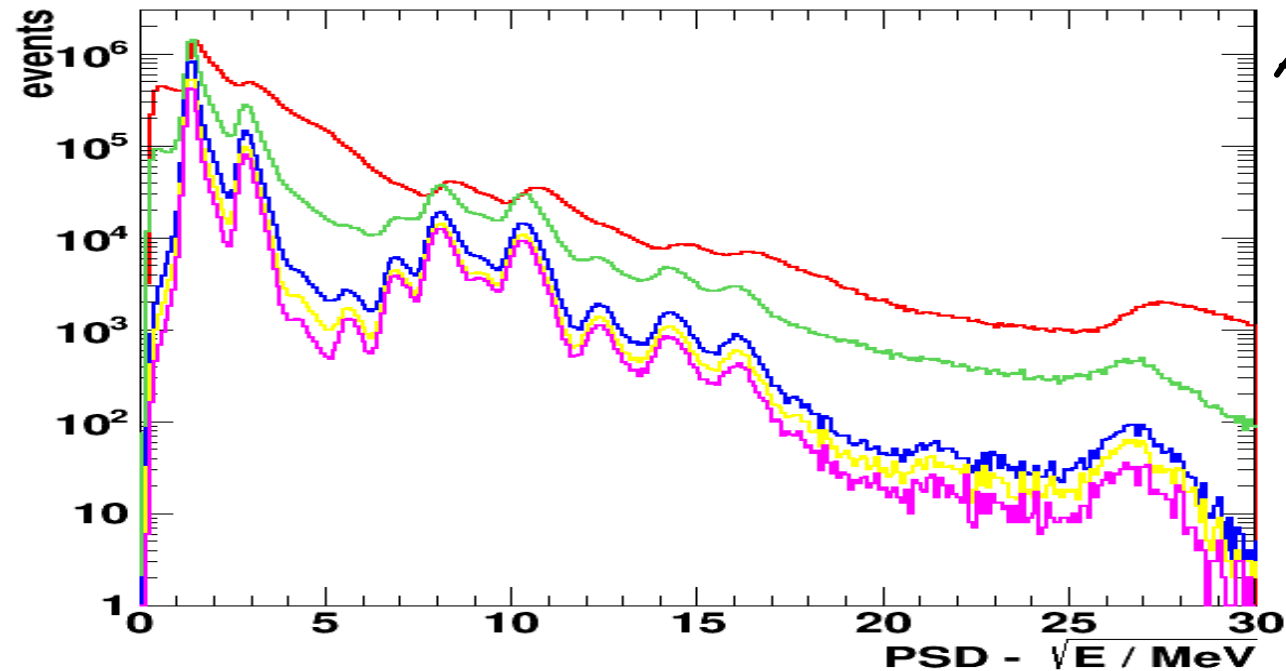
$$-\frac{dE}{dx} = K z^2 \frac{Z}{A} \frac{1}{\beta^2} \left[ \frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2 T_{\max}}{I^2} - \beta^2 - \frac{\delta}{2} \right]$$

$$Z \propto \sqrt{\frac{E_1 + E_2}{L_1 + L_2} L_0}$$



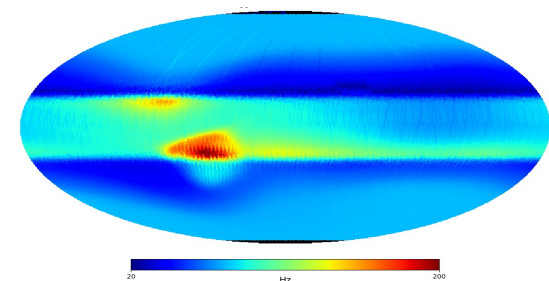
- $L_0$  path in PSD for a vertical trajectory
- $L_1$  path in the 1<sup>st</sup> intercepted PSD bar
- $L_2$  path in the 2<sup>nd</sup> intercepted PSD bar

Z-measurement depends on the quality of the track reconstruction



An example of analysis cuts to get a "clean" sample

Applied to real and simulated data



Preselection:

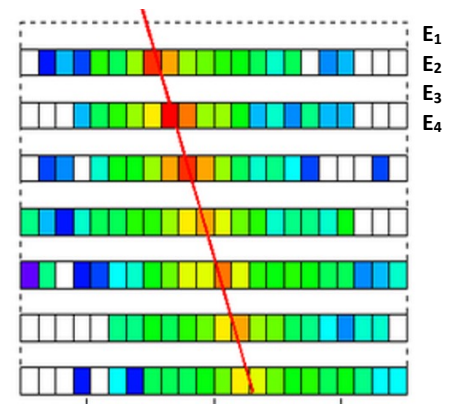
$E_{BGO} > 20 \text{ GeV}$     High Energy Trigger    No events from SAA

Quality of STK track:     $\chi^2/\text{dof} < 35$     maximum signal

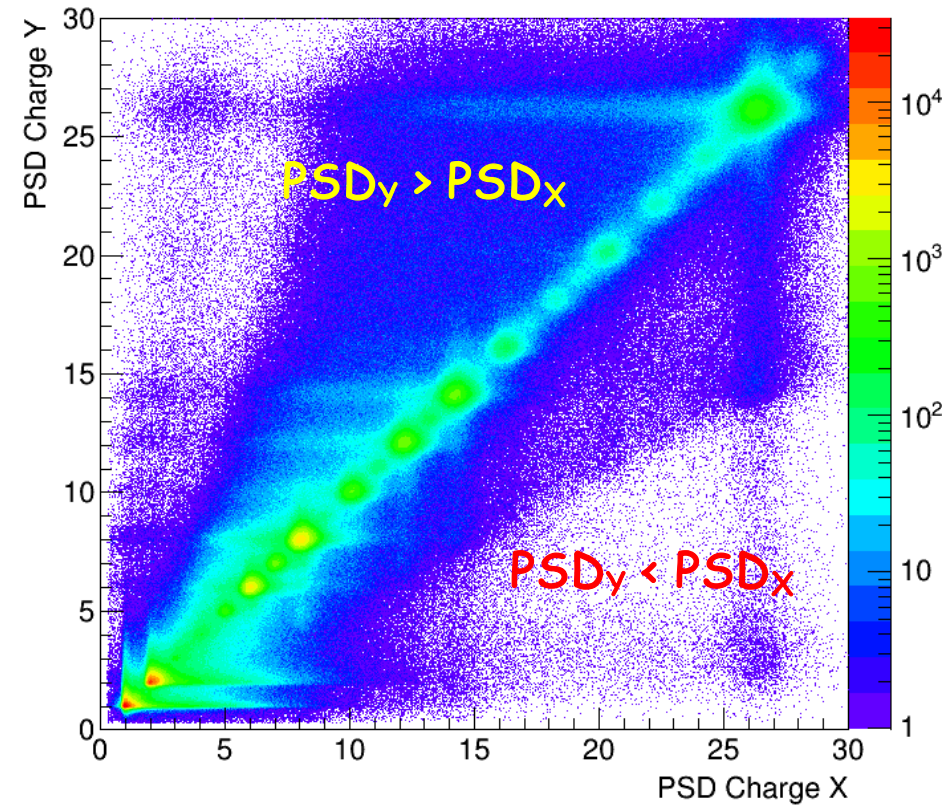
BGO - STK tracks agreement

BGO signal:     $E_{MAX} < 0.35 E_{TOT}$      $E_1 + E_2 < E_3 + E_4$

Minor analysis cuts





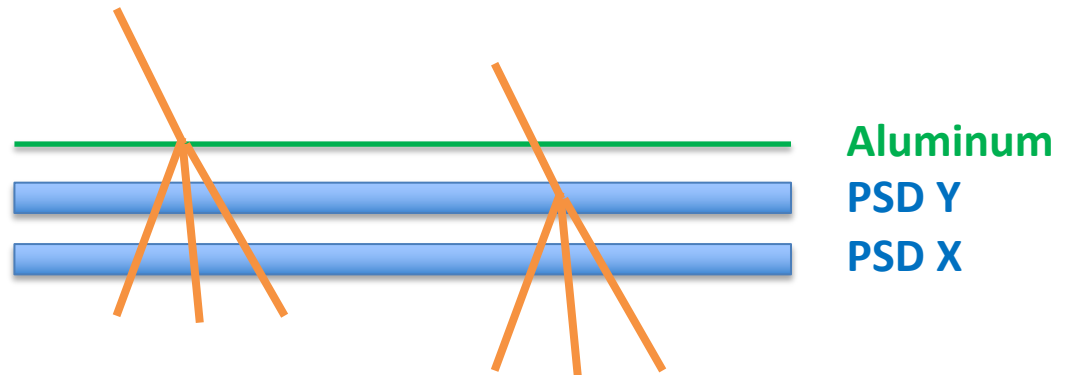


## Coherent signals

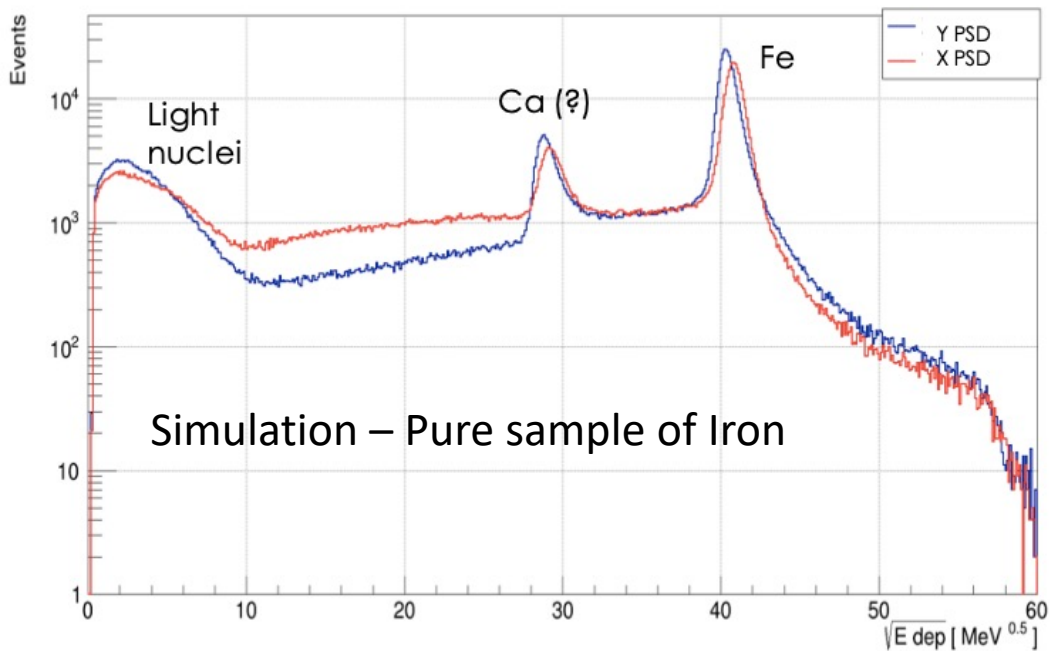
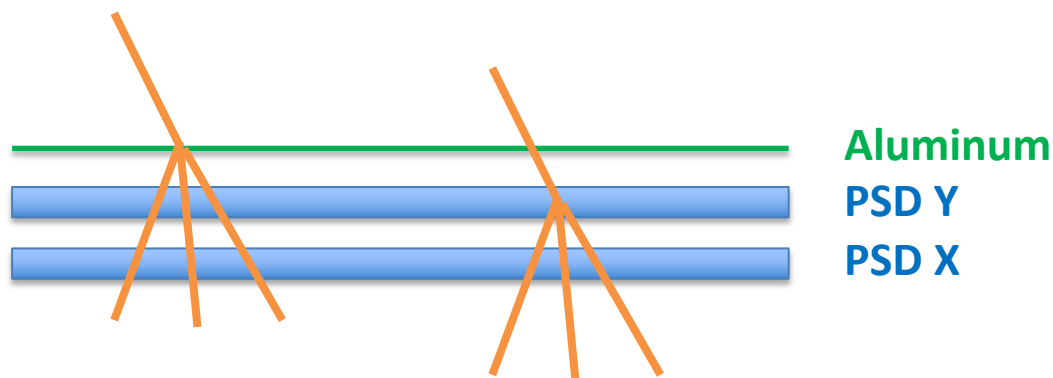
- $PSD_y \sim PSD_x$  (bisector)

## Anomalies

- $PSD_y > PSD_x$  fragmentation
- $PSD_y < PSD_x$  more particles on same x-bar

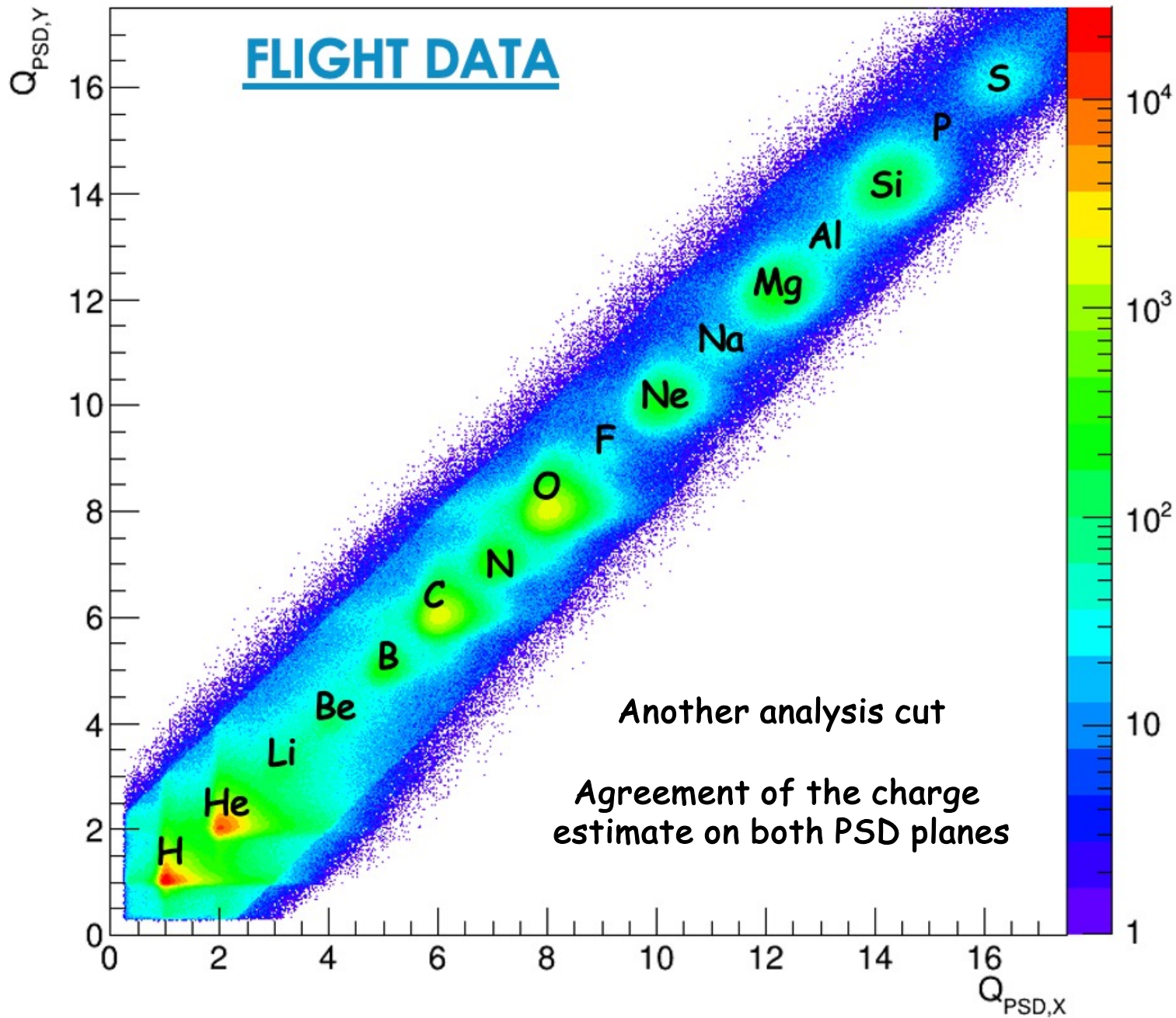


# Check on fragmentation

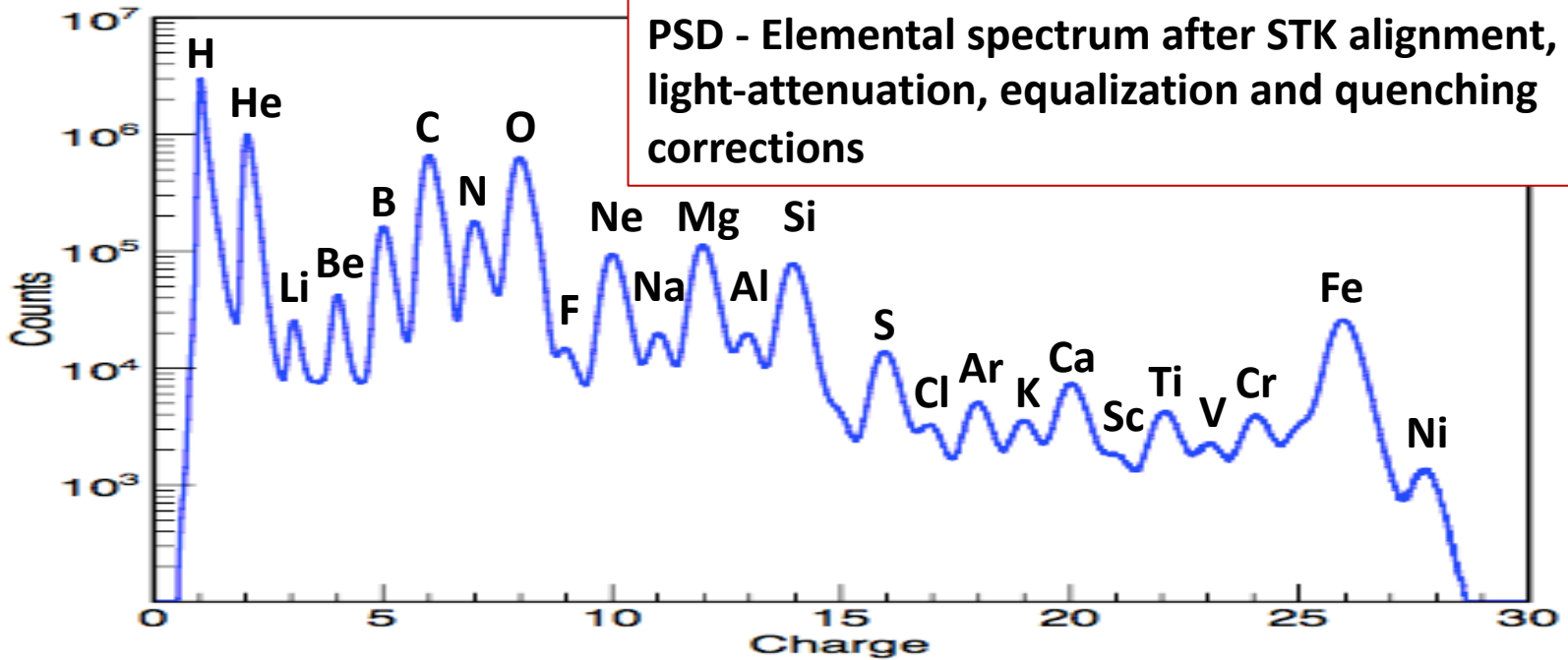


The hadron interaction in PSD or STK is a problem

Analysis is optimized for interaction in the BGO



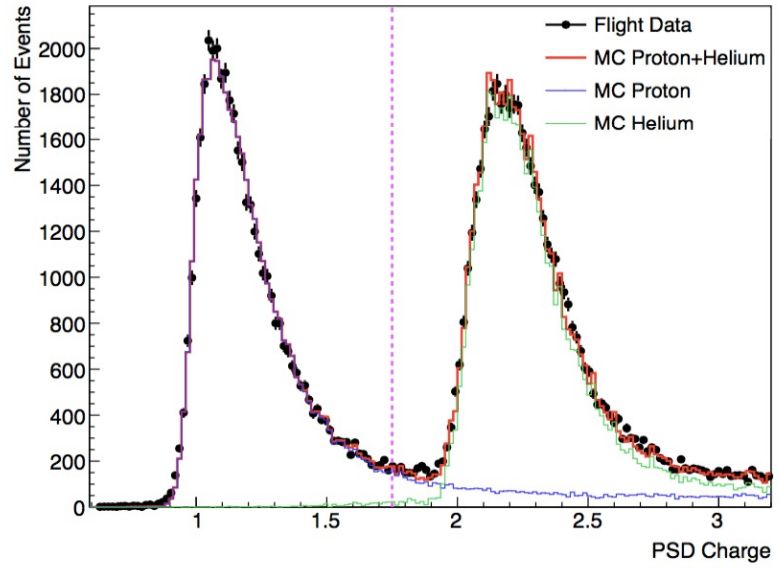
# Nuclei (Z=1-26 and more)



PSD - Elemental spectrum after STK alignment, light-attenuation, equalization and quenching corrections

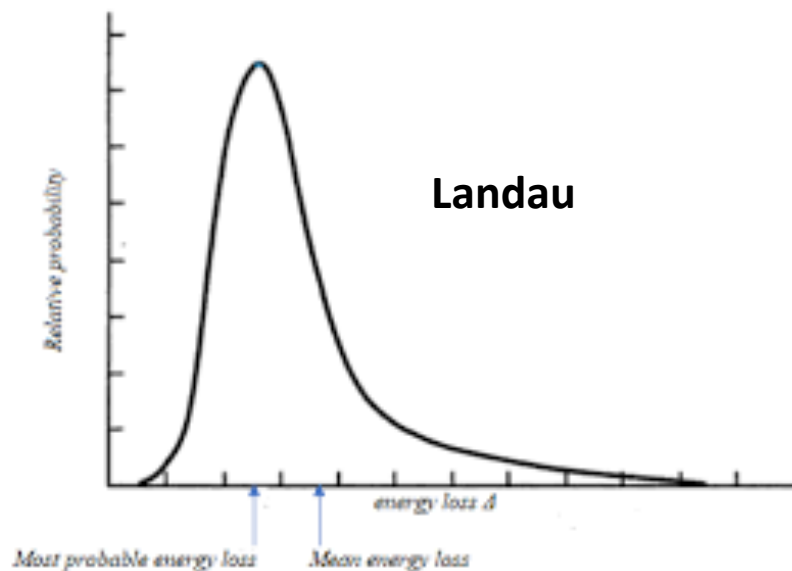
Element	$\sigma_z$
p	0.07
He	0.12
Li	0.14
Be	0.21
B	0.17
C	0.18
N	0.21
O	0.21

Template fit for protons and Helium

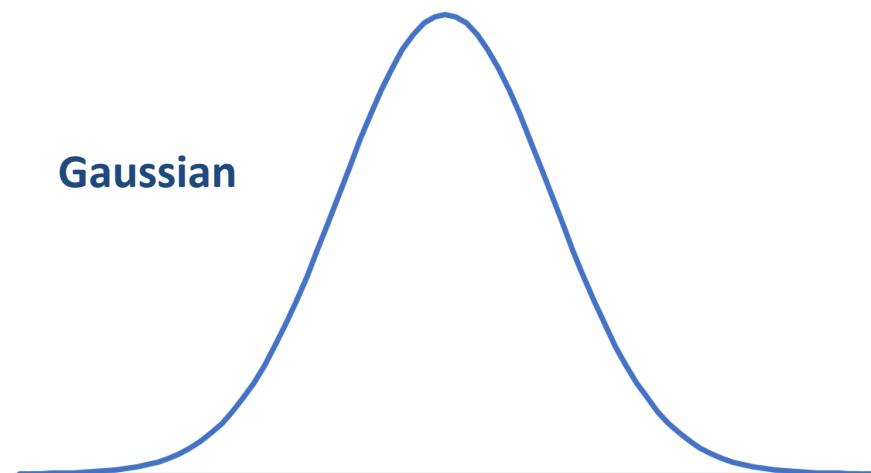




## Energy loss in a thin layer of matter

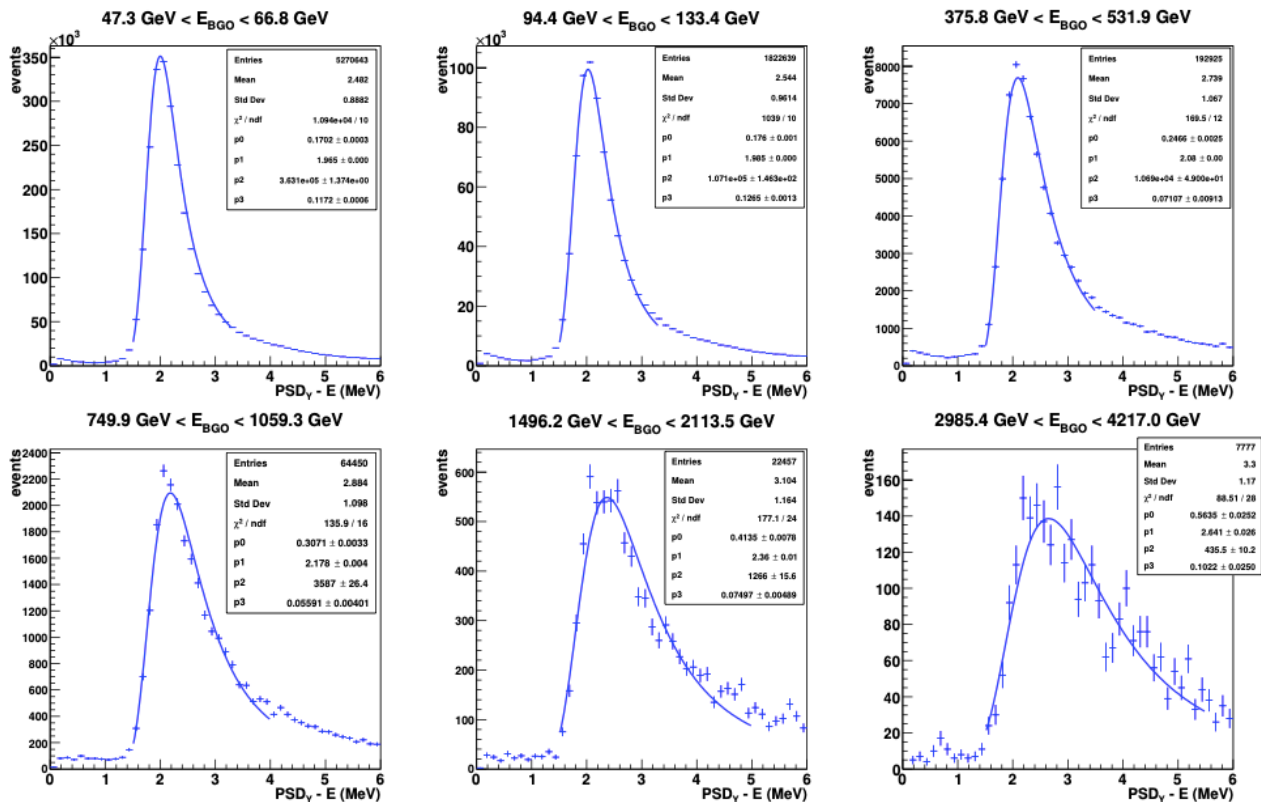


## Measurement uncertainty



# Langaus curve

# Charge selection (protons as an example)



Landau + Gauss (langaus)  
fit of the PSD p-signal  
for different  $E_{BGO}$  ranges

Fit parameters

MPV (Most Probable Value)

$\sigma_{Landau}$

$\sigma_{Gauss}$

must be used to select the Z signal in real  
and simulated samples

# Charge selection

$$MPV - 3\sqrt{\sigma_{Lan}^2 + \sigma_{Gaus}^2} < E_{PSD} < MPV + 4\sqrt{\sigma_{Lan}^2 + \sigma_{Gaus}^2}$$

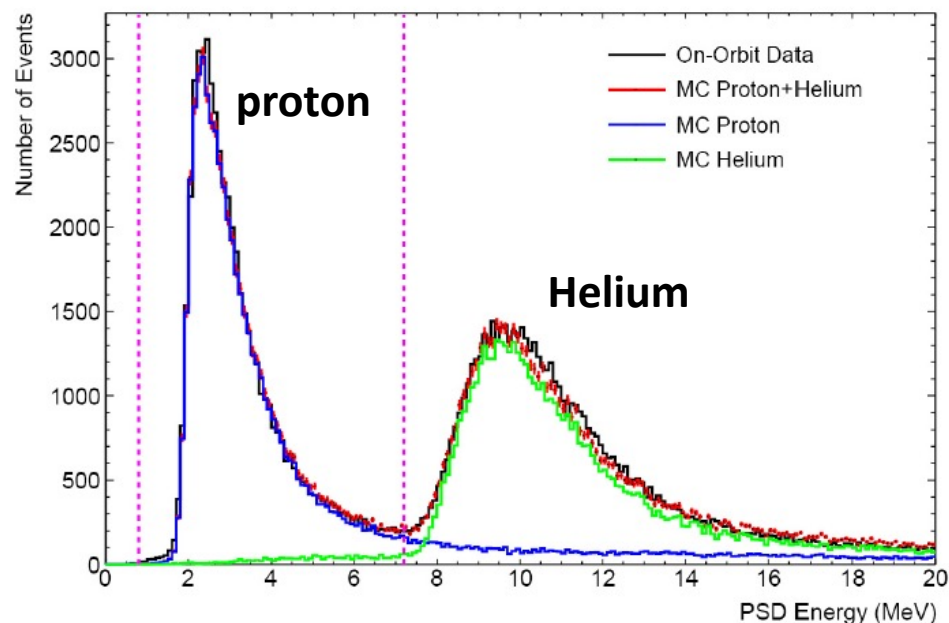
Same selection for real data and simulated ones

Parameters ( $MPV$ ,  $\sigma_{Landau}$ ,  $\sigma_{Gauss}$ ) are functions on  $E_{BGO}$

## Background estimate

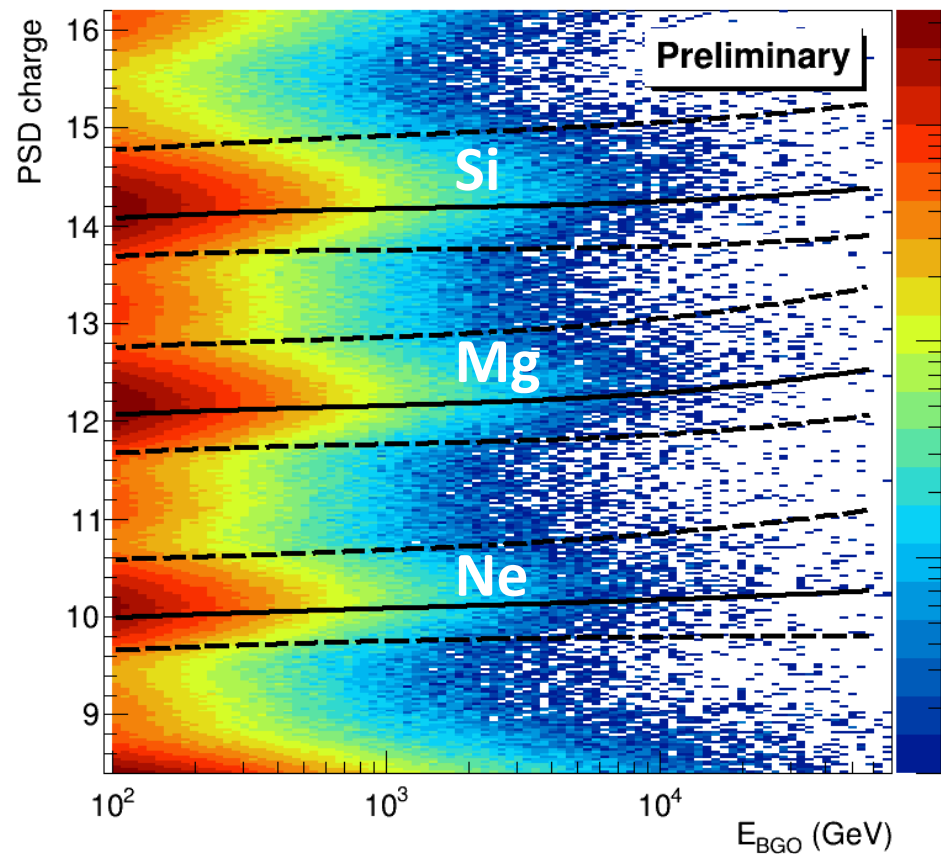
Template fit for protons and Helium

The background can be subtracted or taken into account as systematical error

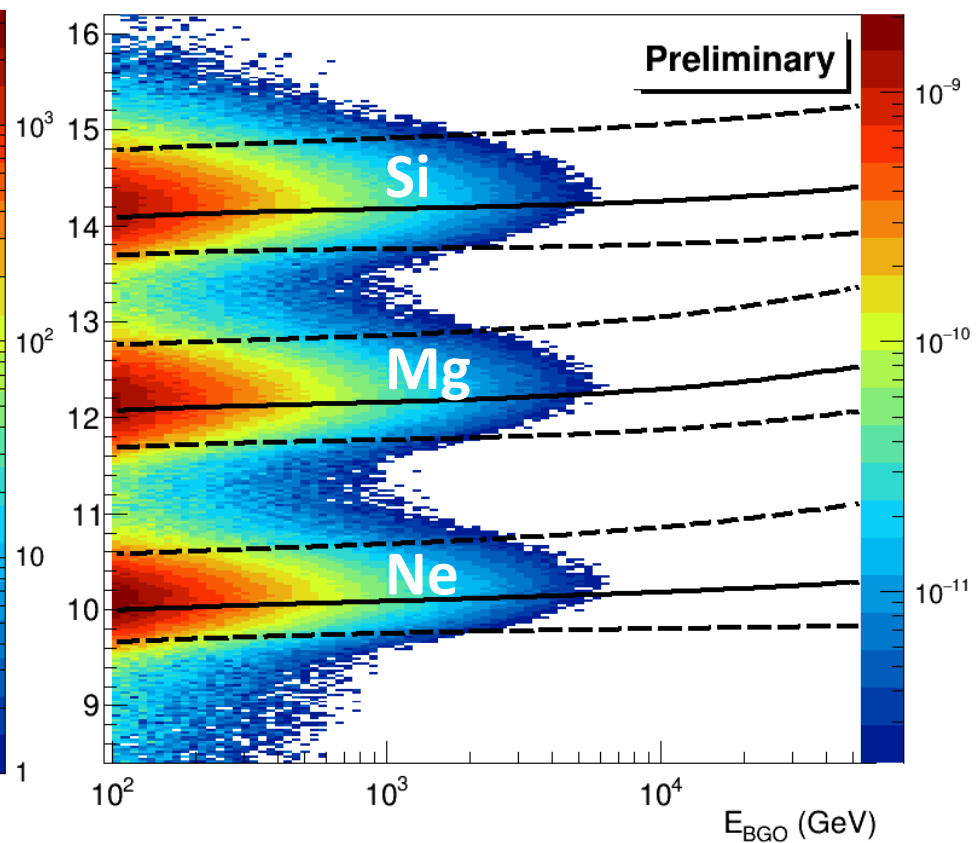


# Charge selection vs BGO energy

Flight data



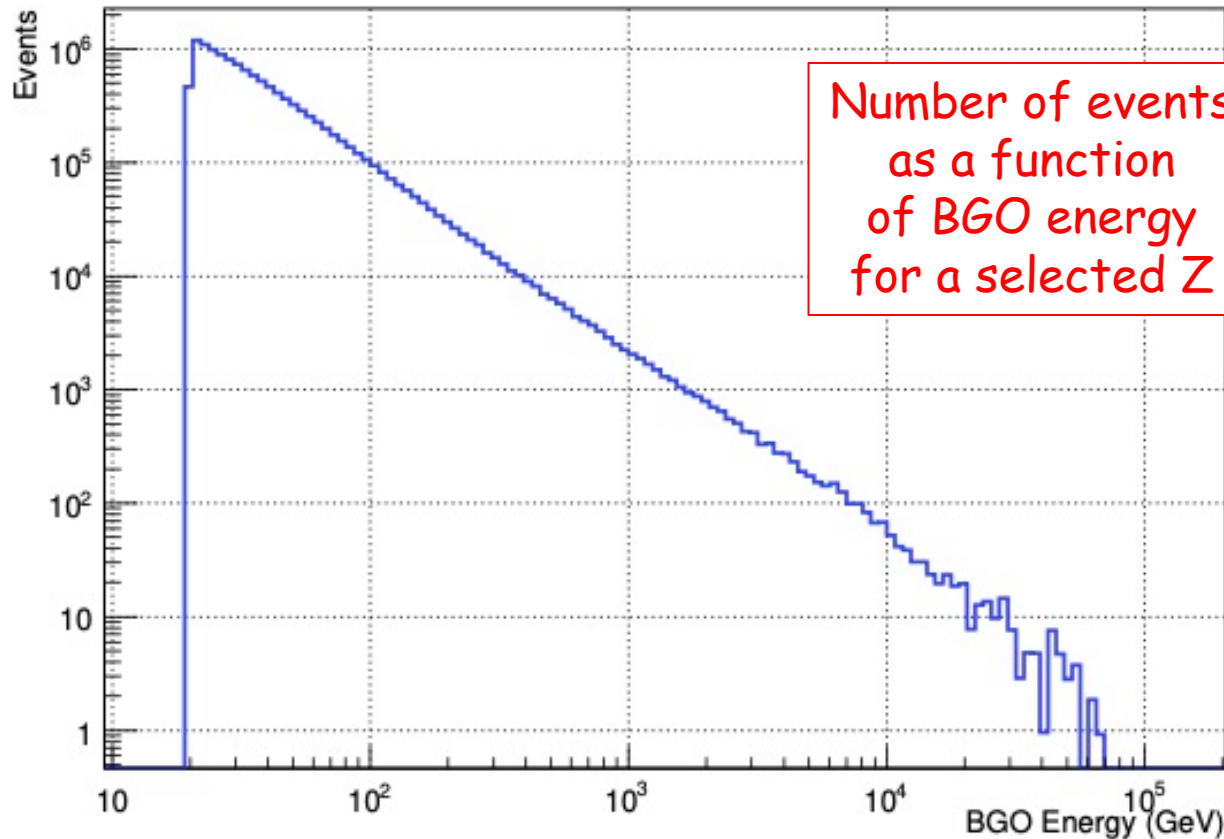
Simulation



In these plots: PSD energy  $\rightarrow$  PSD charge



# Unfolding procedure (I)

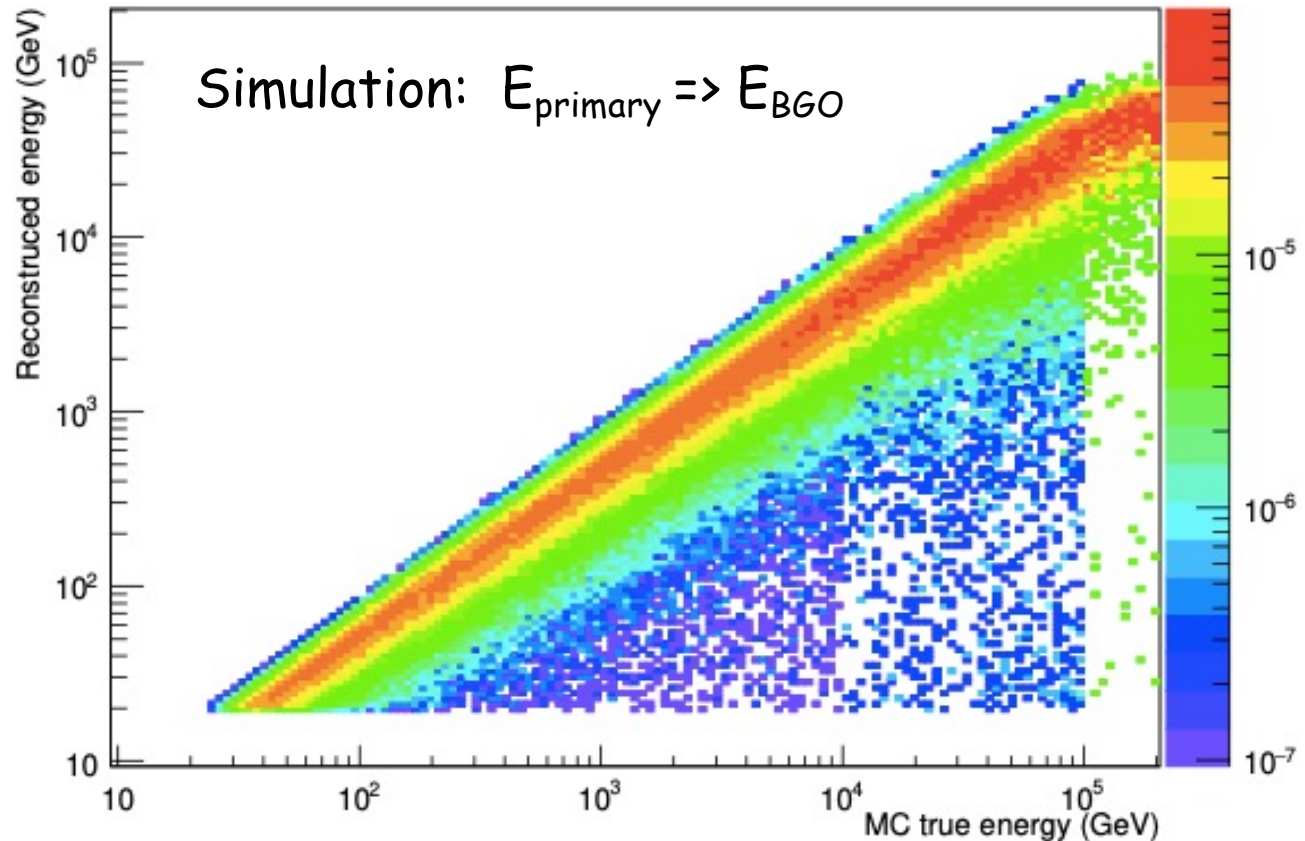


$$N(E_{\text{BGO}}) \neq N(E_{\text{RC}})$$

We need the number of events as a function of primary energy

# Unfolding procedure (II)

What energy in the calorimeter assuming the primary energy ?



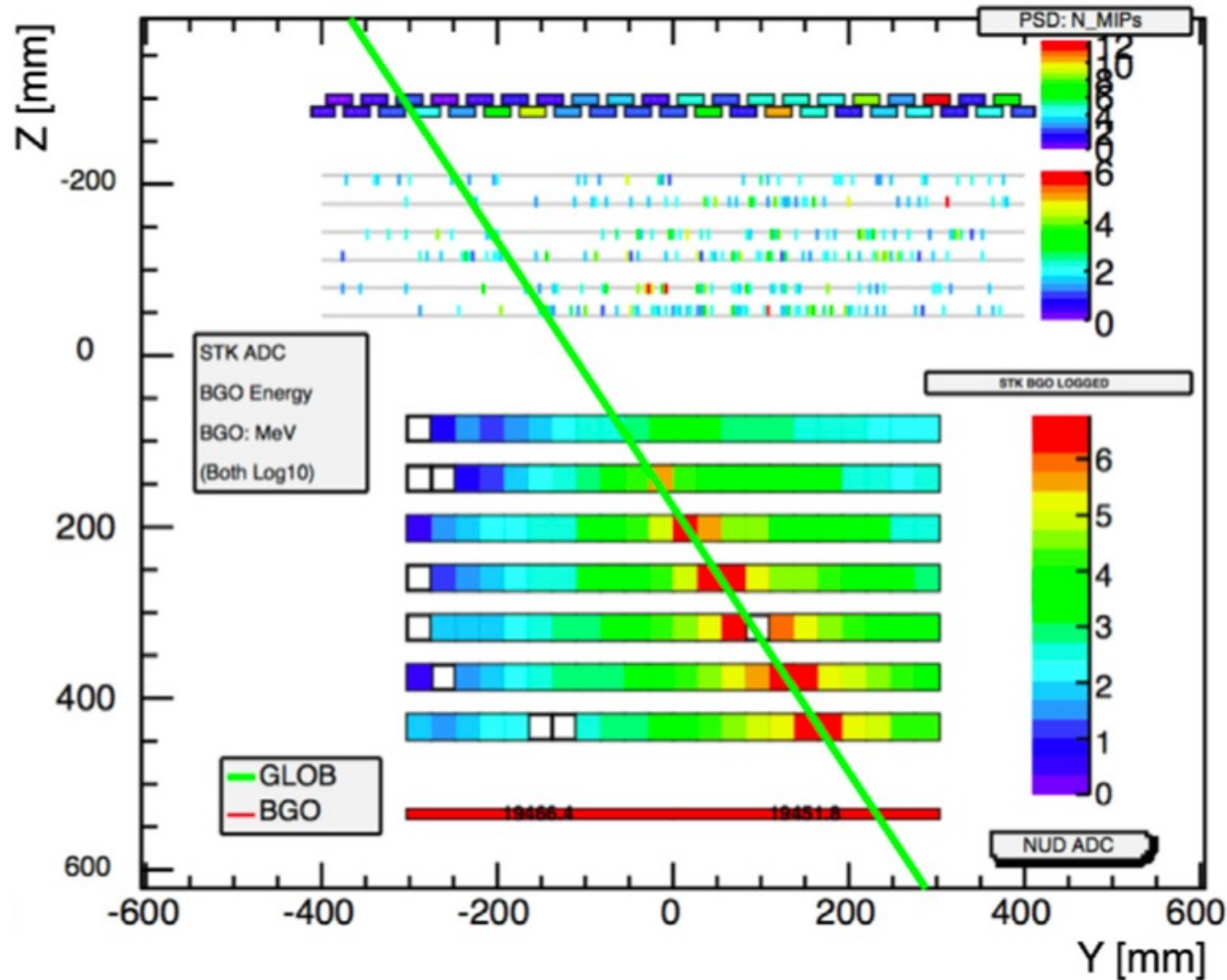
By means of the **unfolding** matrix

$$E_{\text{BGO}} \Rightarrow E_{\text{primary}}$$

(estimate on statistical basis)

# BGO saturation - white box close to red boxes

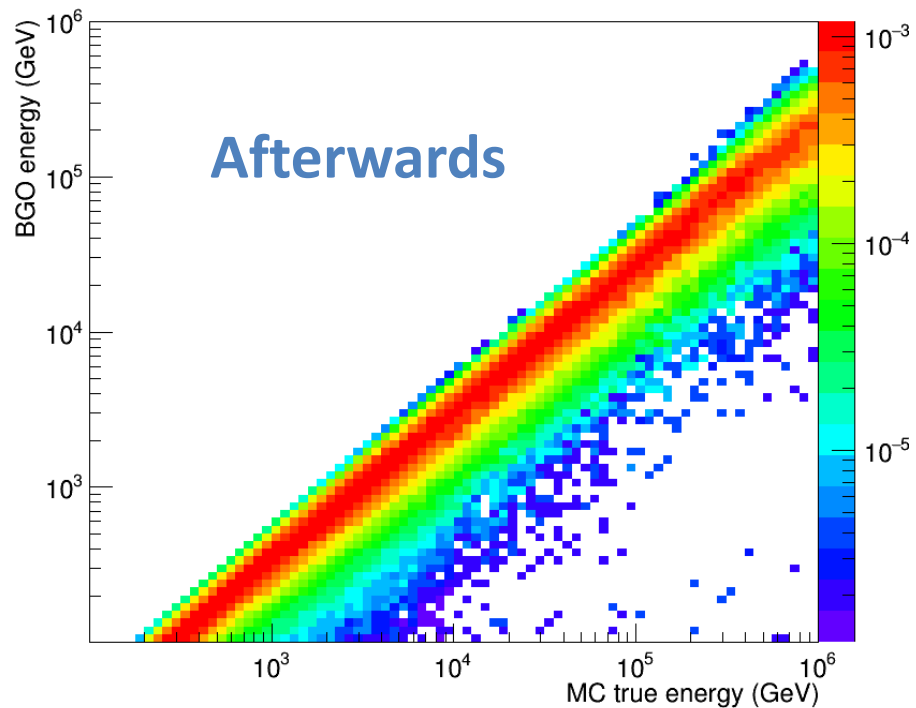
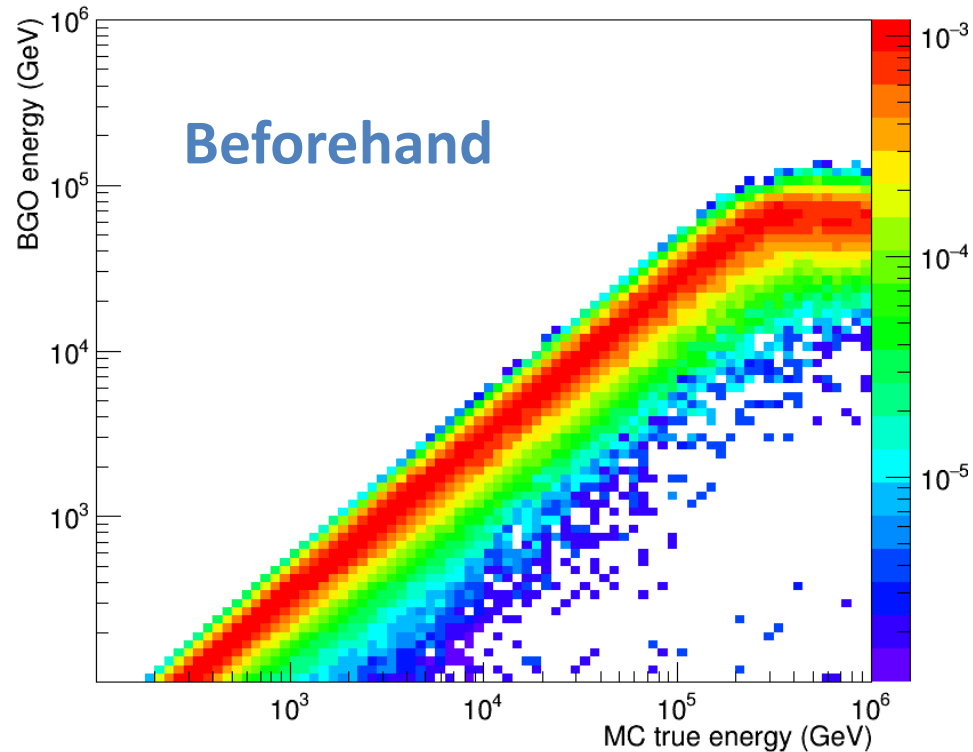
## Y-Z View



Correction  
by means of  
simulation

# Effect of the saturation correction on the unfolding matrix

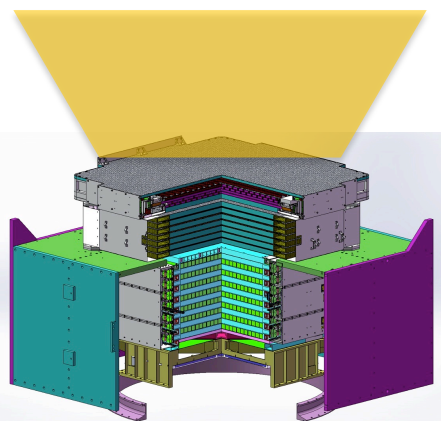
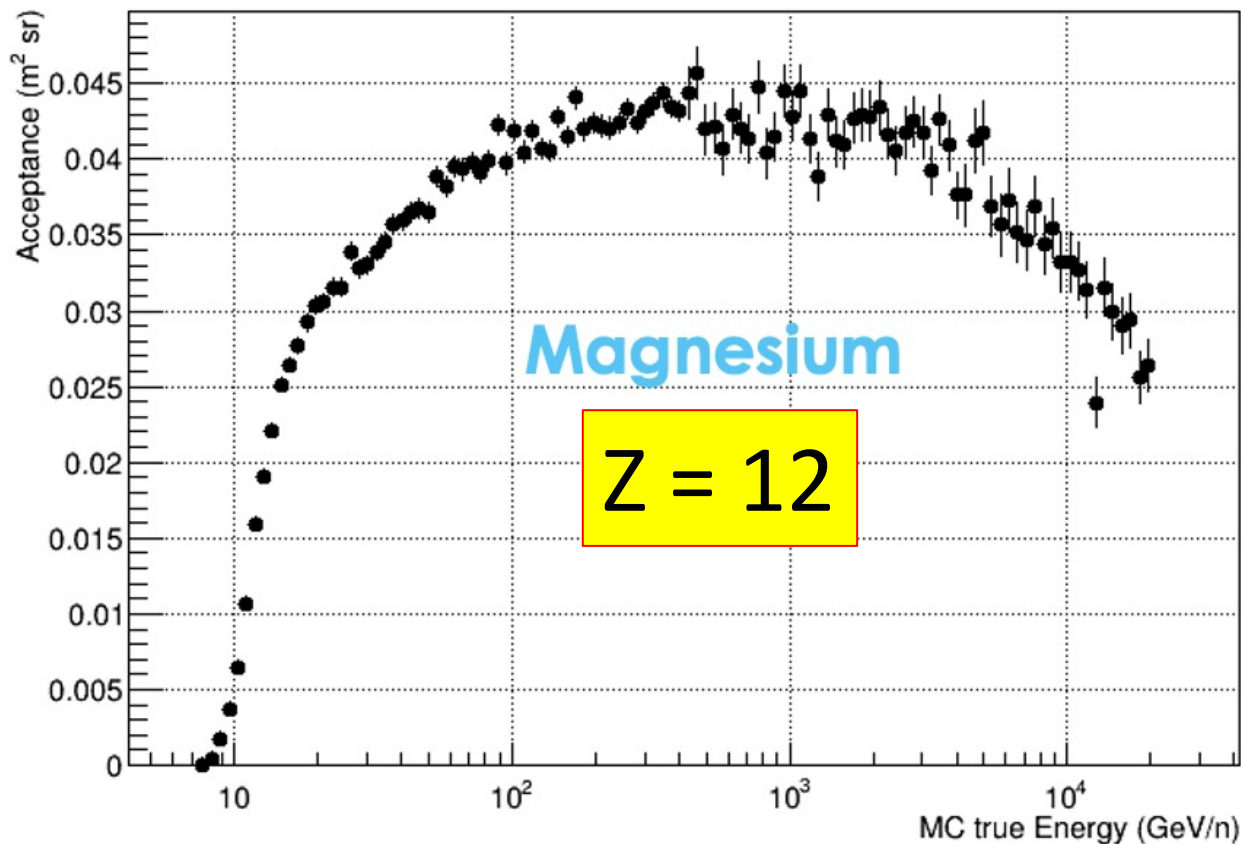
- Mg example -





# Estimate of acceptance

$$A_{GEOM} = \text{area} \times \text{solid angle} \text{ [m}^2 \text{ sr]}$$



$$A_{EFF} = \varepsilon A_{GEOM} = (N_{surv} / N_{simu}) A_{GEOM}$$

# Measurement of CR fluxes

## Protons (Z=1), Helium (Z=2) and so on

$$\Phi_Z(E) = \frac{N_Z(E)}{A_{\text{EFF}} \Delta E \Delta t}$$

$\Phi_Z(E)$  differential flux for the Z element

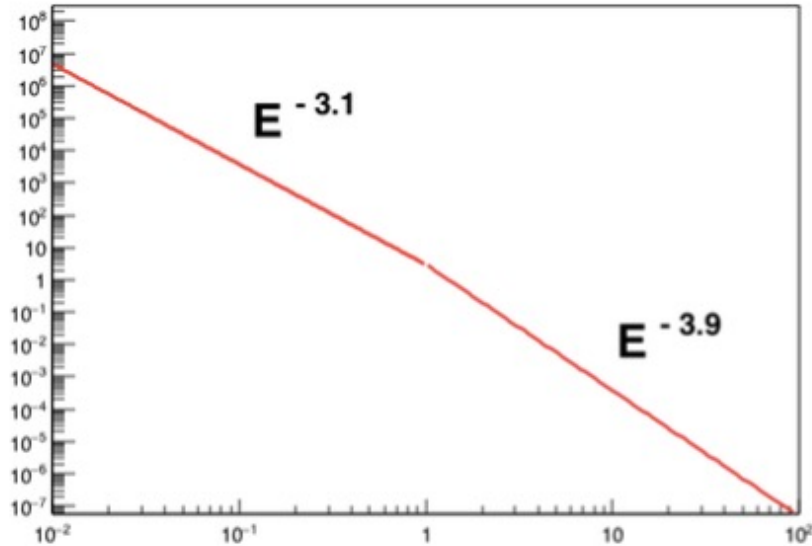
$N_Z(E)$  number of observed Z events in the energy range

$A_{\text{EFF}}$  effective acceptance, that is selection efficiency ( $\varepsilon$ ) times geometrical acceptance ( $A_{\text{GEOM}}$ )

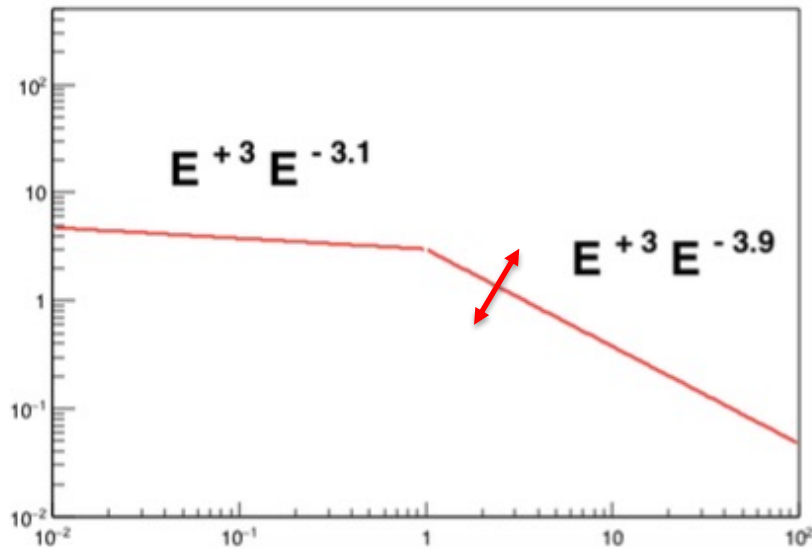
$\Delta E$  width of the energy bin

$\Delta t$  total live time

# Power law



The change of the spectral index is not so evident



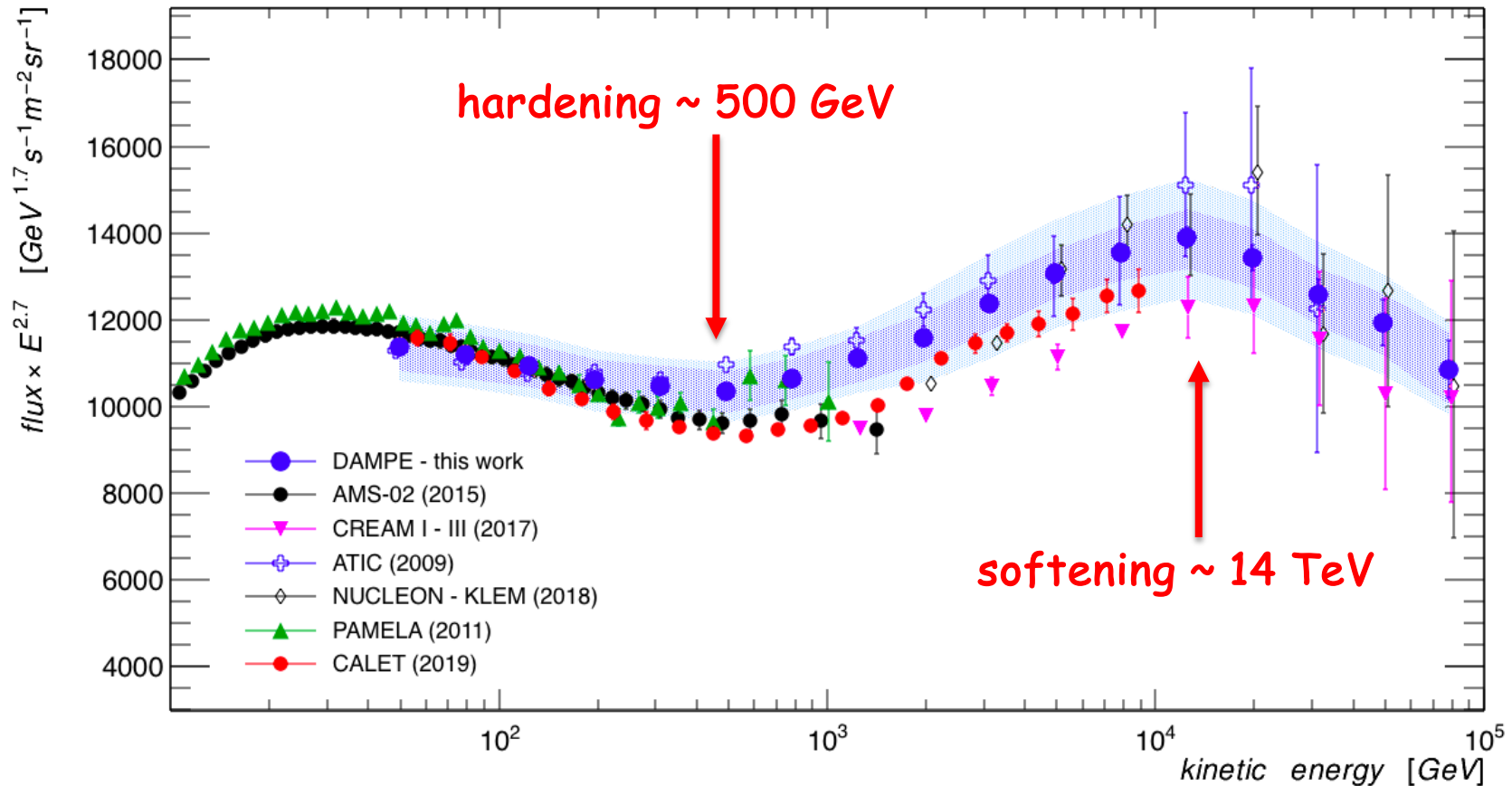
The change of the spectral index is made evident by the factor  $E^{+3}$

Energy-correction affects both the axes

# Proton spectrum (I)

30 months of data (Jan 2016 - Jun 2018)

Spectral hardening followed by softening  
(measured with unprecedented precision)



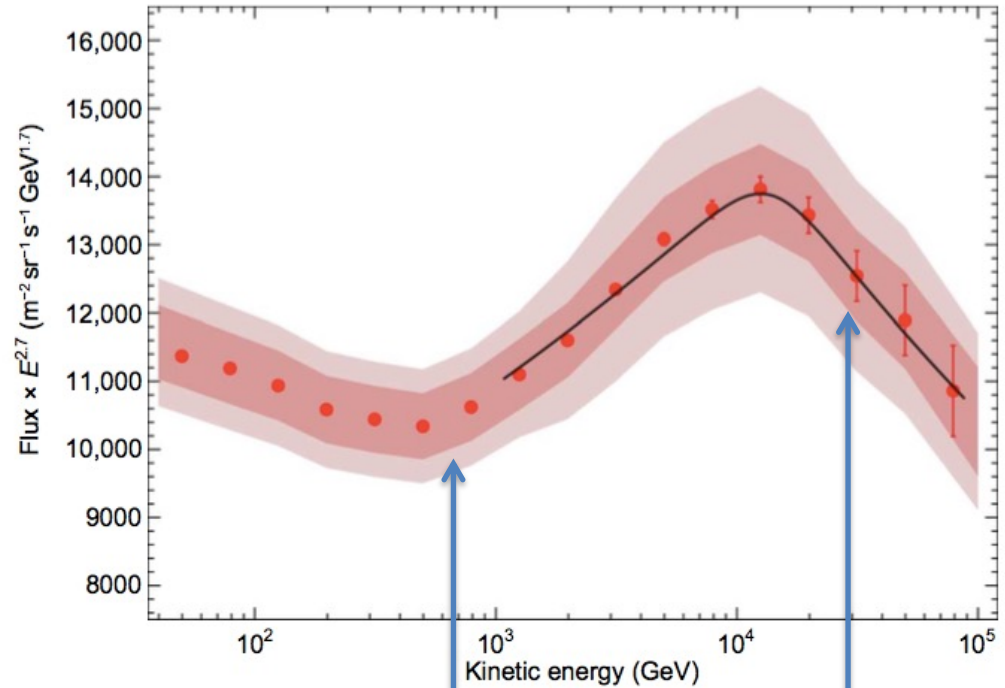
Sci. Adv. 2019; 5:eaax3793



# Proton spectrum (II)

Fit with a Smoothly Broken Power Law ( $s = 5$ )

$$\Phi(E) = \Phi_0 \left( \frac{E}{\text{TeV}} \right)^{-\gamma_1} \left[ 1 + \left( \frac{E}{E_B} \right)^s \right]^{-(\gamma_2 - \gamma_1)/s}$$



**100 GeV – 6.3 TeV**

**Hardening**

$E_B = 480 \text{ GeV}$

$\gamma_1 = 2.772$

$\gamma_2 = 2.599$

**1 – 100 TeV**

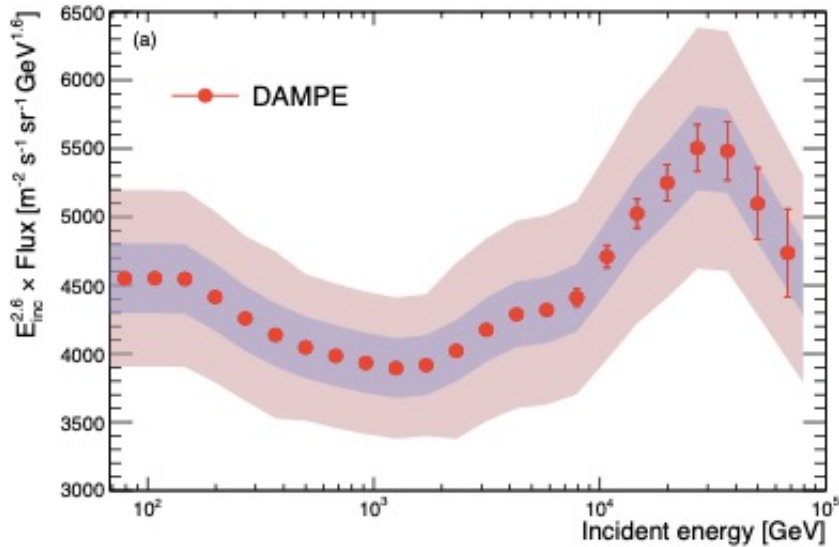
**Softening**

$E_B = 13.6 \text{ TeV}$

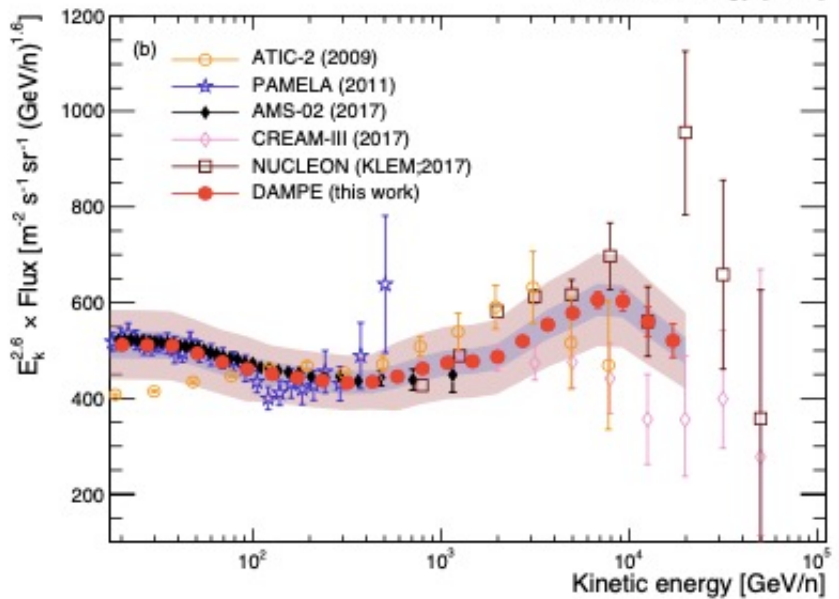
$\gamma_1 = 2.60$

$\gamma_2 = 2.85$

# Helium spectrum



*Physical Review Letters  
126 (2021) 201102*



Hardening at  $\sim 1.2$  TeV

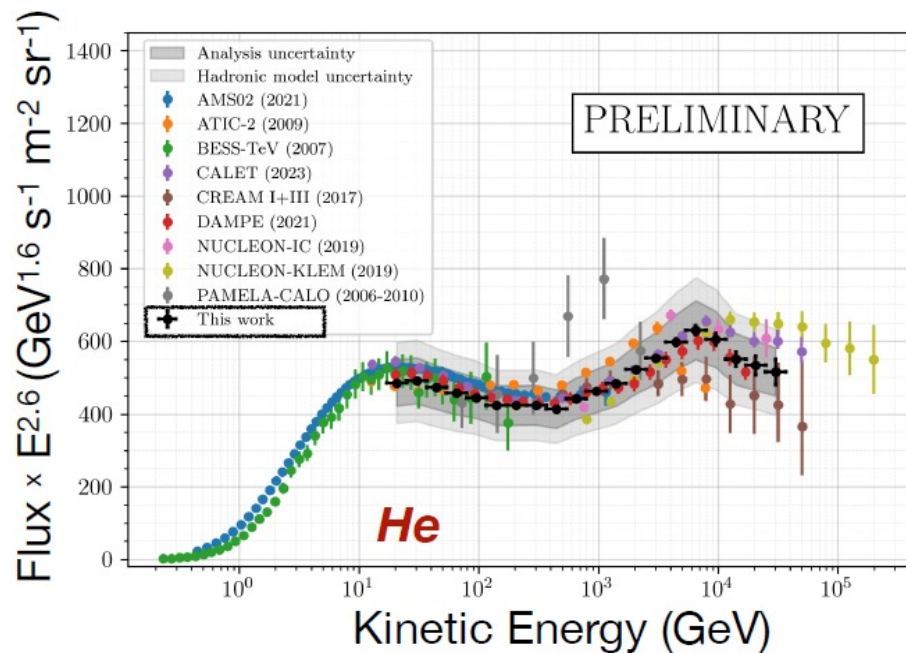
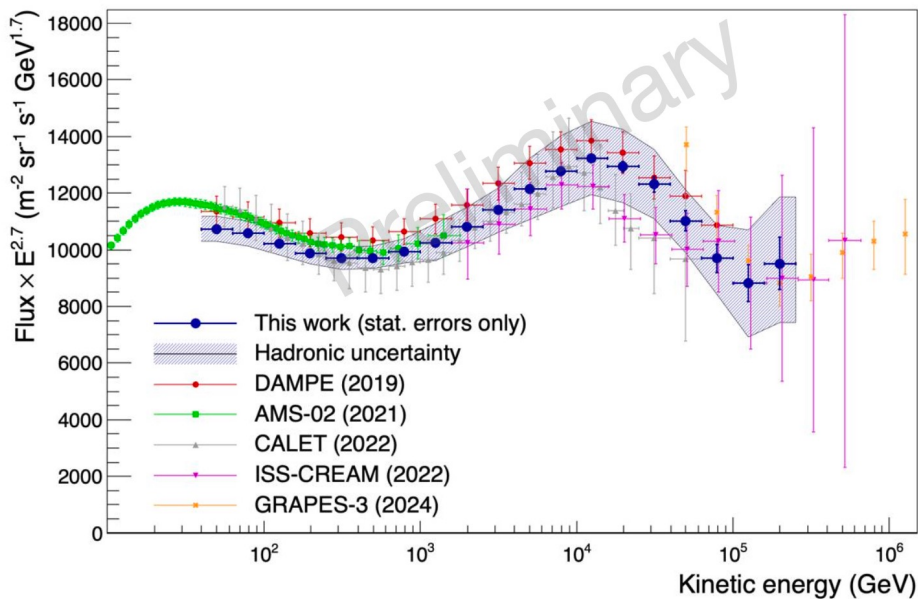
First clear evidence of  
a softening at  $\sim 34$  TeV

Z-dependent hardening  
and softening energies  
are suggested

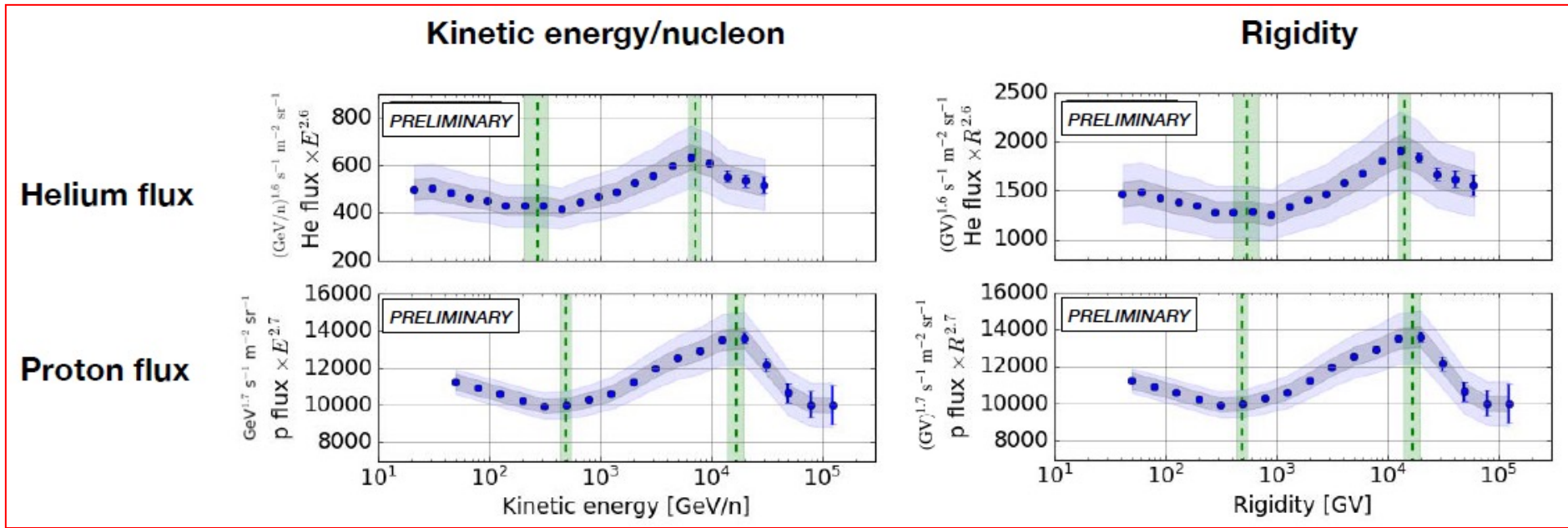
# Proton and He spectra: updates



## - Neural Network tracking -



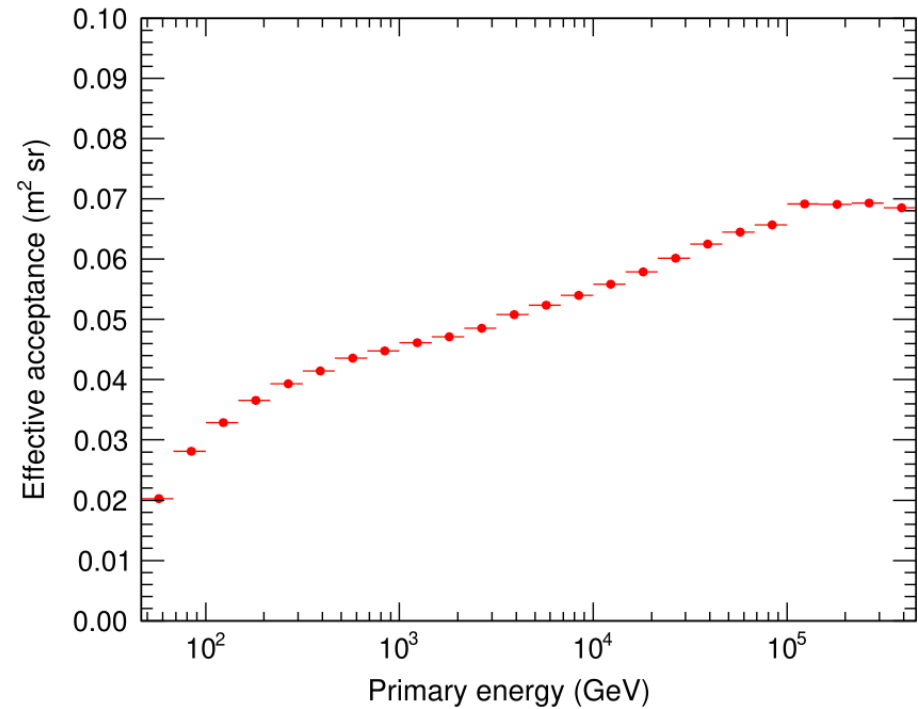
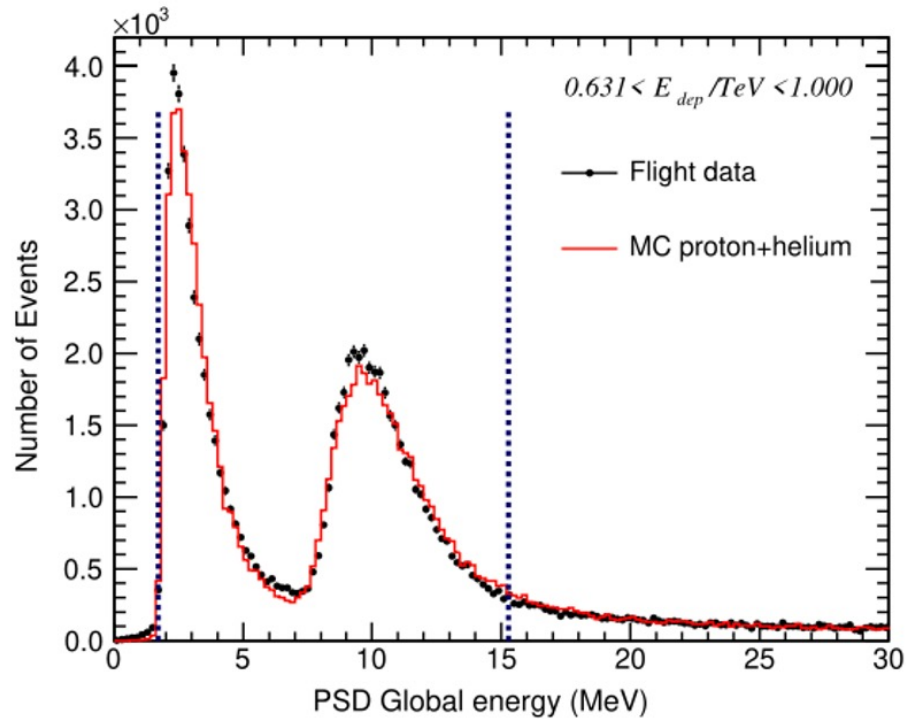
# Hardening/softening energies



A rigidity dependence of hardening (500 GV) and softening (15 TV) is favoured by data

# Protons + Helium - Bridge between space and ground measurements

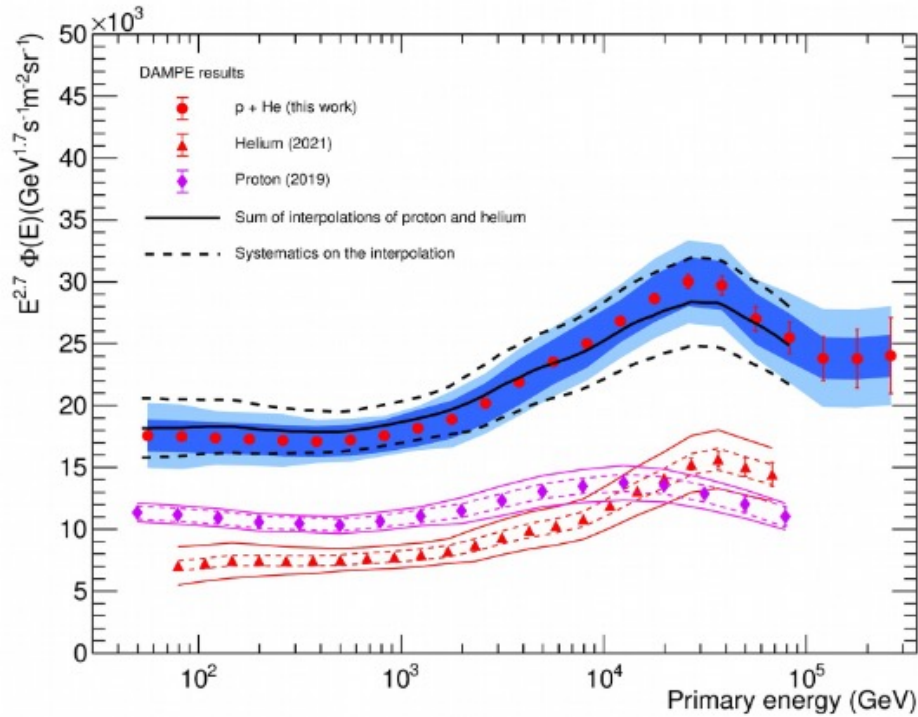
PRD 109 (2024) L121101



- Higher effective acceptance
- Lower systematics
- Measurements at higher energies

# Protons + Helium - Bridge between space and ground measurements

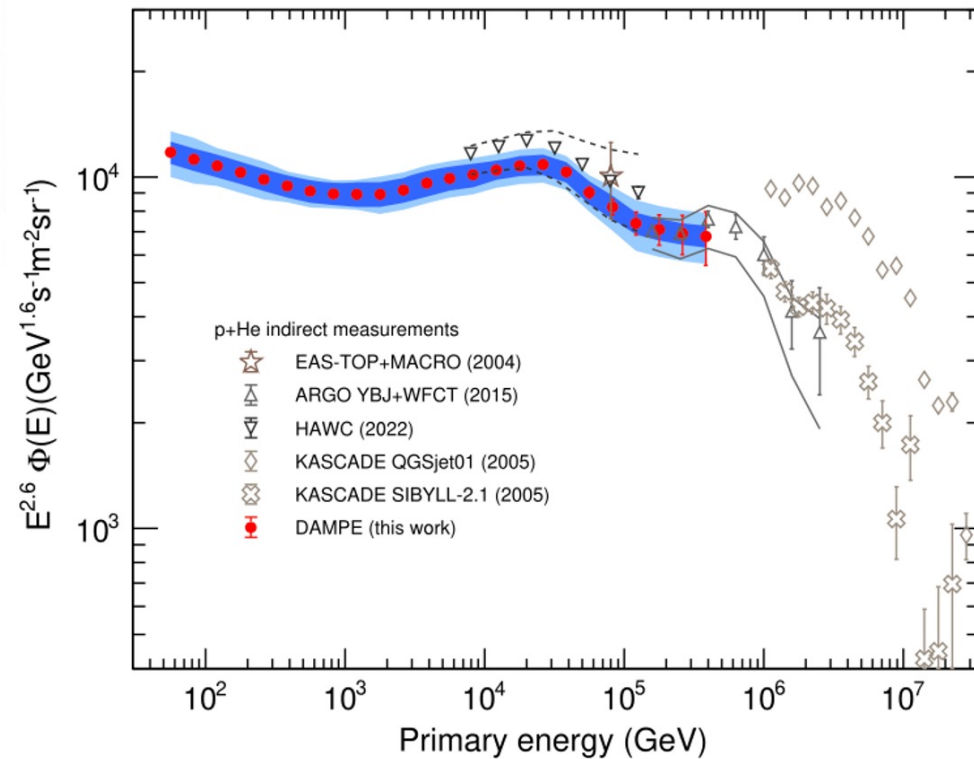
PRD 109 (2024) L121101



Softening confirmed at 15 TV

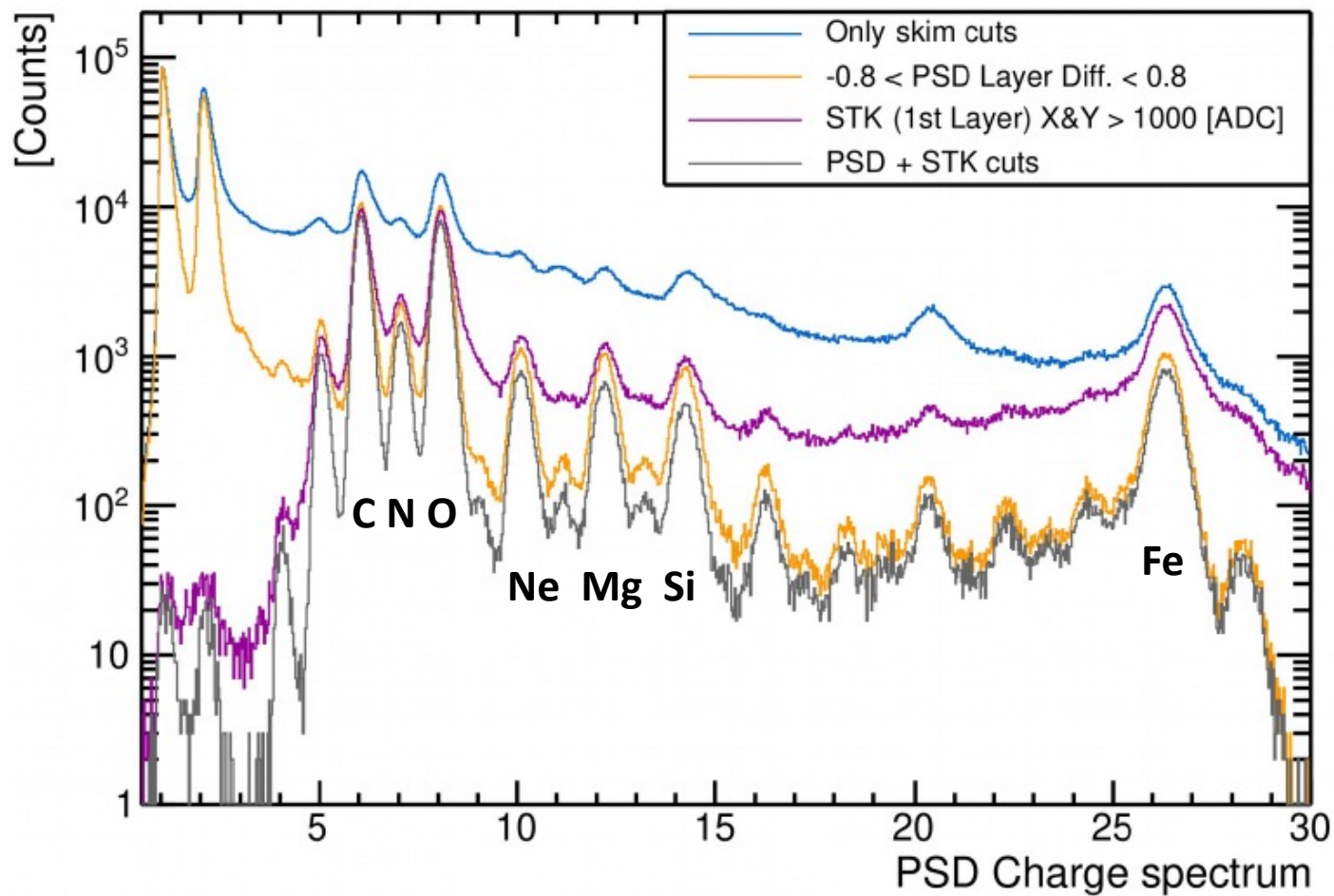
Hint of a following hardening

Bridge with ground experiment

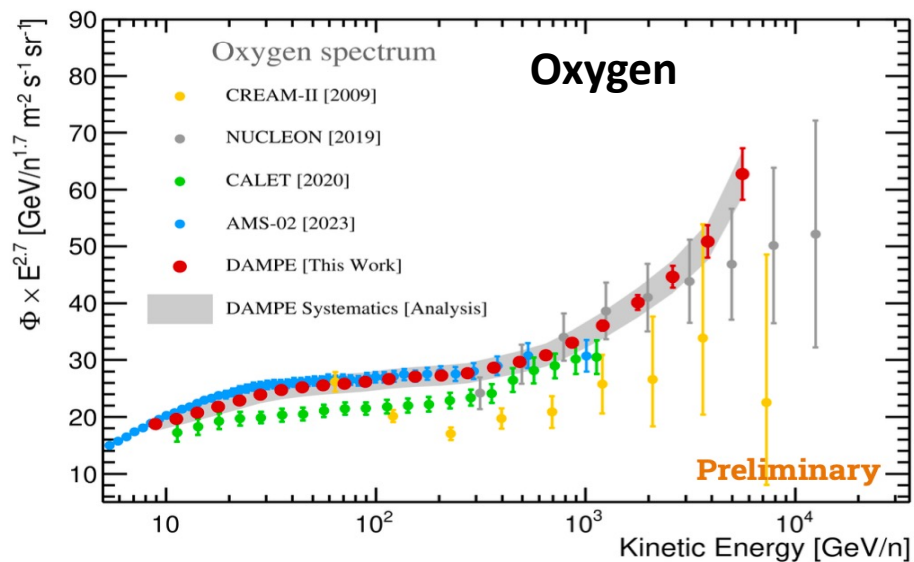
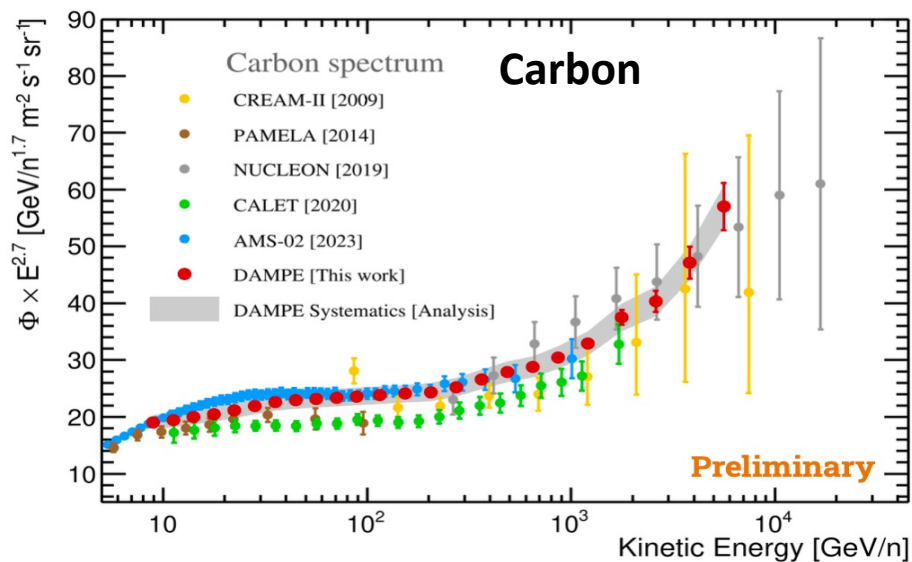
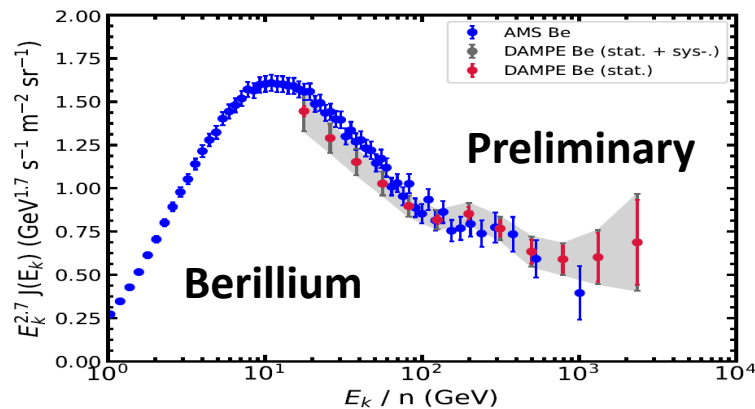
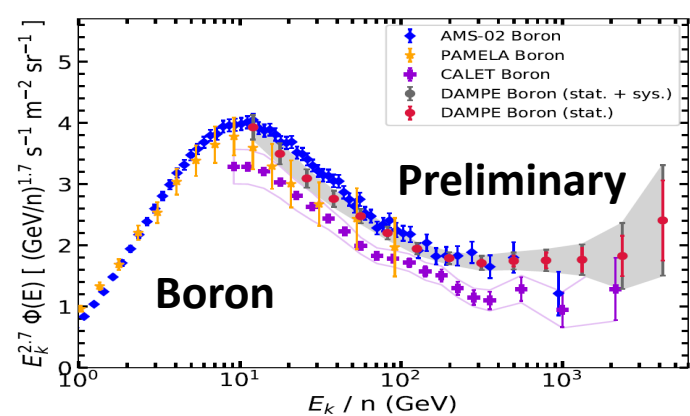
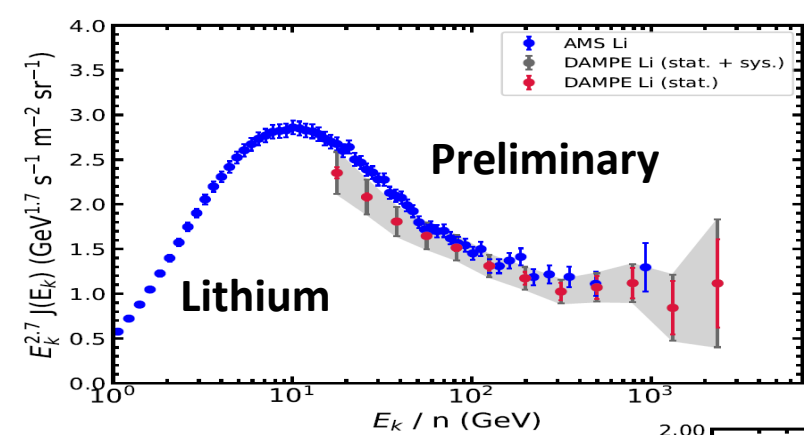


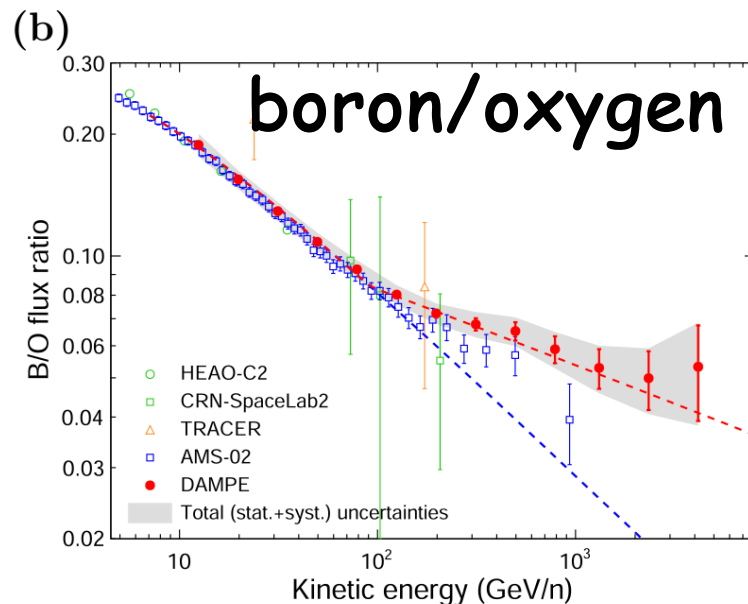
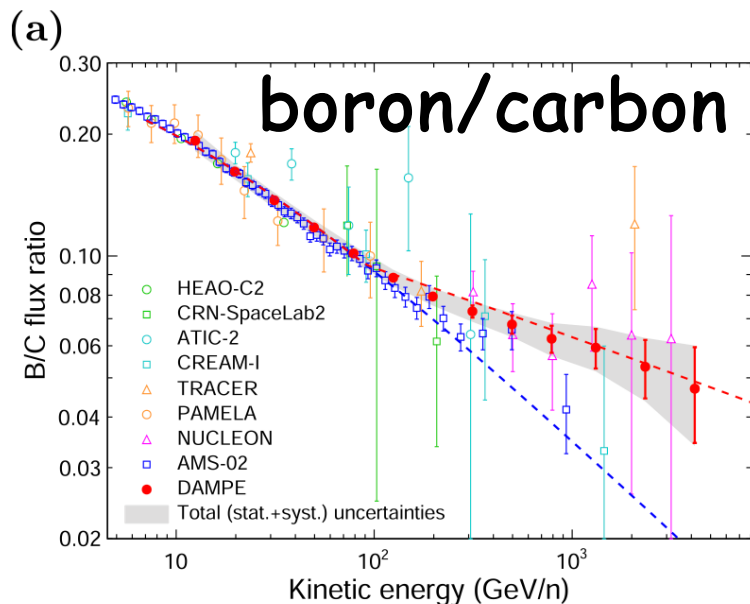


# Many independent analyses from Li up to Fe

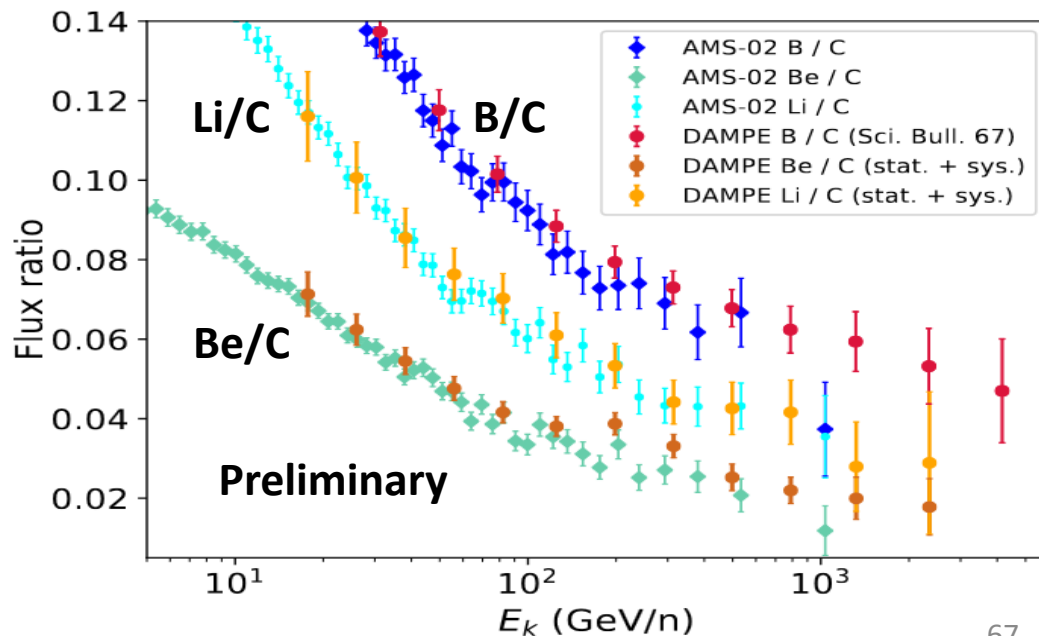


Just an example  
Elemental spectrum  
optimized for the  
CNO group

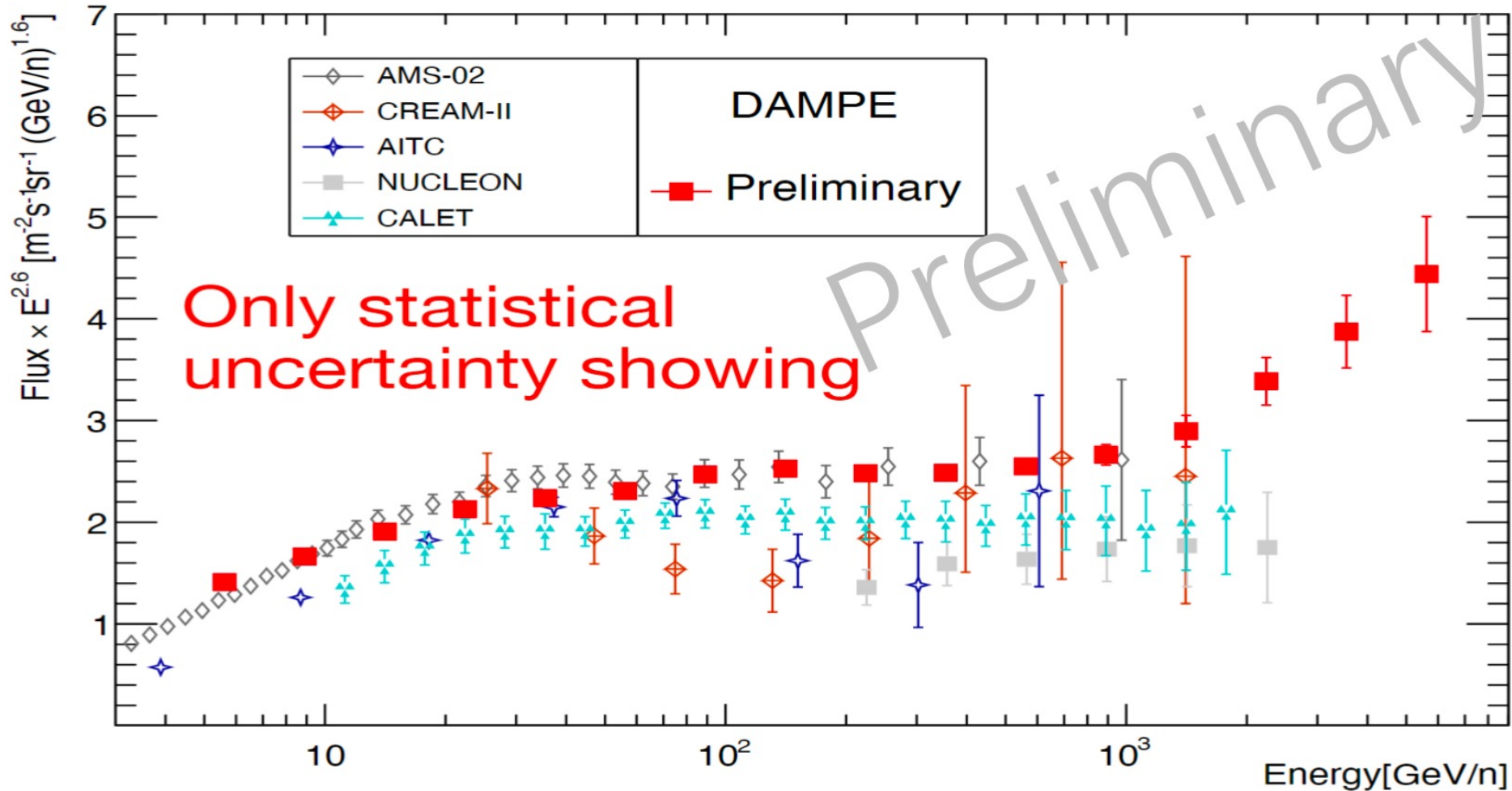




Measurements devoted to investigate the CR propagation in the galaxy



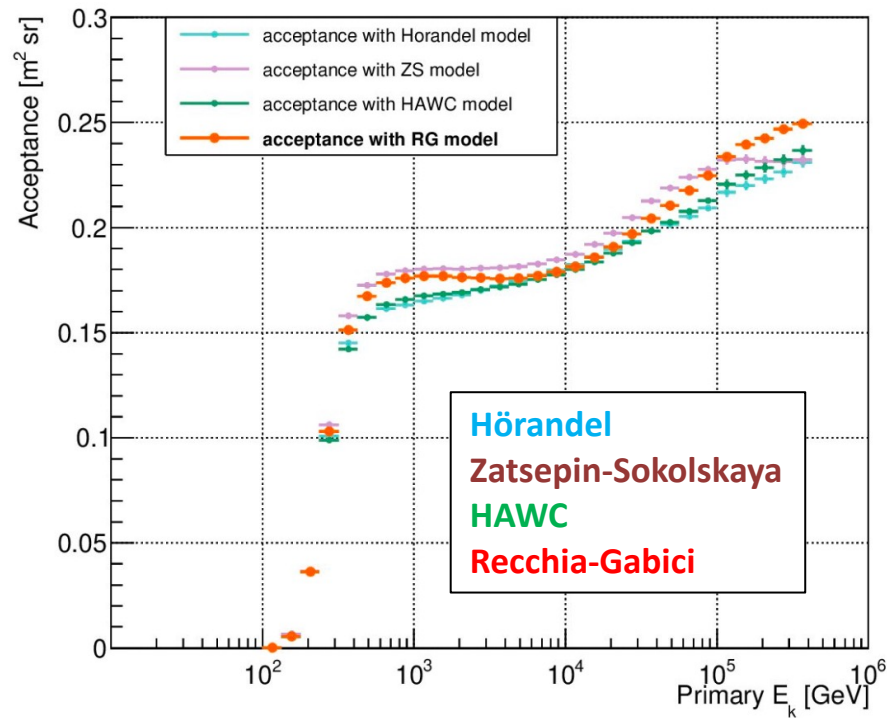
# Iron spectrum



# All-particle spectrum

## Weighted mean acceptance

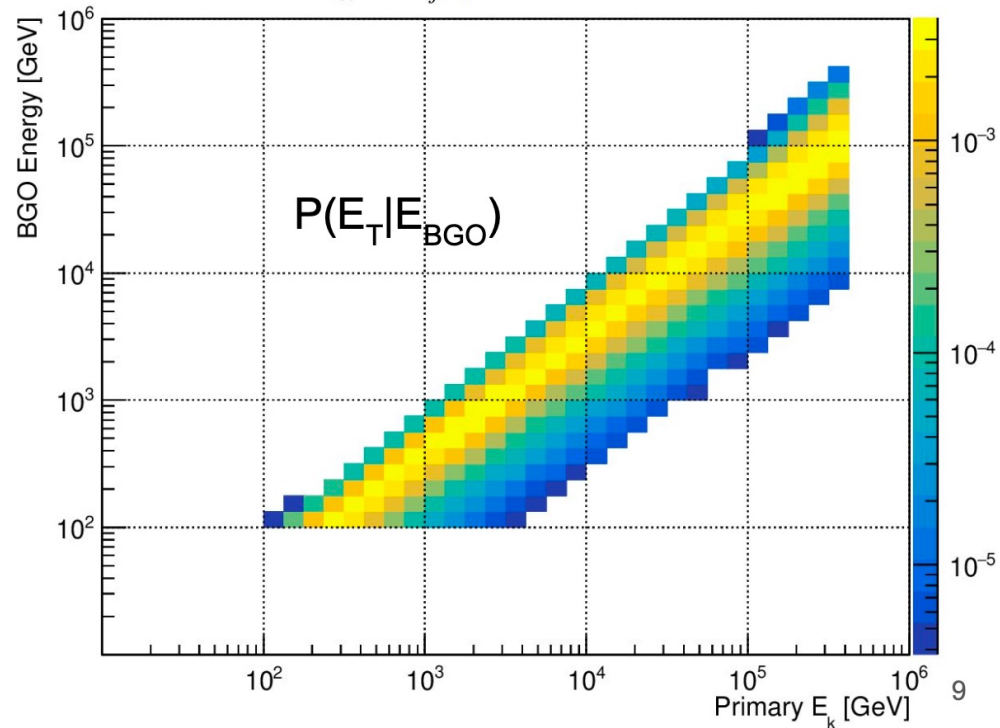
$$A_i = \sum_{el} w^{el} G_{gen}^{el} \frac{N_{sel}^{el}(E_T^i)}{N_{gen}^{el}(E_T^i)}$$



## Weighted mean response matrix

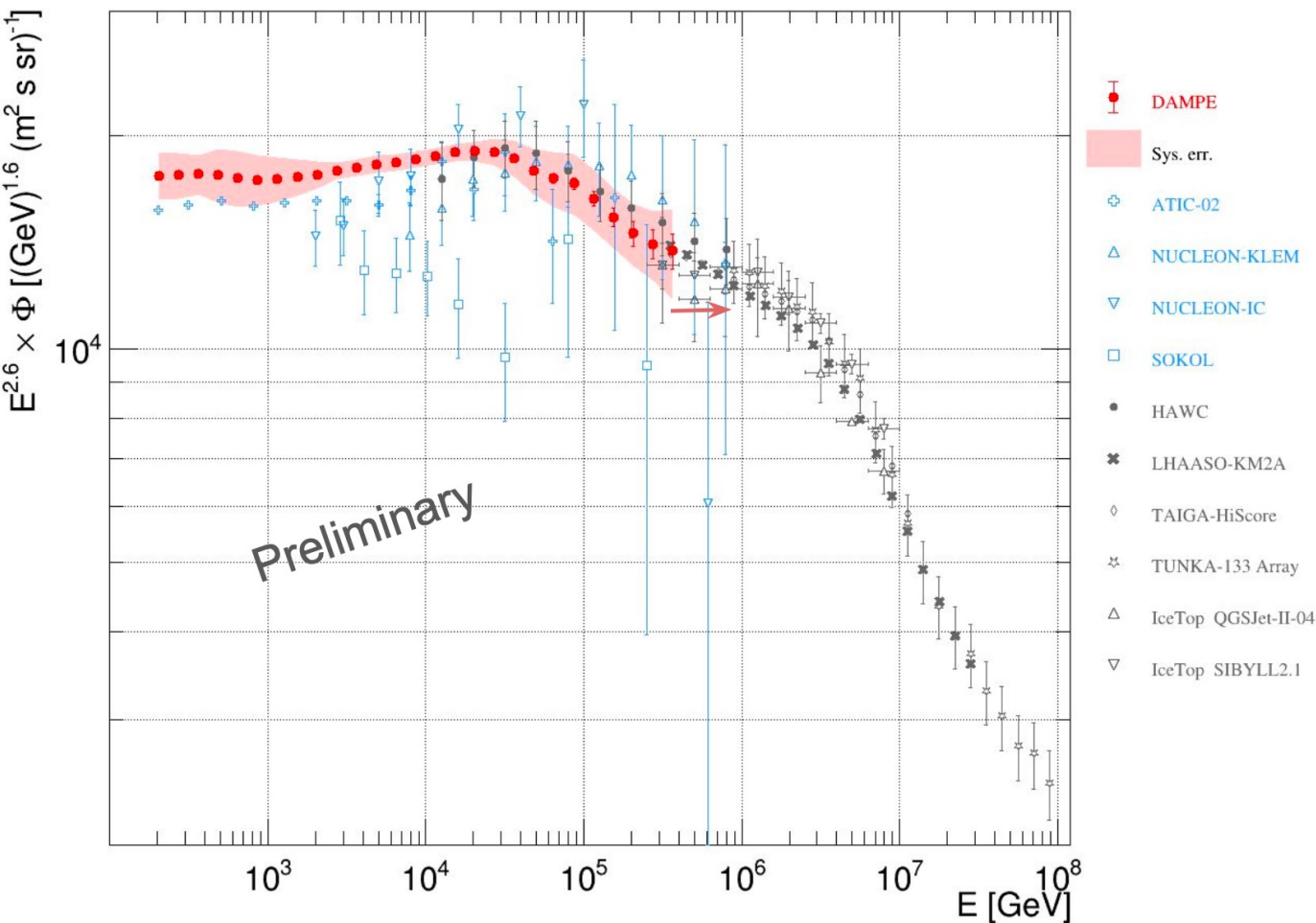
for the unfolding: iterative Bayesian procedure is adopted to reconstruct the primary energy of the events

$$N_i = \sum_{el} w^{el} \sum_{j=1}^n P^{el}(E_T^i | E_{BGO}^j) N^{el}(E_{BGO}^j)$$





# All-particle spectrum



- DAMPE
- Sys. err.
- ◆ ATIC-02
- ▲ NUCLEON-KLEM
- ▼ NUCLEON-IC
- SOKOL
- HAWC
- ✱ LHAASO-KM2A
- ◇ TAIGA-HiScore
- ☆ TUNKA-133 Array
- △ IceTop QGSJet-II-04
- ▽ IceTop SIBYLL2.1

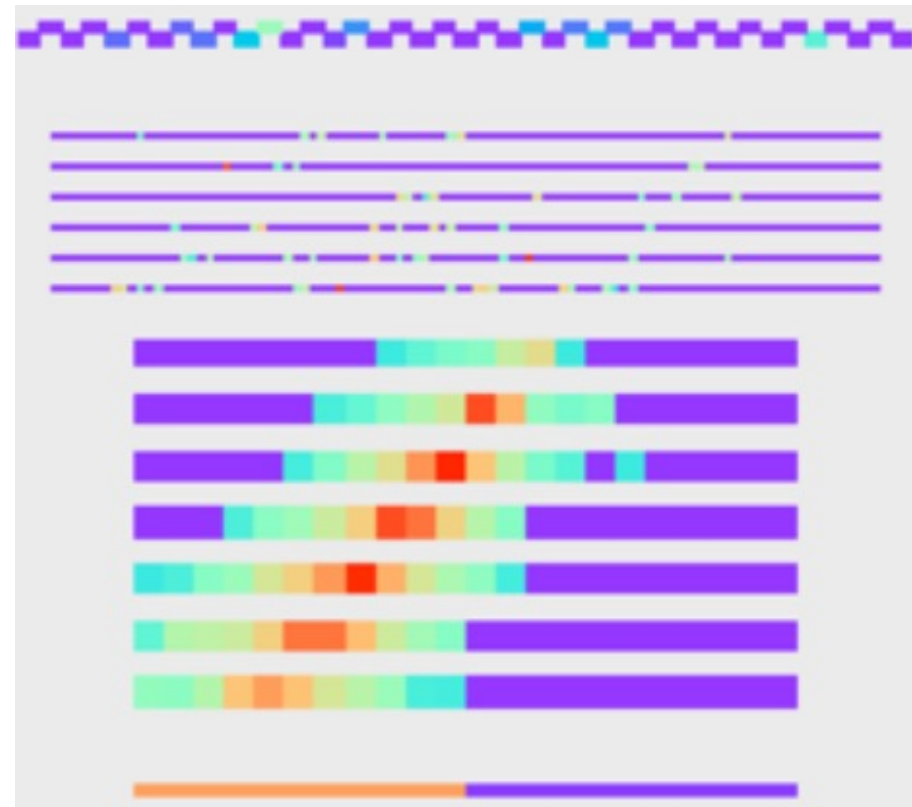
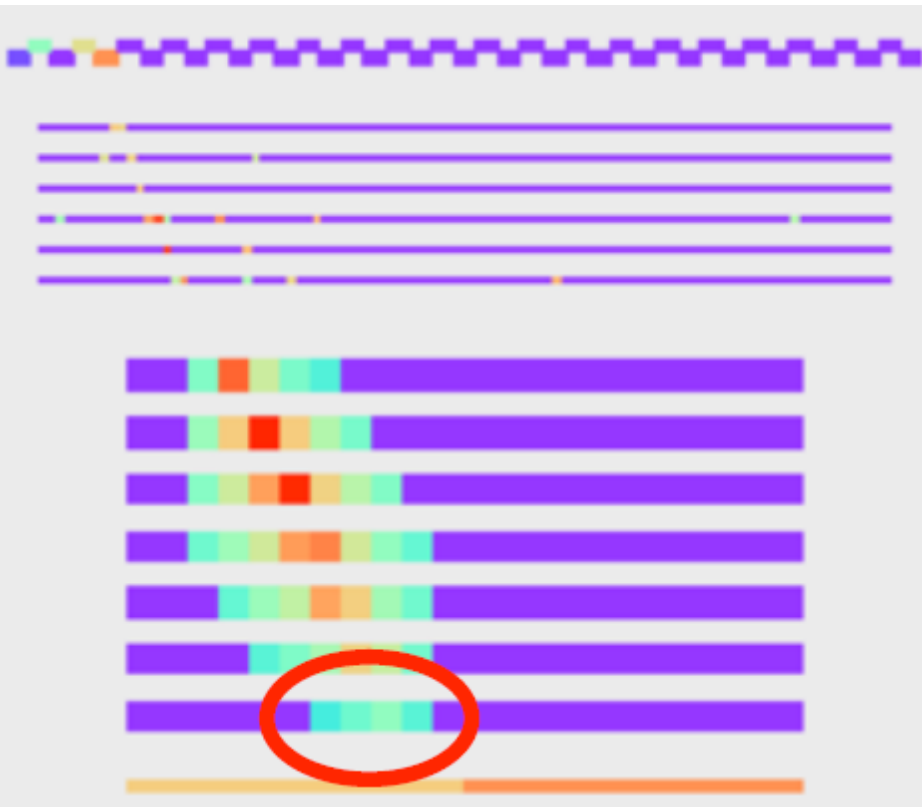
- Differential flux:
 
$$\Phi_i = \frac{N_i}{\Delta T \times A_i \times \Delta E_i}$$
- Evaluated systematics
  - Unfolding
  - Composition model
- Contribution from hadronic model is under evaluation
- In agreement with indirect experiments results
- Structure at tens TeV (convolution of the softening of different nuclei?)
- Work in progress to extend the measurement up to 0.7/0.8 PeV



# Electron+positron identification (I)

Selected events with  $Z_{\text{PSD}} = 1$

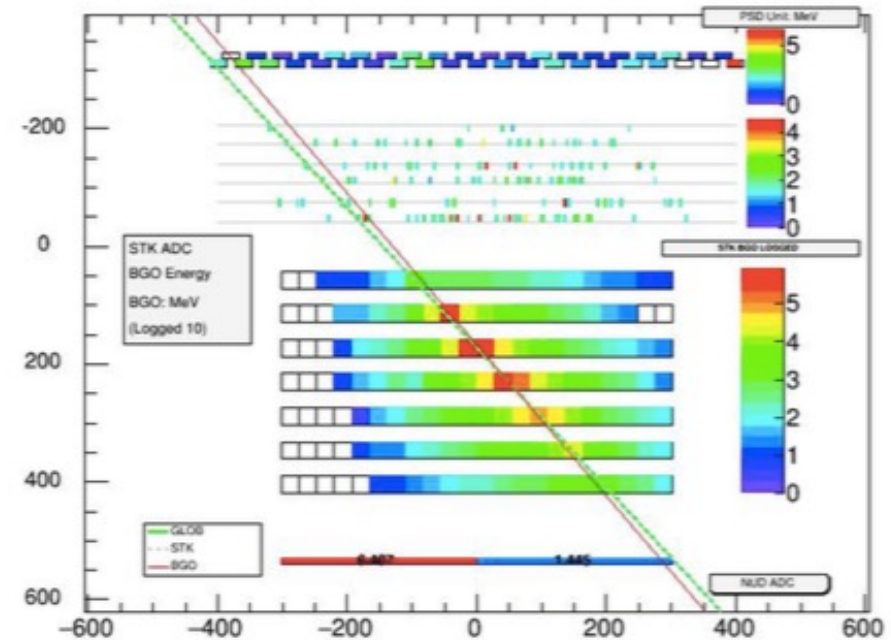
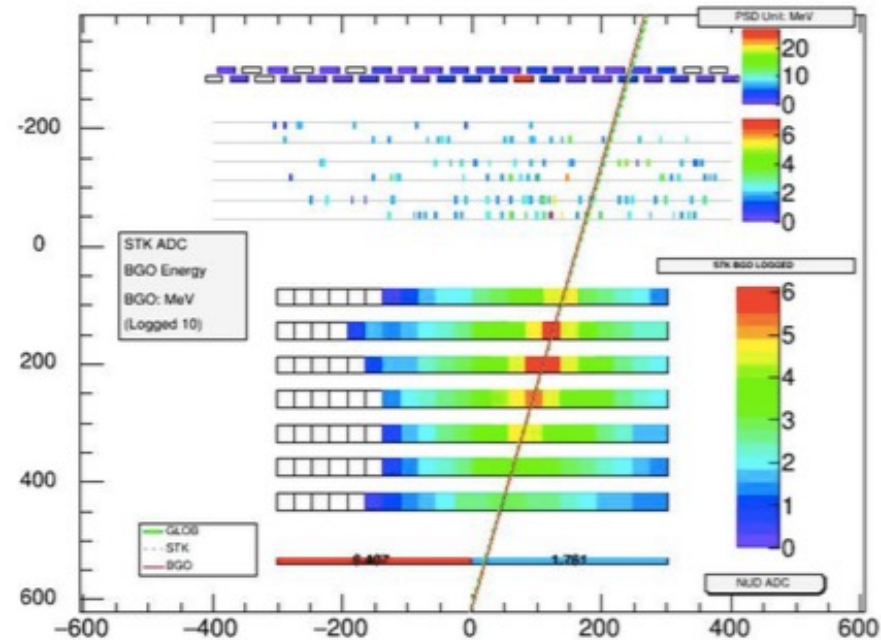
Exploiting the imaging CALO-features



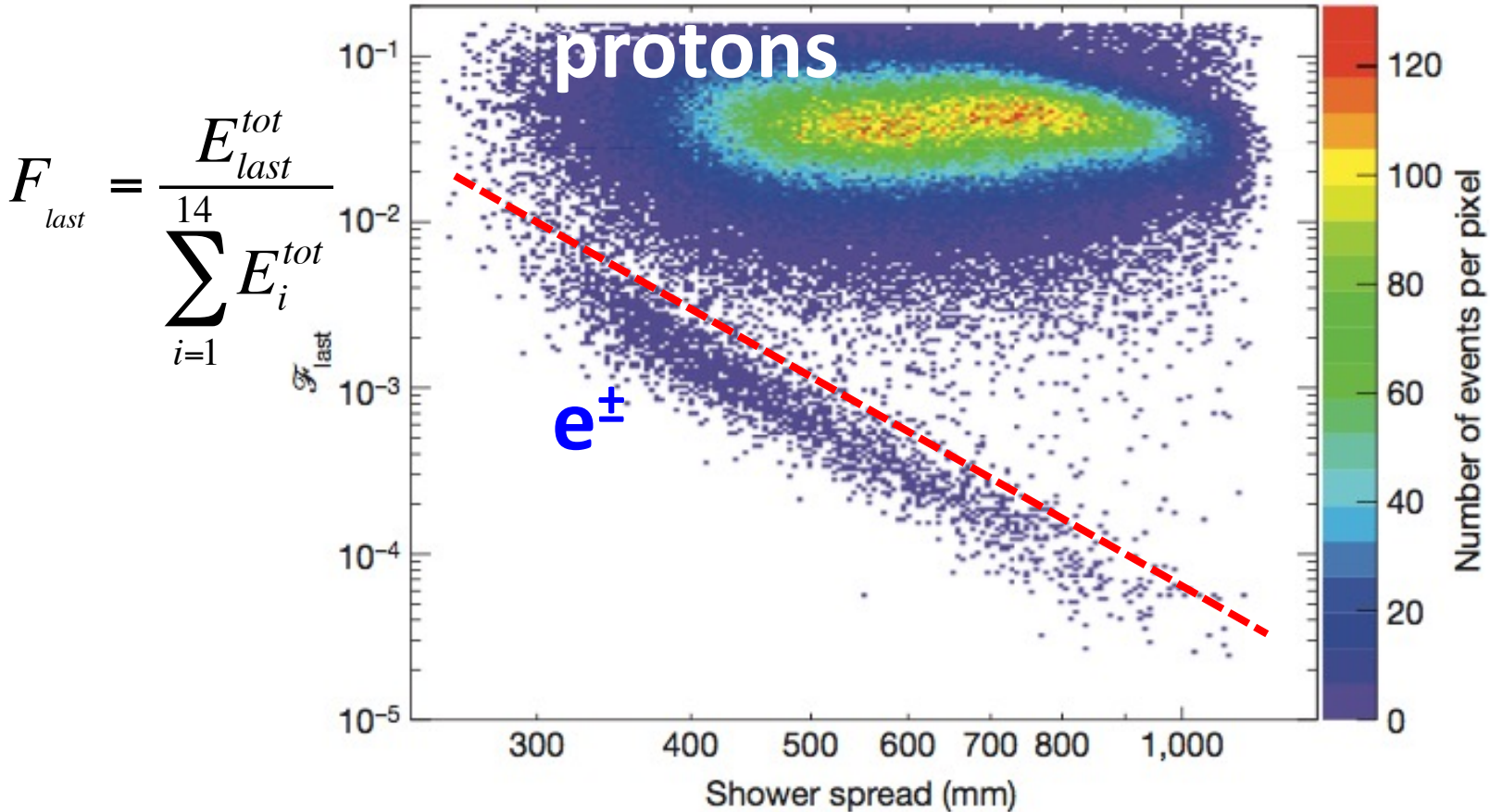
# Electron+positron identification (II)

Selected events with  $Z_{\text{PSD}} = 1$

Exploiting the imaging CALO-features

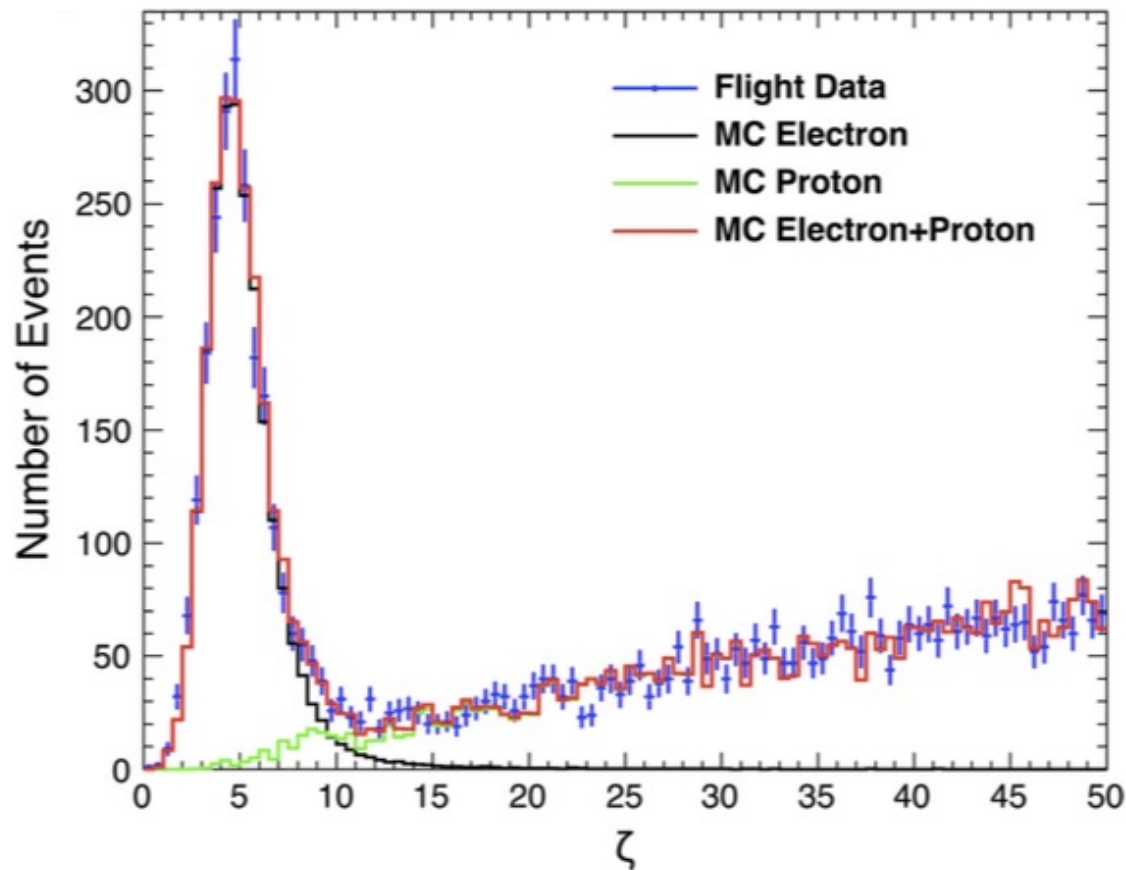


# BGO imaging to separate electrons and hadrons (I)



$$spread = \sum_{i=1}^{14} RMS_i = \sum_{i=1}^{14} \sqrt{\frac{\sum (x_{ij} - x_{iC})^2 E_{ij}}{E_i^{tot}}}$$

# BGO imaging to separate electrons and hadrons (II)

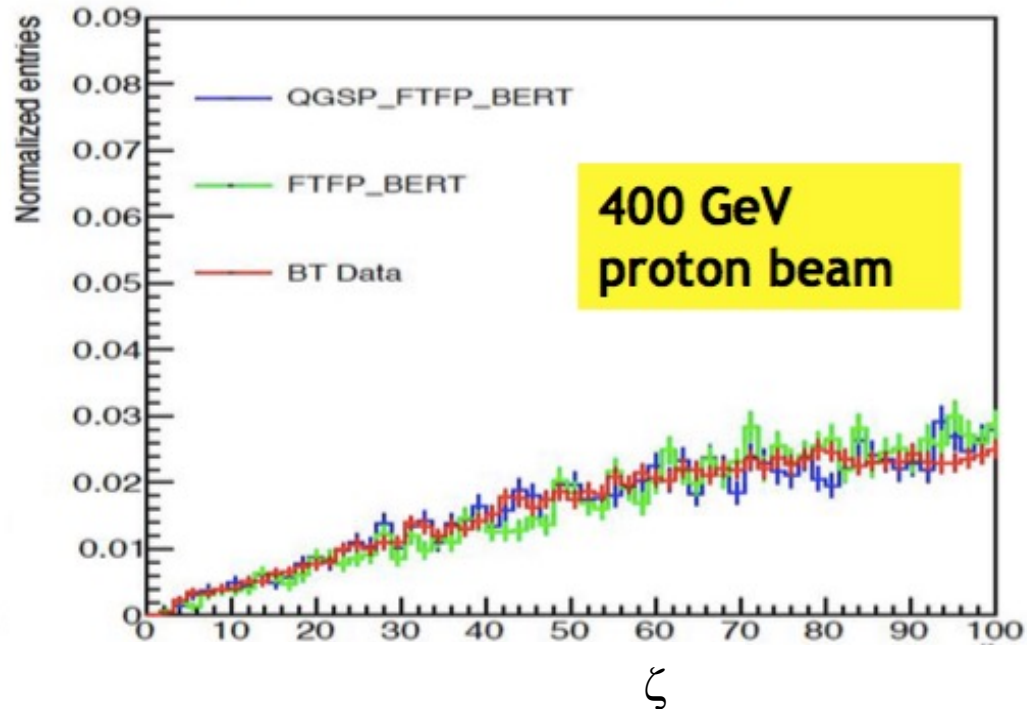
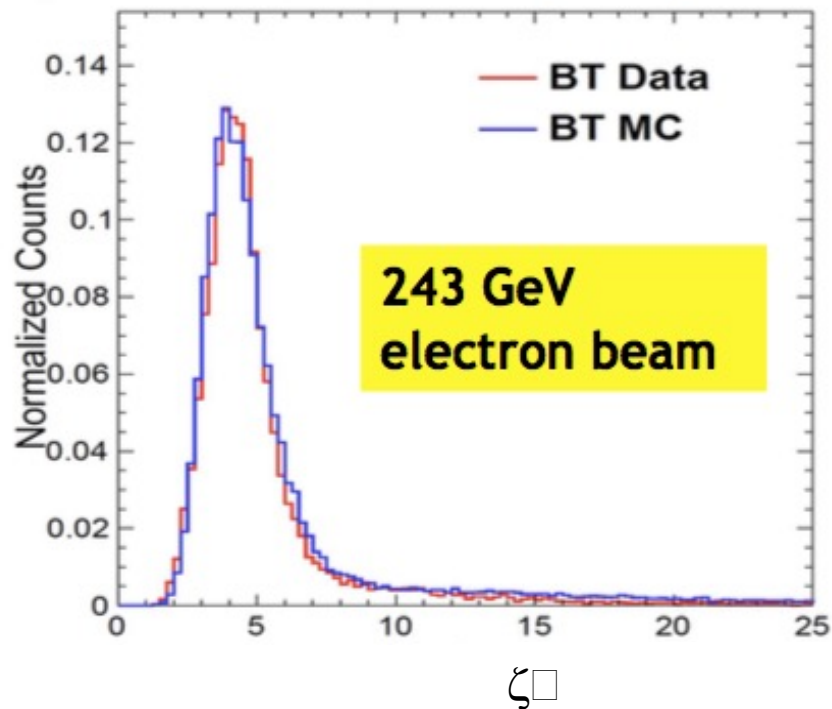


Electron and positron selection

Estimate of proton pollution

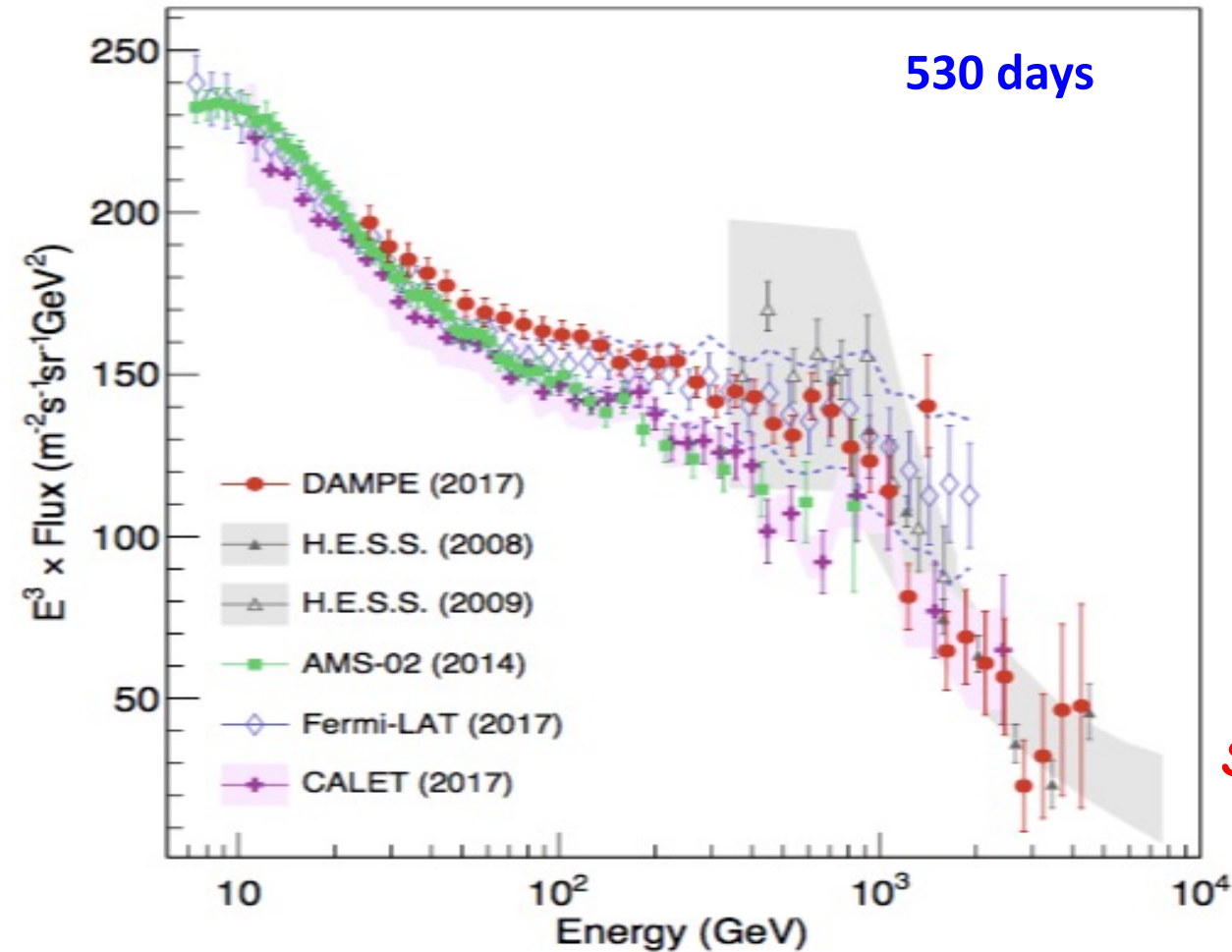
$$\zeta = F_{last} \frac{(\text{spread} / \text{mm})^4}{8 \times 10^6} < 8.5$$

# Validation of parameter $\zeta$ with beam-test data





# Electron+positron spectrum



Energy range:  
25 GeV - 4.6 TeV

Energy resolution <1.2%  
for  $E > 100$  GeV

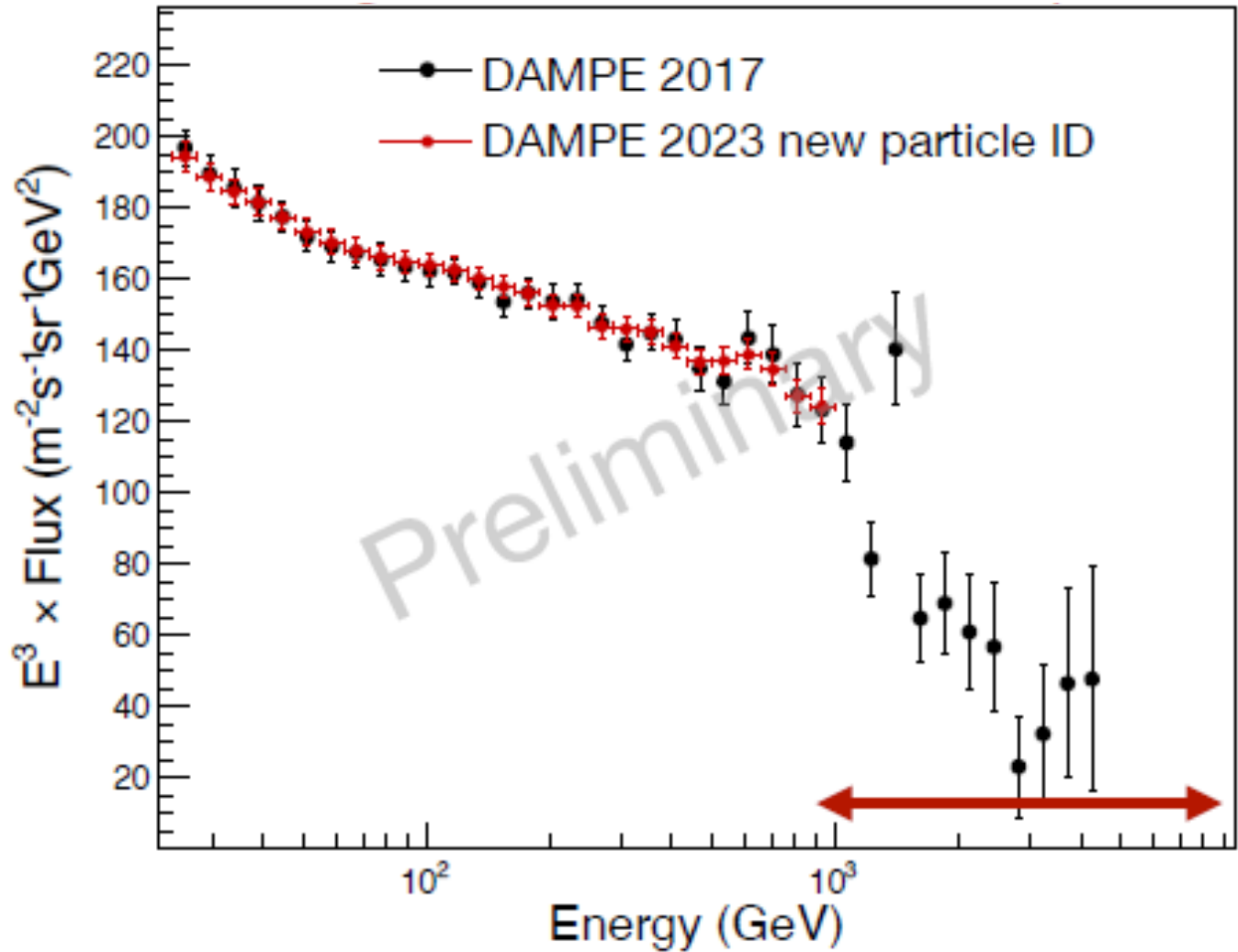
Spectral break  $\sim 1$  TeV

Uncertainties mainly due to the statistics at high energy

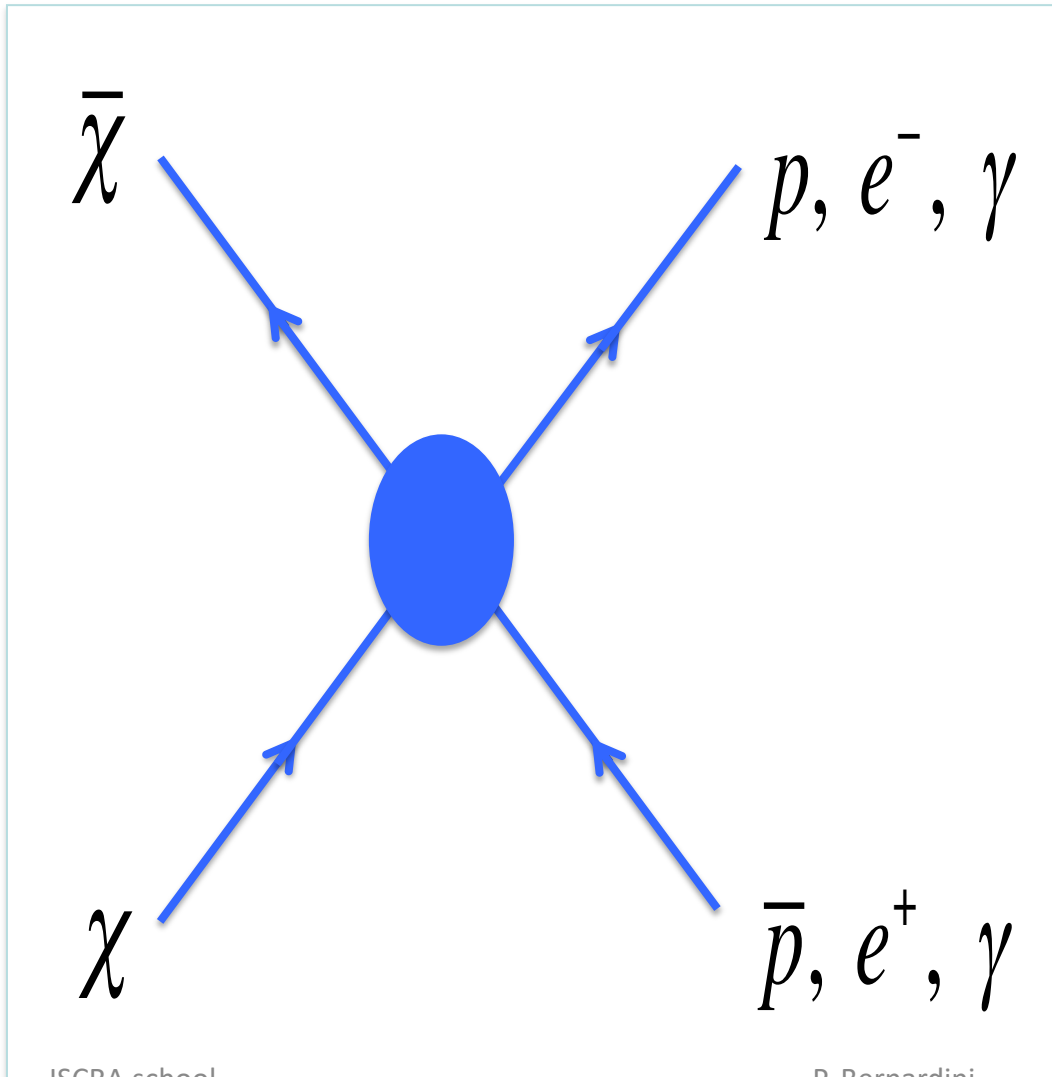


New analysis with Neural Network is ongoing

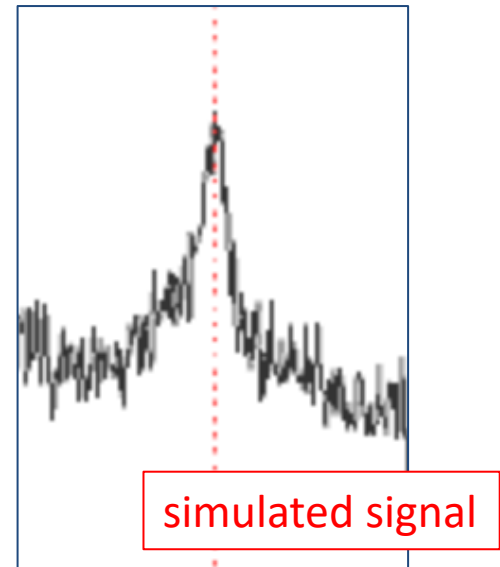
Excellent agreement with standard particle ID



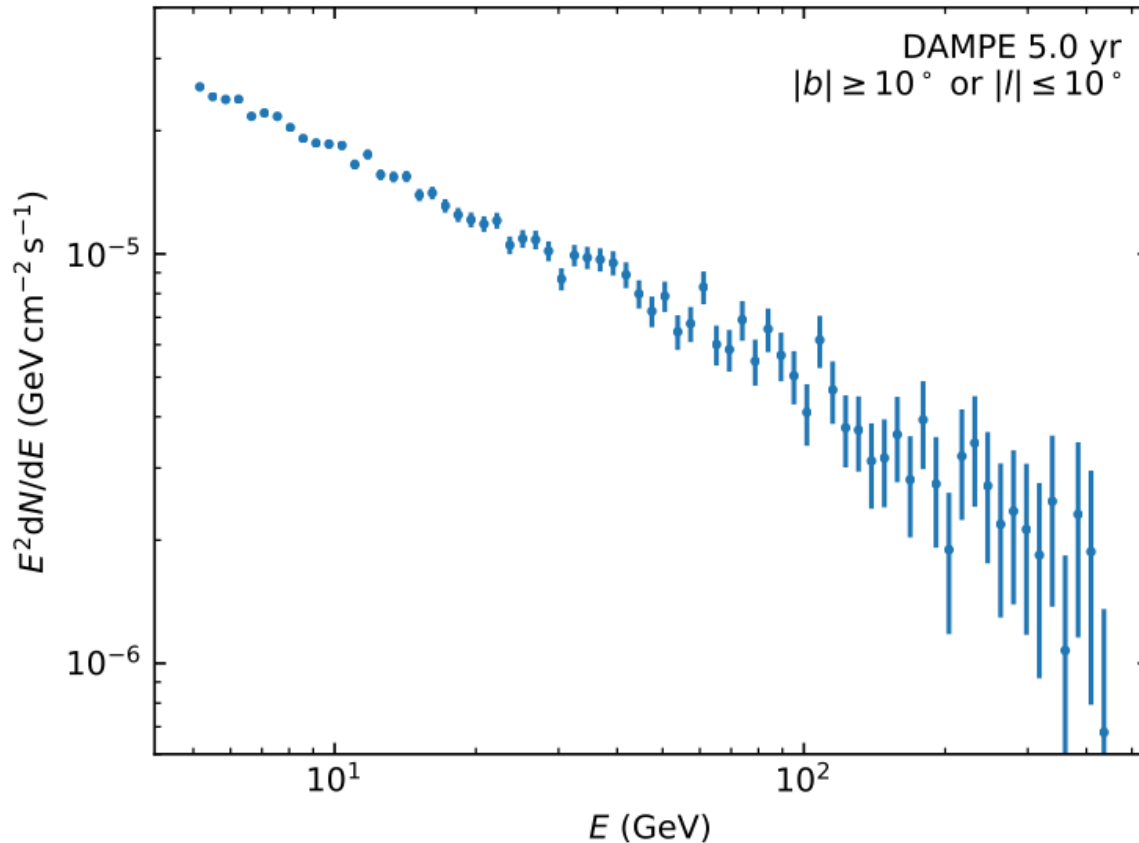
# Possible indirect detection of Dark Matter in space



$\gamma$  line as possible signal ?



# Looking for dark matter lines in photon spectrum



*No visible lines  
up to now*

**Fig. 1.** (Color online) The average SED of the LineSearch and BgoOnly photons from the region with Galactic plane removed.

# Summary

DAMPE is "very deep" ( $\sim 32 X_0$ ) detector in orbit. It works properly since its launch more than 8 years ago. We expect DAMPE will continue to take data for some more years

Long exposure and high energy resolution are the unique tools to find a dark matter signal in the cosmic ray flux (if it exists)

## Main DAMPE scientific results

- Evidence for a break at  $\sim 1$  TeV in the all-electron spectrum
- Clear structures (hardening + softening) have been detected in the proton and Helium spectra
- Z-dependence for hardening and softening is suggested
- Measurement of p+He and all-particle are a bridge toward ground exp.s
- New spectral measurements are in progress (Li, Be, B ... Fe)



# Backup slides

# CALET ratios

